

# Statistics and Data analysis

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**RAL Graduate Lectures**  
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- Why use statistics?
- Have I found something?
- Parameter extraction
- Goodness of fit



# What do we want to achieve?

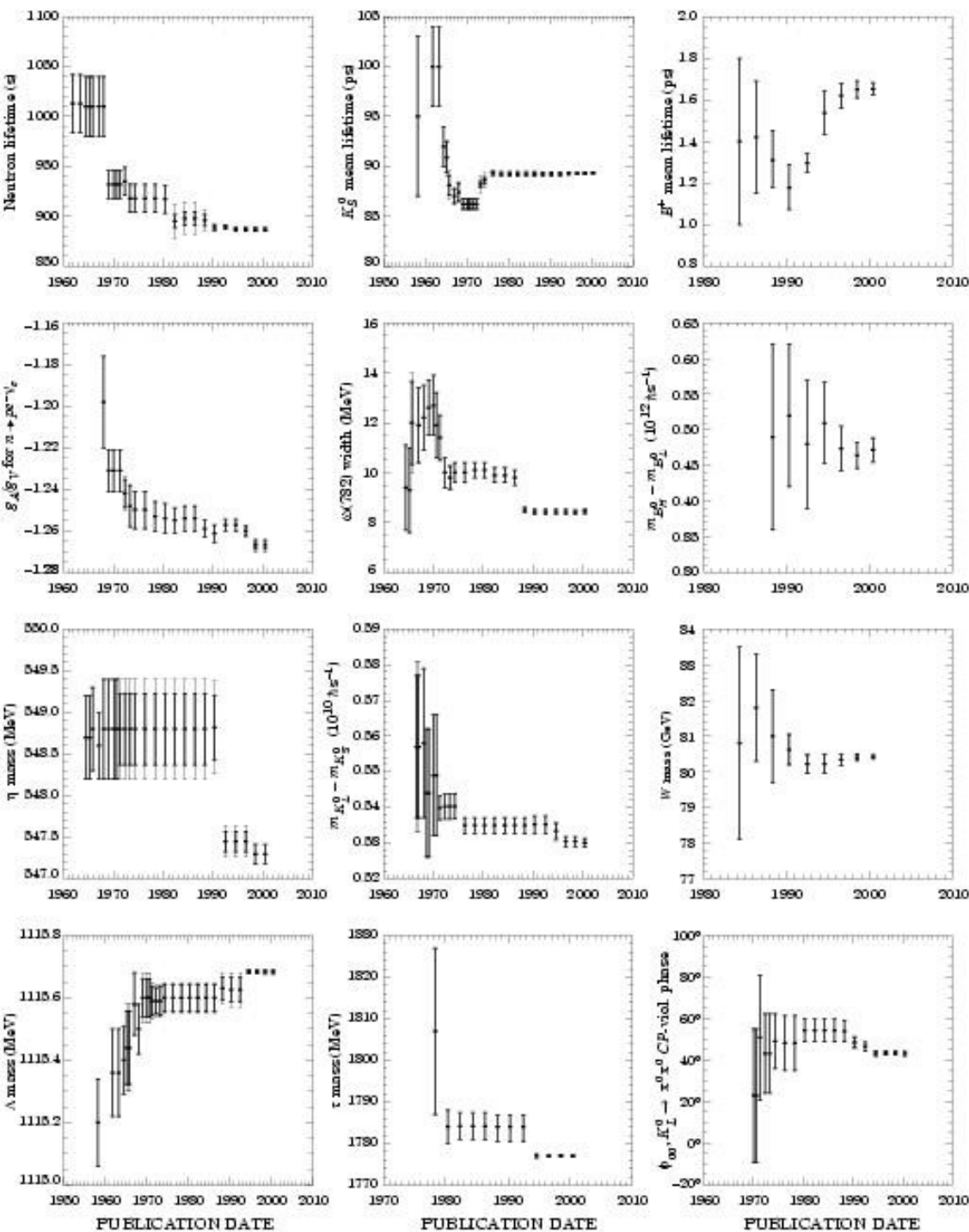
- We aim to make optimal use of the data collected
  - Data are expensive:
    - Use powerful techniques
  - Data processing is also expensive:
    - Mathematical perfection is not the only criterion
  - Systematic errors may well dominate
    - We need to be able to justify our results.



# Three Classes of problem

- Hypothesis testing
  - Does this signal exist?
  - Bayes and Frequentist limits
- Parameter extraction
  - What is the mass of the  $W$ ?
  - Systematic and statistical errors
- Goodness of fit
  - Chi2 test
  - Other tests

# History of measurements



Each measurement agrees with preceding one!

Publications which disagree with the standard model/previous estimates are checked more carefully.

If you look you can usually find something

**Hence Blind Analysis**

Figure 2: An historical perspective of values of a few particle properties tabulated in this Review as a function of date of publication of the Review. A full error bar indicates the quoted error; a thick-lined portion indicates the same but without the "scale factor."



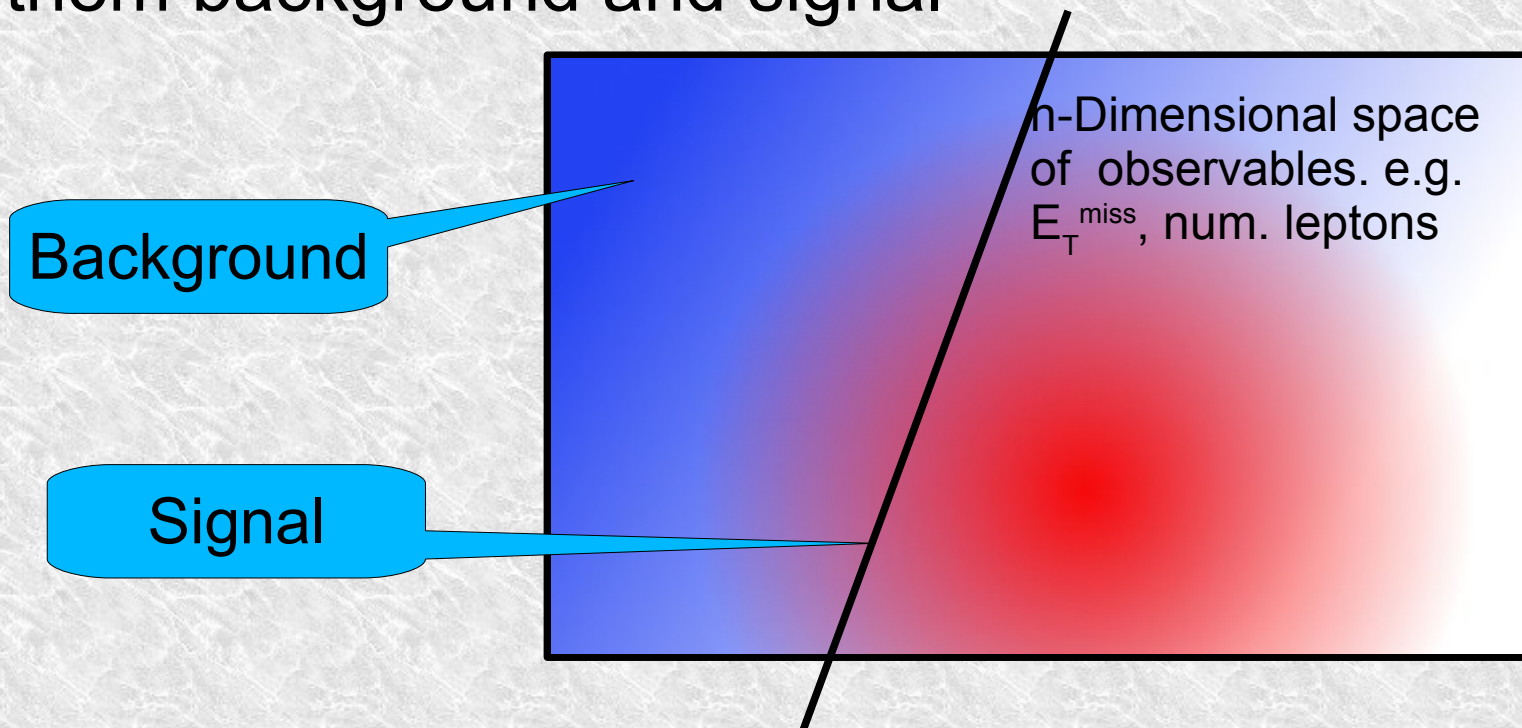
# Hypothesis Testing

- Have we found a new signal?



# Signal Recognition

- Consider separating a dataset into 2 classes
  - Call them background and signal

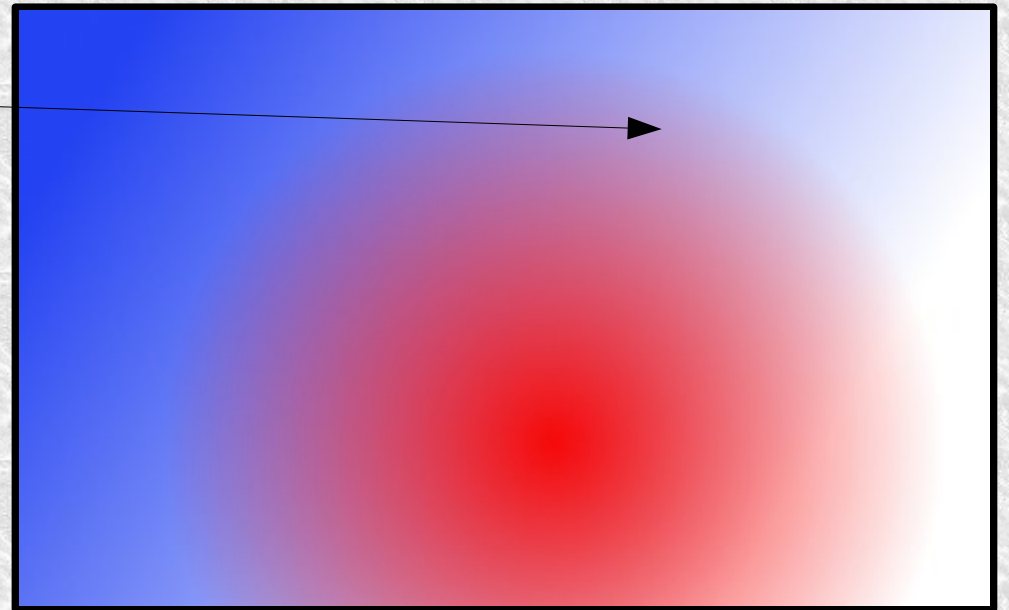


- A simple cut is not optimal



# The right answer II

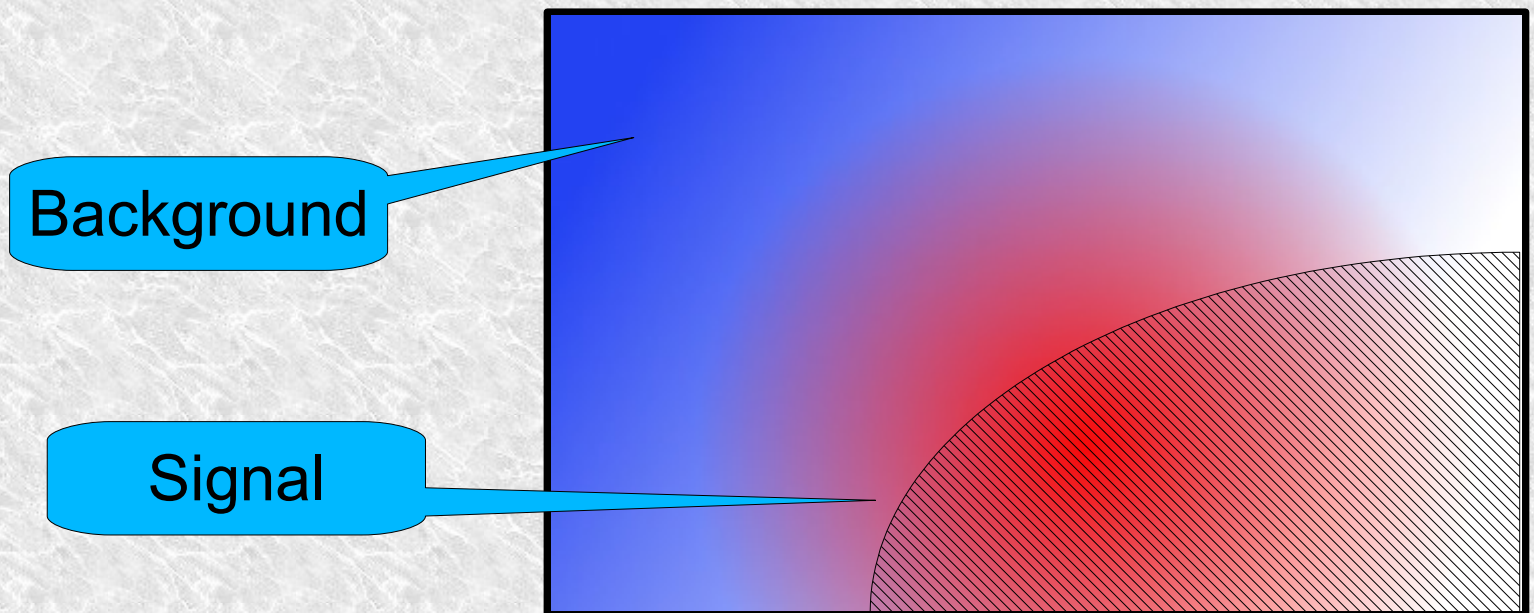
- For a given efficiency, we want to minimize background
- For each event, find ratio of signal and background at that point
- Accept all areas with  $s/b$  above some threshold
- Leading to the ***Likelihood Ratio***





# The right answer III

- Plot region where events are accepted:



- This is the best possible *cut*





## Is that the best we can do?

- Ordering events by  $\rho_s/\rho_b$  and selecting above some threshold gives best possible sample
- But when we ask the question “is there a signal there?” we can extract more information.
- Do a fit to the likelihood-ratios:
  - Use:  $\mathcal{L}=1+\rho_s/\rho_b = (\rho_s+\rho_b)/\rho_b$
  - How much more probable if signal also present?
- Take product over all events
  - One event with  $\rho_b=0$  disproves  $b$
  - 2 events with  $\mathcal{L}=10$  same as 1 with  $\mathcal{L}=100$
- This product is the most sensitive estimator
  - Neymann-Pearson Lemma



# Determination of $s$ , $b$ densities

- We may know matrix elements (“Matrix Element method”)
  - Not for e.g. a  $b$ -tag
  - But anyway there are detector effects
- Usually taken from simulation



# Using MC to calculate density

- Brute force:
  - Divide our n-D space into hypercubes with m divisions of each axis
  - $m^n$  elements, need  $100 m^n$  events for 10% estimate.
  - e.g. 1,000,000,000 for 7 dimensions and 10 bins in each
- This assumed a uniform density – actually need far more
  - The purpose was to separate different distributions



# Better likelihood estimation

- Clever binning
  - Starts to lead to tree techniques
- Kernel density estimators
  - Size of kernel grows with dimensions
  - Edges are an issue
- Ignore correlations in variables
  - Very commonly done **'I used likelihood'**
- Assume measured=true, correct later
  - e.g. 'Matrix element' method
  - Bias correction an issue

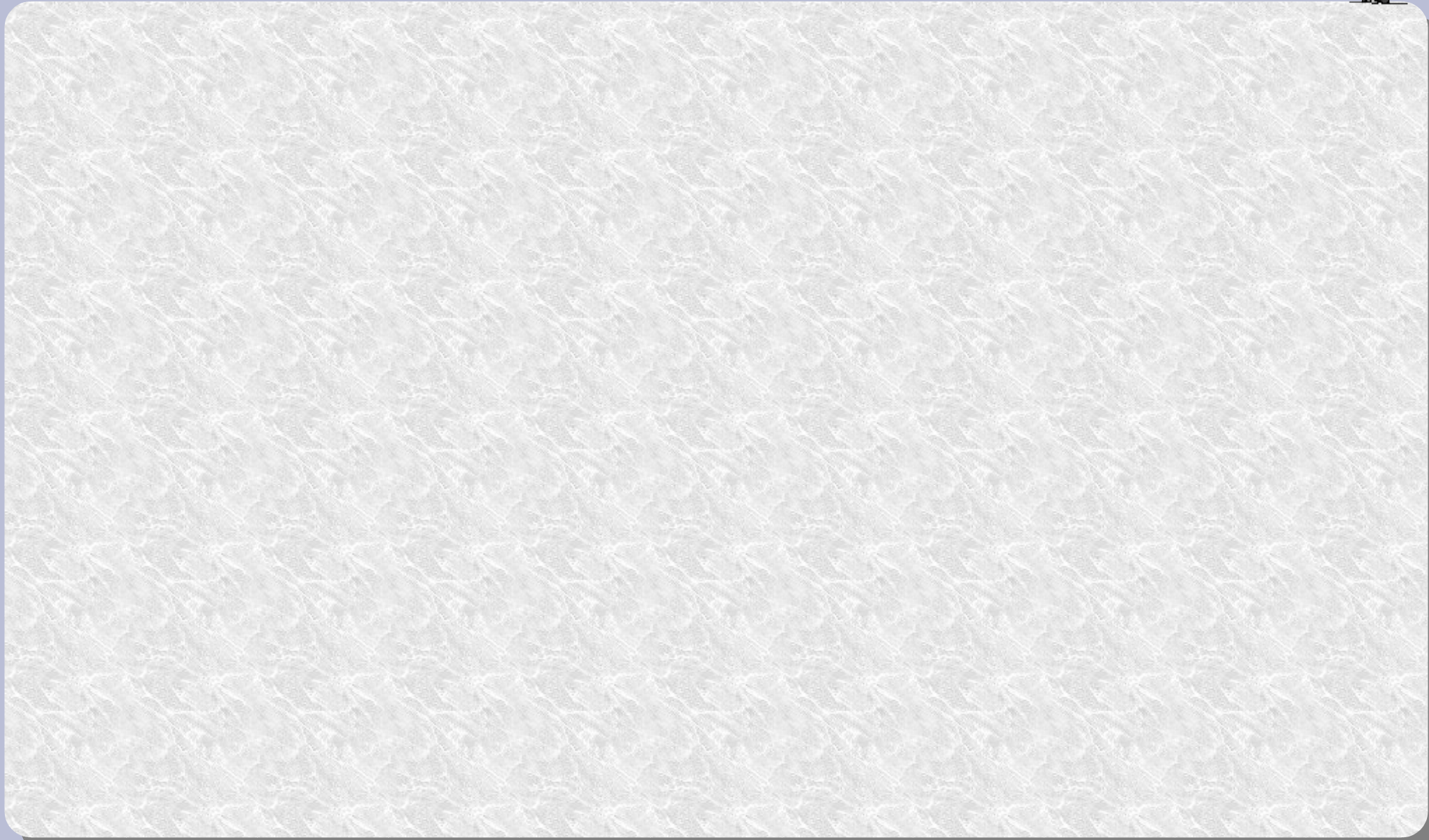


# Alternative approaches

- Neural nets
  - Well known, good for high-dimensions
- Support vector machines
  - Computationally easier than kernel
- Decision trees
  - Boosted or not?



# How to calculate densities



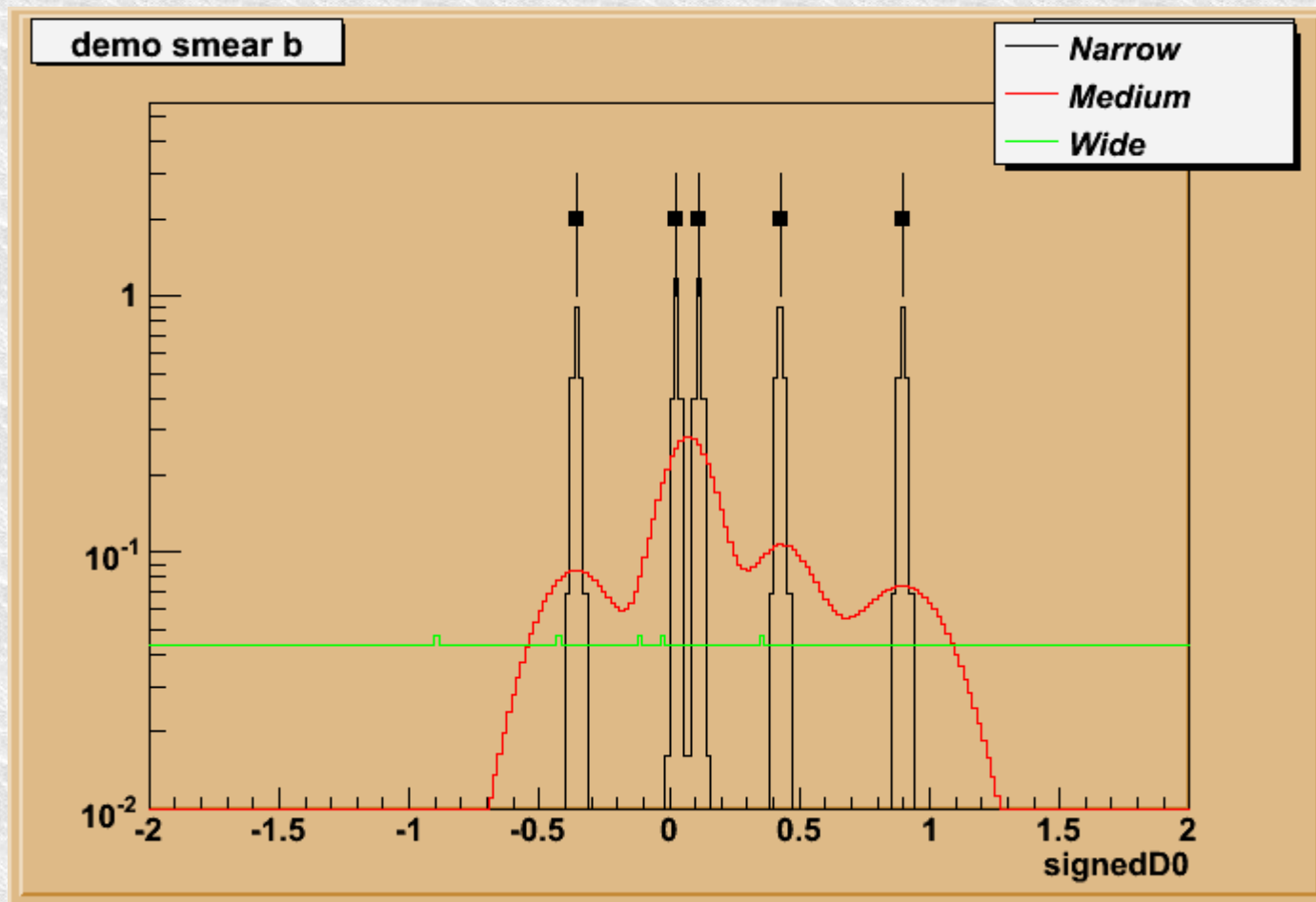


# Kernel Likelihoods

- Directly estimate Probability Density Function of distributions based upon training sample events.
- Some kernel, usually Gaussian, smears the sample
  - increases widths
  - Width of kernel must be optimised
- Fully optimal *if* infinite MC statistics
  - Then we can use narrow kernels



# Smearing 5 events, 1D

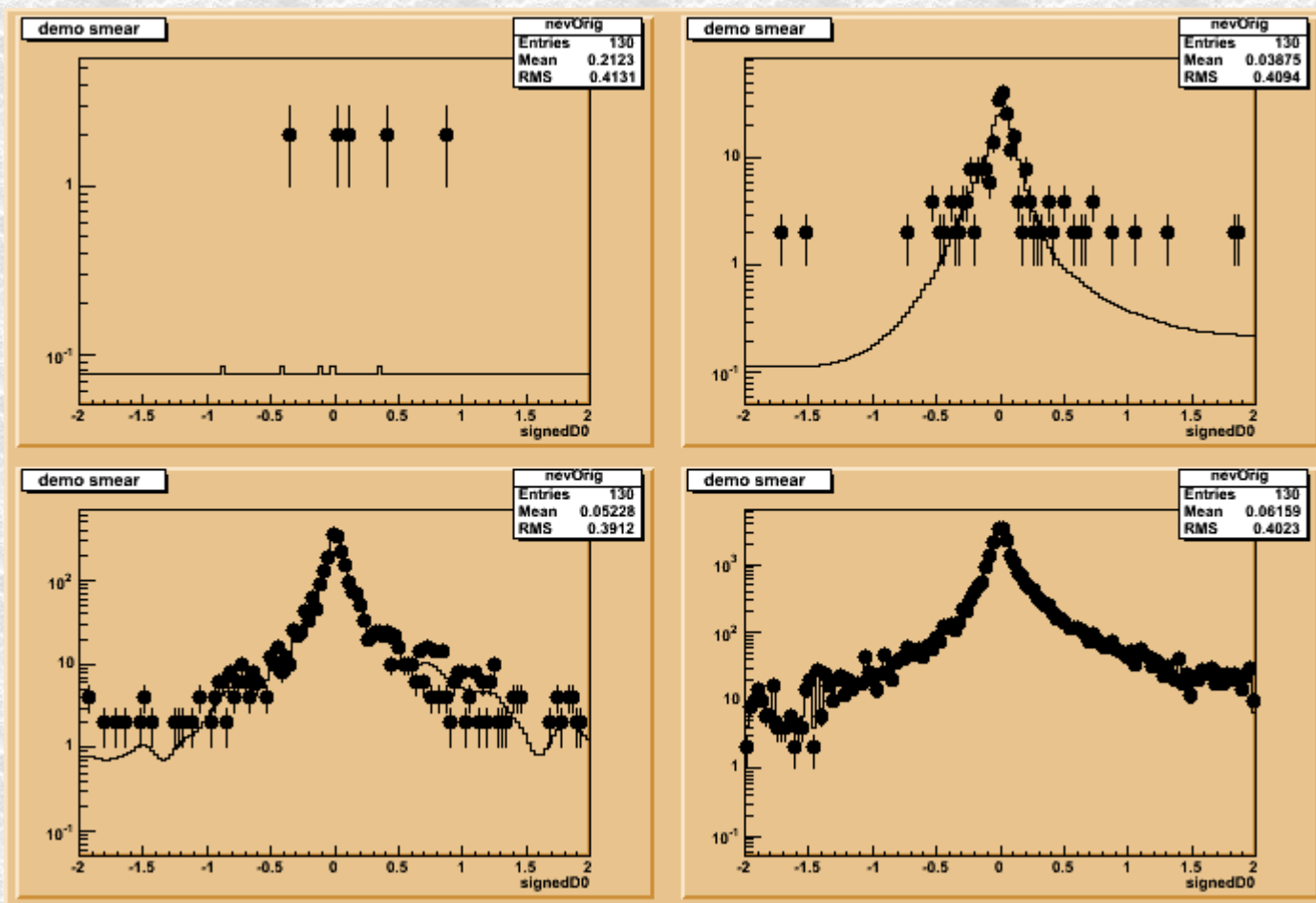


- The kernel width is crucial
- Problem much worse in higher dimensions

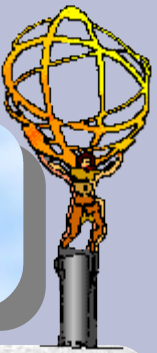




# Varying event nos.



- More events always helps



# Kernel Likelihoods: nDim

- Metric of kernel, (**size, aspect ratio**) hard to optimize
  - Watch kernel size dependence on stats.
- Kernel size must grow with dimensions;
  - Lose precision if unnecessary dimensions added
  - Need to choose which variables to use
- Big storage/computational requirements



# Taming $m^n$

Use of an approximately sufficient statistic or likelihood estimate

- No large resolution and acceptance effects:

Perform fit with uncorrected data and undistorted likelihood function.

- Acceptance losses but small distortions:

Compute global acceptance by MC and include in the likelihood function.

- Strong resolution effects:

Perform crude unfolding.

All approximations are corrected by the Monte Carlo simulation. The loss in precision introduced by the approximations is usually completely negligible.



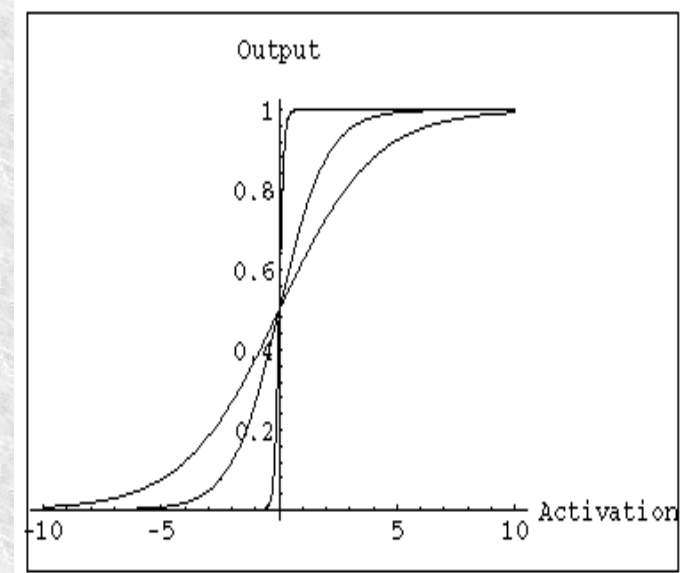
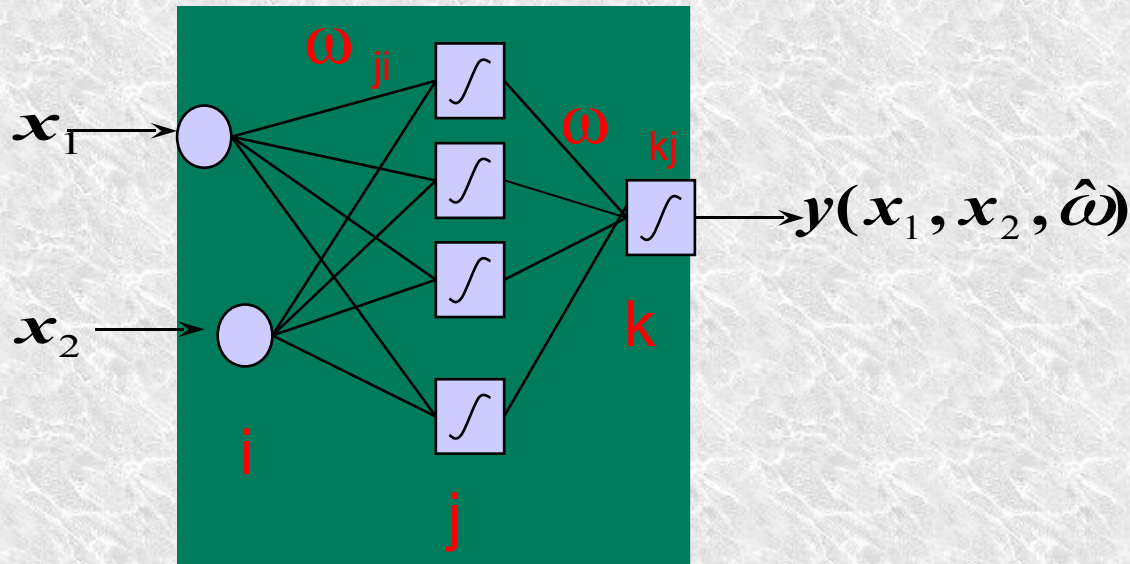
# Neural Nets

- Well known in HEP



# Multi-Layer Perceptron NN

- A popular and powerful neural network:  
(in root)



$$F = \sum_j \omega_{kj} f(\sum_i \omega_{ji} x_i + \theta_j) + \theta_k;$$

$$y = \frac{1}{1 + e^{-F}}$$

Need to find  $\omega$  's and  $\theta$  's, the free parameters of the model



# Neural Network Features

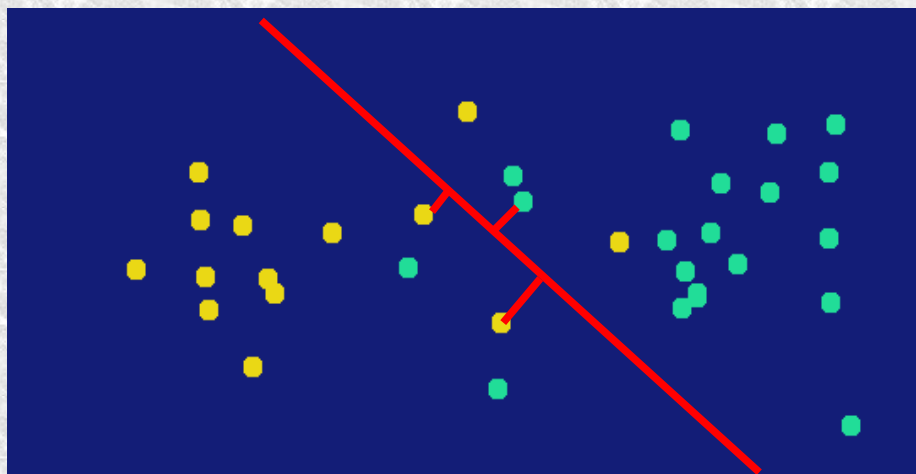
- Easy to use – packages exist everywhere
- Performance is good
  - Especially at handling higher dimensionality
  - No need to define a metric
  - But not optimal (we aim to approx. likelihood)
  - And only trained at one point
- Training is an issue
  - Optimization of nodes/layers can be difficult
  - Can over-focus on fluctuations
  - This is a problem for all machine learning
- Often worth a try
  - But it is solving the wrong problem



# Support Vector Machines

Vapnik 1996

- Simplify storage/computation of separation by storing the 'support vector'

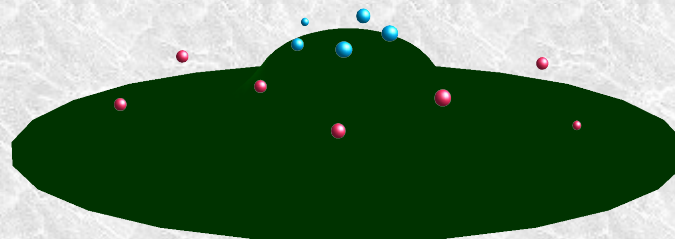


- The straight line is defined by closest points – the support vectors



# Straight lines???

- Straight lines are not adequate, in general. Trick is to project from observed space into higher (infinite?) dimensionality space, such that a simple hyperplane defines the surfaces



- Projection done implicitly by kernel choice
- Only inner-products are ever evaluated, and these are metric independent, so can be calculated in normal space.
- Never need to explicitly define the higher dimensional space





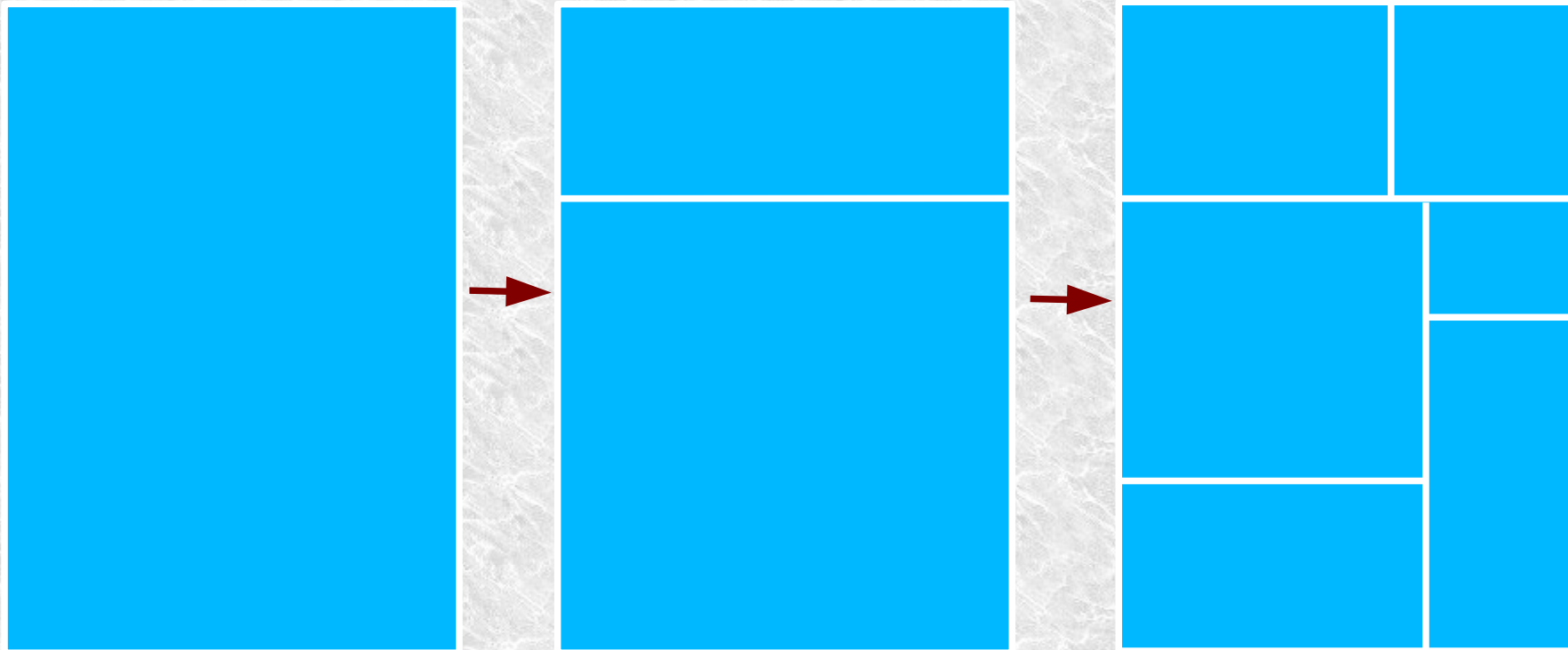
# Support Vector Machines

- Storing just those points lying closest to the line is much easier than storing the entire space
- Defect is that the cut is only well defined near the line
- Computationally much easier than kernel likelihood



# Decision Trees

- A standard decision tree divides a problem in a series of steps.

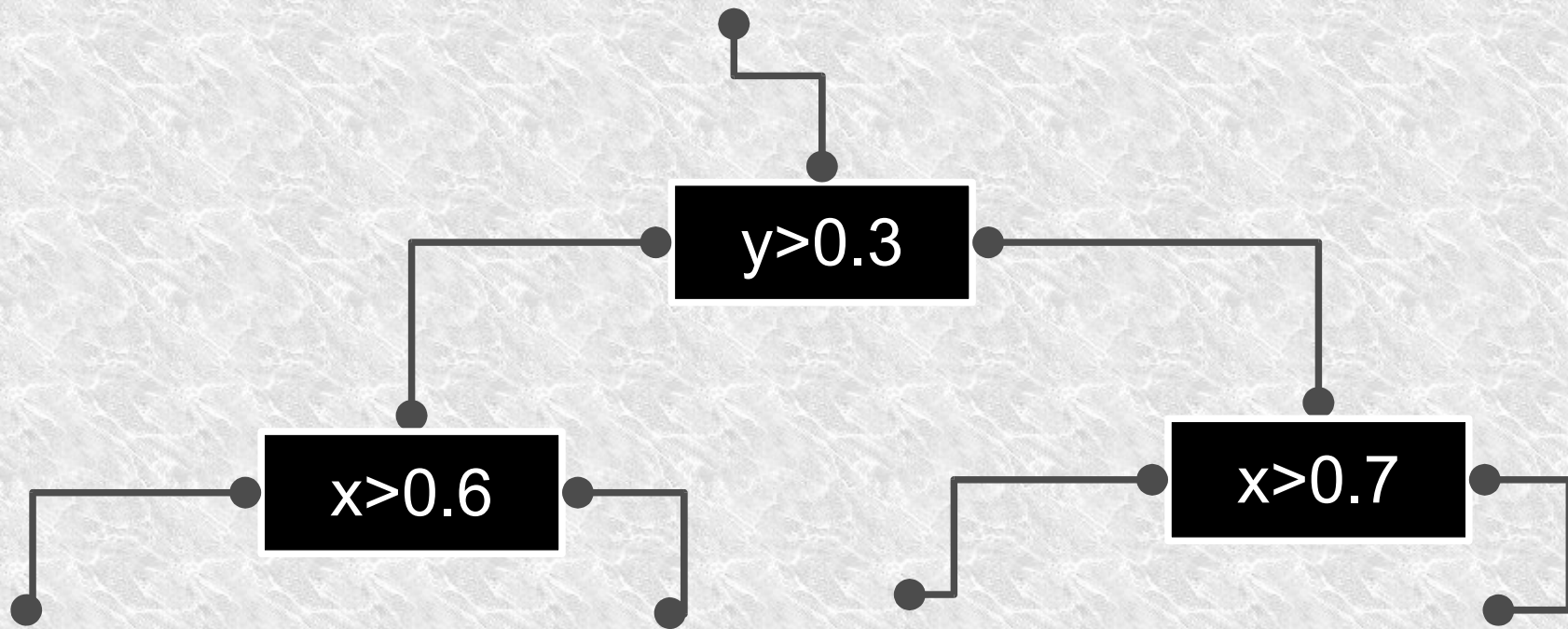


Signal/background evaluated in each box



# Decision Trees II

- Very clear: Each decision is binary, and whole tree can be represented as a tree:



This display works for any dimensionality of problem  
**But how much do we value clarity?**



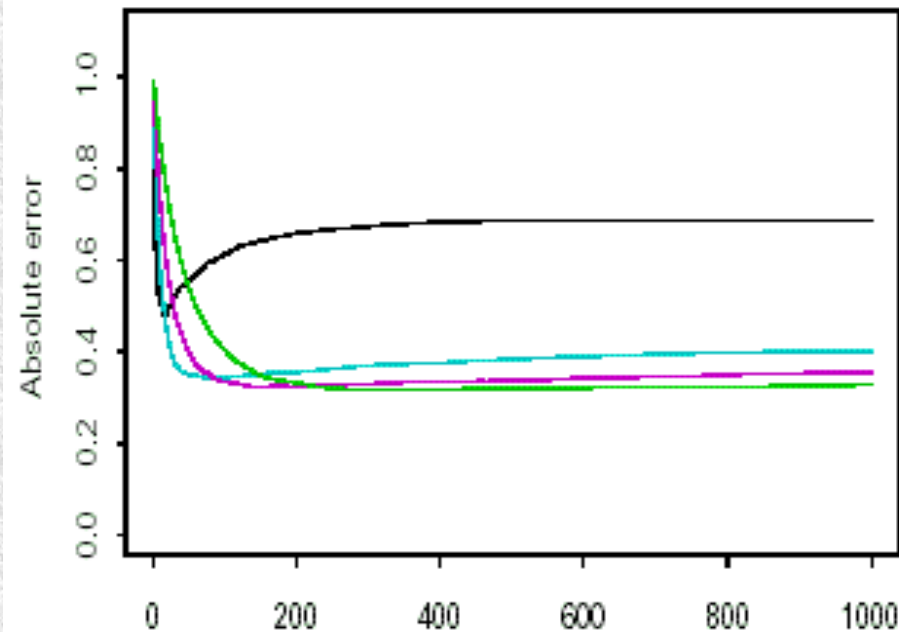
# Decision Trees III

- Downside is lack of stability
  - First 'cut' affects all later ones, classification can vary widely with a different training set
- Power somewhat below NN/Kernel likelihood for typical HEP problems
- Stopping rule important, affected by sample size.



# Boosted Decision Trees

- A first tree is made
- Add more, increasing the weight of the misclassified events
- Final s/b is average (in some sense) over all trees.



Black, cyan, purple and green reflect increasing down-weighting of new trees

Number of trees



# *Boosted Decision Trees*

- Lack of stability removed by averaging
- Computer intensive – but not  $N^3$
- Power very good
- Trees individually small, whole data set is in each  
– good use of statistics
- Fairly fast.

*Breiman: Boosted trees best off the shelf classifier in the world*

**I do not have direct experience here**



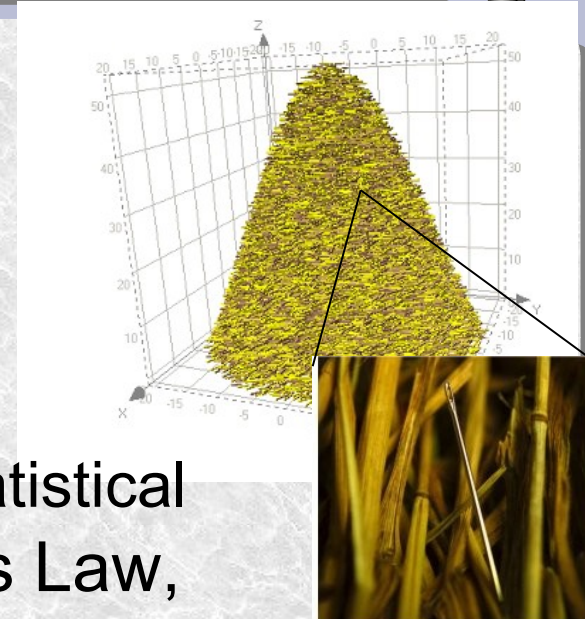
# How to use these?

- Much is available from root
  - TMVA package widely used
  - I have produced 'TNSmooth'
    - Root add-on for kernel likelihood method



# Summary of classification:

- Looking for
  - Needles in haystacks – the Higgs particle
- Needles are easier than haystacks
- ‘Optimal’ statistics have poor scaling
  - likelihood techniques:  $N^3$
  - For large data sets main errors are not statistical
- As data and computers grow with Moore’s Law, we can only keep up with  $N \log N$
- A way out?
  - Discard notion of optimal (data is fuzzy, answers are approximate)
  - Don’t assume infinite computational resources or memory
- Requires combination of statistics & computer science





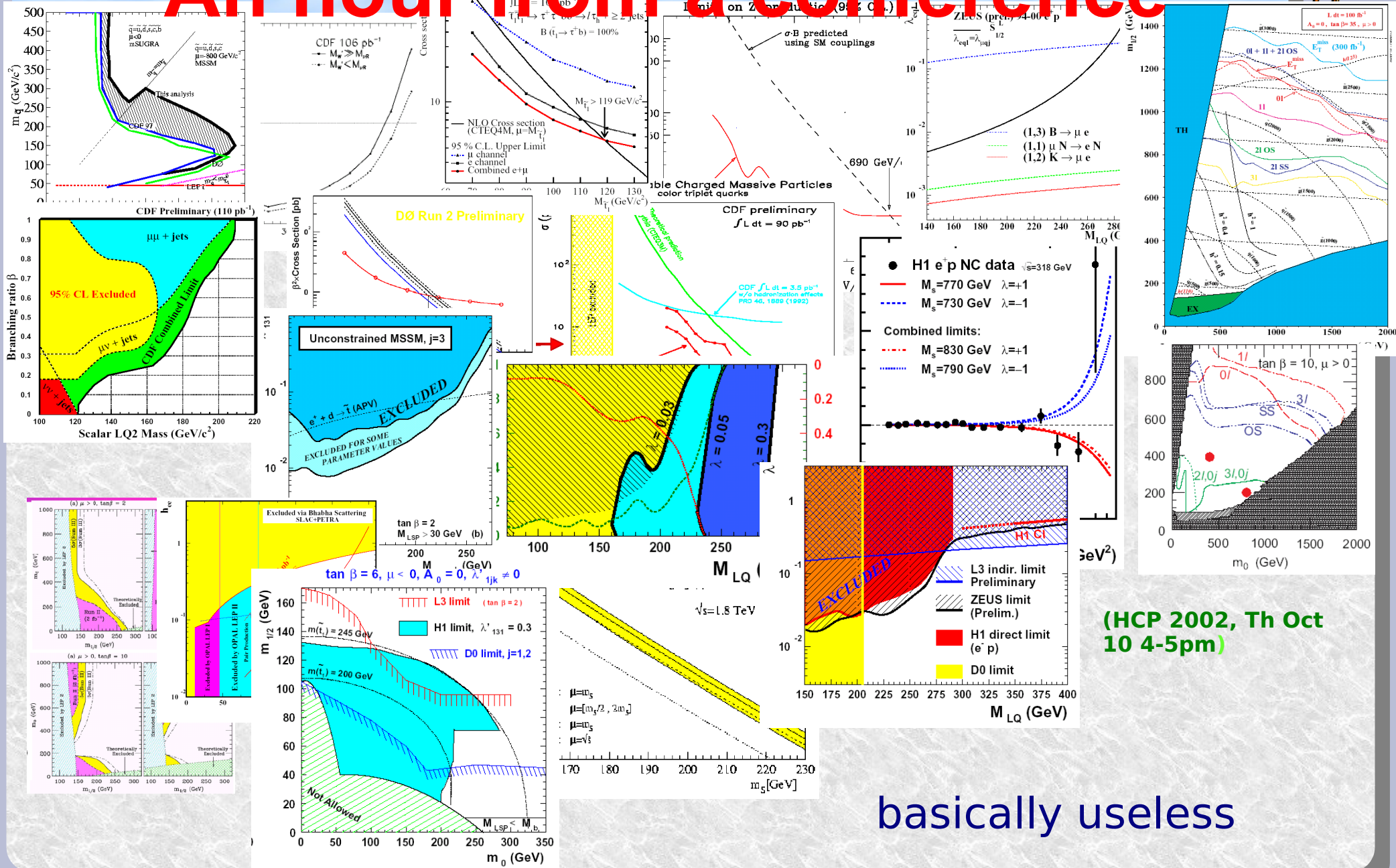


# Limits

- Definition of a limit
  - Bayes vs Frequentist
- Errors on limits
- Combining limits
- You need to know: What is being limited?



# An hour from a conference



(HCP 2002, Th Oct 10 4-5pm)

basically useless



# Bayes and Frequentist Statistics

- Bayes confidence level
  - Probability of theory given data
  - `There is a 5% probability  $X < 0$ `
  - Right question, but **subjective**
- Frequentist/Classical P-value
  - Probability of data given theory
  - `If  $X < 0$ , probability of these (or more extreme) data is 5% or less`
  - Wrong question



# Bayes Theorem

- Bayes theorem modifies probabilities in light of data
- But needs an a-priori probability

$$p(a \wedge b) = p(a) p(b|_a) = p(b) p(a|_b)$$

$$p(b|_a) = \frac{p(b)}{p(a)} p(a|_b)$$

Uncontroversial

Requires  $p(\text{theory})$

$$p(\text{theory}|_{\text{data}}) = \frac{p(\text{theory})}{p(\text{data})} p(\text{data}|_{\text{theory}})$$

- If data plentiful frequentist and Bayesian converge
- Bayesian statistics much easier to use



# Frequentist limits

'For this theory, probability of these data is  $<5\%$ '

- About as useful as:

'It was raining when I went to the theatre'

- e.g. a Higgs Search, background=3, observed=0
  - S=3: P = 0.2%
  - S=1: P=1.8%
  - S=0: P=5%
- So even a production of 0 Higgses is excluded at 95% CL.  
( $M_H=1000$  excluded!)
- Mathematically fine, but not useful



# Bayes vs Frequentist Statistics

## II



- For measurement, almost all results are implicitly Bayesian - but could be justified frequentistically.
- Limits are much less clear - **read the small print**
  - Many modifications (e.g. Feldmann & Cousins, RPP 2000) try to give Bayesian properties to classical limits
  - Without such modification, frequentist limits can be meaningless. (e.g. Example on previous page)

**Fundamental problem:**

$$P(\text{data}|\text{hypothesis}) \neq P(\text{hypothesis}|\text{data})$$

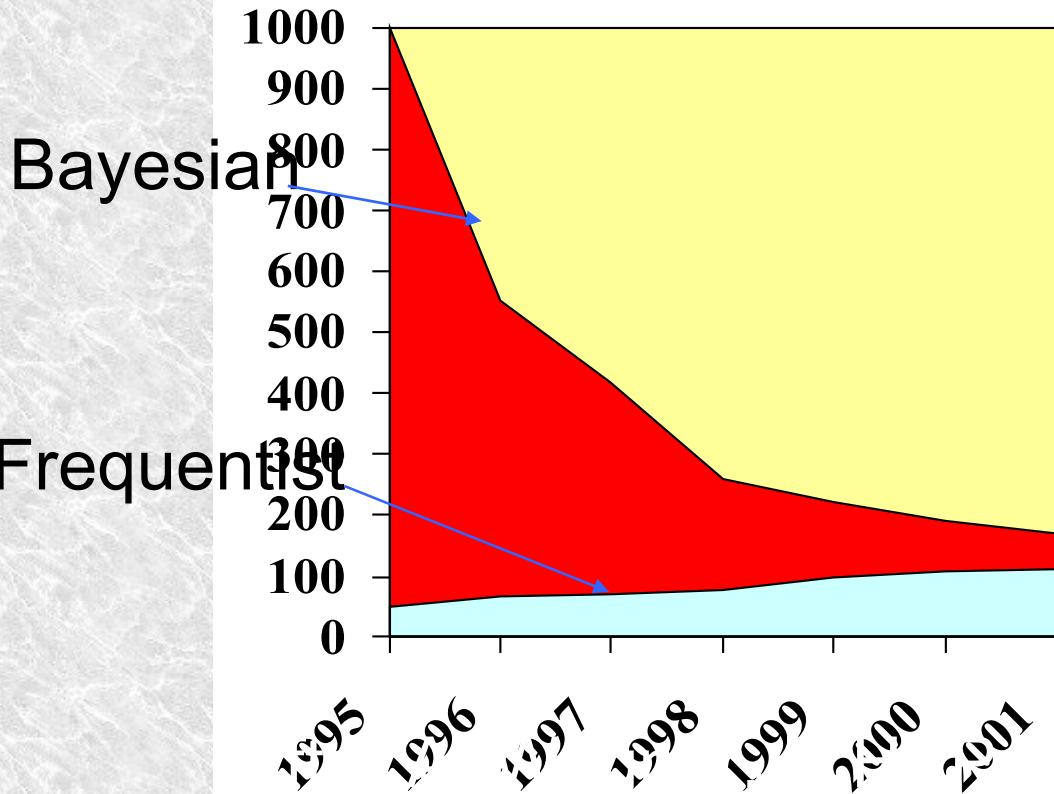


# Error on a limit

- Sometimes people ask: What is the error on this limit?
- *This is the wrong question, it hides two others:*
  - How different would the limit be if
    - the data was a bit different?
    - you quoted 90% or 99% instead of 95%
- It is wrong, because  $M_H > 113 \pm 1$  helps no-one.  $M_H = 115 \pm 1$  is useful



# Mixed up results!



## Higgs limits

EW fits assume a Higgs

Search looks for one

- Excluded by EW fits
- Allowed
- Excluded by Direct Search

To be fair, he had no choice - this is what is provided





# So what does a limit mean?

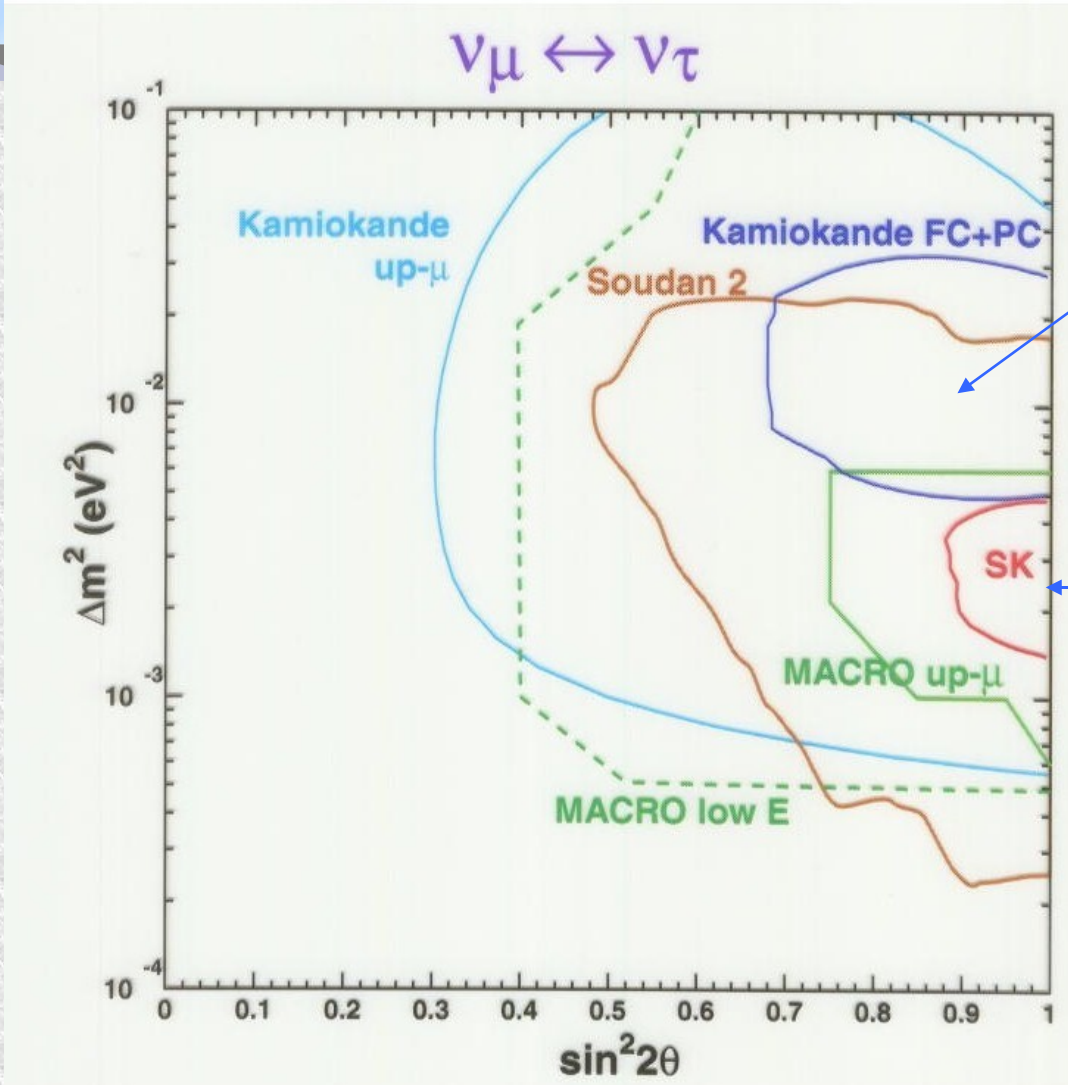
- Usually, a statement like:

$$M_H > 113.2 \text{ GeV}/c^2 @ 95\% \text{ CL}$$

- Does NOT Mean that there is a 95% probability that  $M_H > 113.2$
- DOES mean that **IF**  $M_H < 113.2$  **then** there was at most a 5% probability we missed it.
- But you should not CARE anyway - it is a *probability*



# Combination of limits

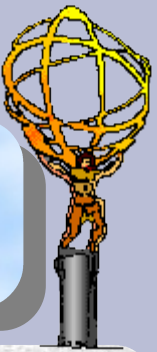


Kamiokande 90%

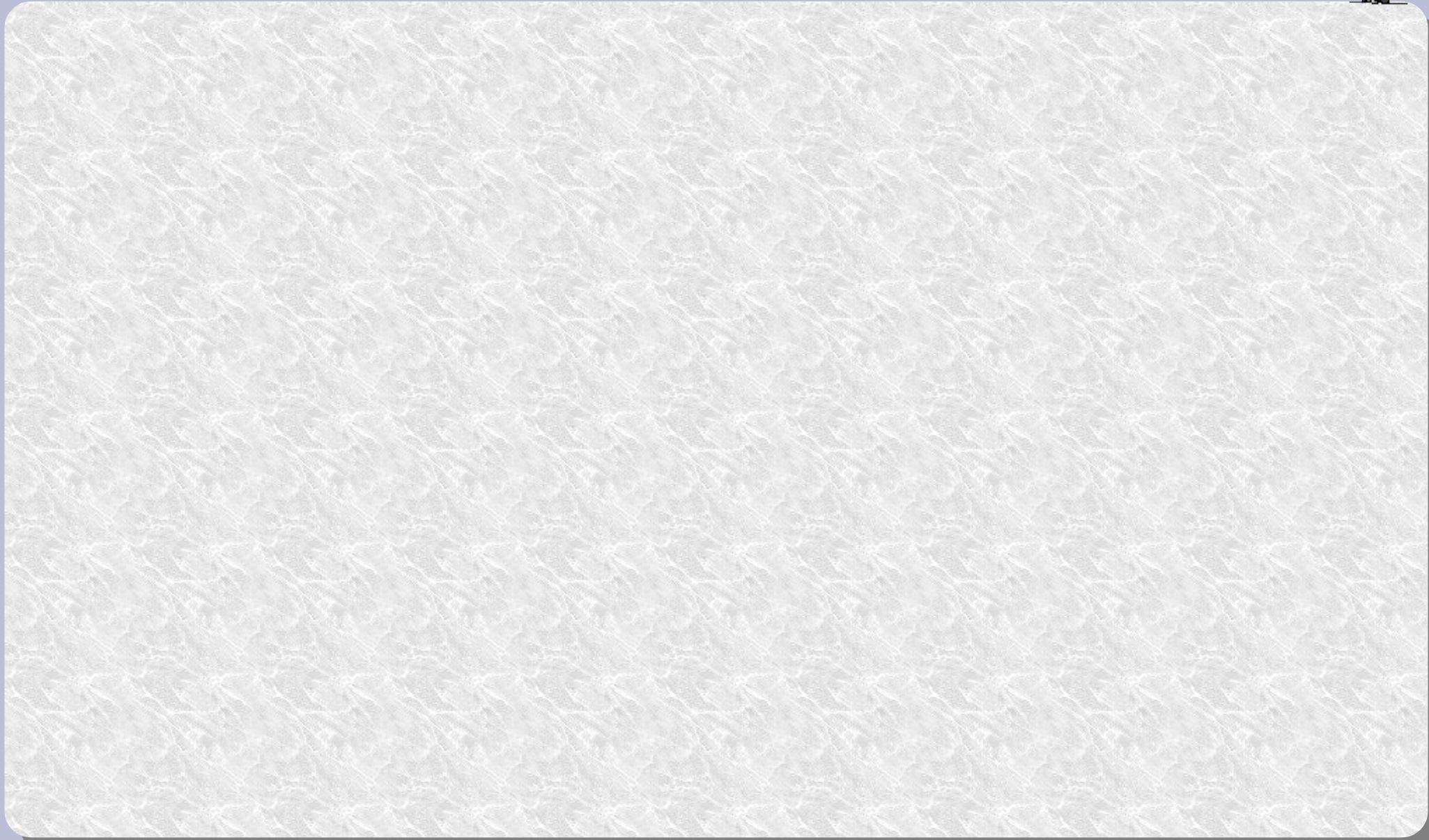
Super-K 90%

Results do not overlap! How can you combine?

Please, don't do it!



# Parameter Extraction





# Statistics and Systematics

- Data are collected very often in terms of **numbers of events**
- These are converted into cross-sections; rates at which things happen.
- This is true of almost all HEP measurements
- The statistical precision is fairly straightforward to estimate, and will have a Poisson or Binomial form which can often be approximated with a Gaussian (CLT)
- Various systematic effects will also affect the result.



# Evaluation of Systematics

- **Usually the hardest part of a measurement**
- **Frequently done badly**
- **Can be under- or over- estimated**
  - **Sometimes an error source is forgotten**
  - **Frequently statistics get taken twice**



# Systematics

- Typical experimental statements:
  - $M_W = 80.336 \pm 0.055(\text{stat}) \pm 0.028(\text{sys}) \pm 0.025(\text{FSI}) \pm 0.009(\text{LEP})$   
(DELPHI)
  - $M_W = 80.329 \pm 0.029$  (EW group)
- First has statistical and systematic errors
- The second does not. **Why?**



# Systematics in combination

- Take an example: Two measurements of 'x'

- $x_a = 50 \pm 10(\text{stat}) \pm 1(\text{sys})$  (expt. a)

- $x_b = 60 \pm 1(\text{stat}) \pm 10(\text{sys})$  (expt. b)

- To combine, use:

Correlations?

$$\frac{x_{tot}}{\sigma_{tot}^2} = \sum_i \frac{x_i}{\sigma_i^2}$$

$$\frac{1}{\sigma_{tot}^2} = \sum_i \frac{1}{\sigma_i^2}$$

- $x_{tot} = 55 \pm 7.11$  (total)

- But either stat. or syst. combined separately is <1!

Errors are completely mixed in combination  
be evaluated on equal footing

- must



# Why is fitting important?

- Experimental data usually used to obtain theoretical parameter values
- There are two distinct questions:
  - What is the best value of  $X$  (Parameter optimization)
  - Does the theory explain the data (Goodness of fit)
- The use of is very different



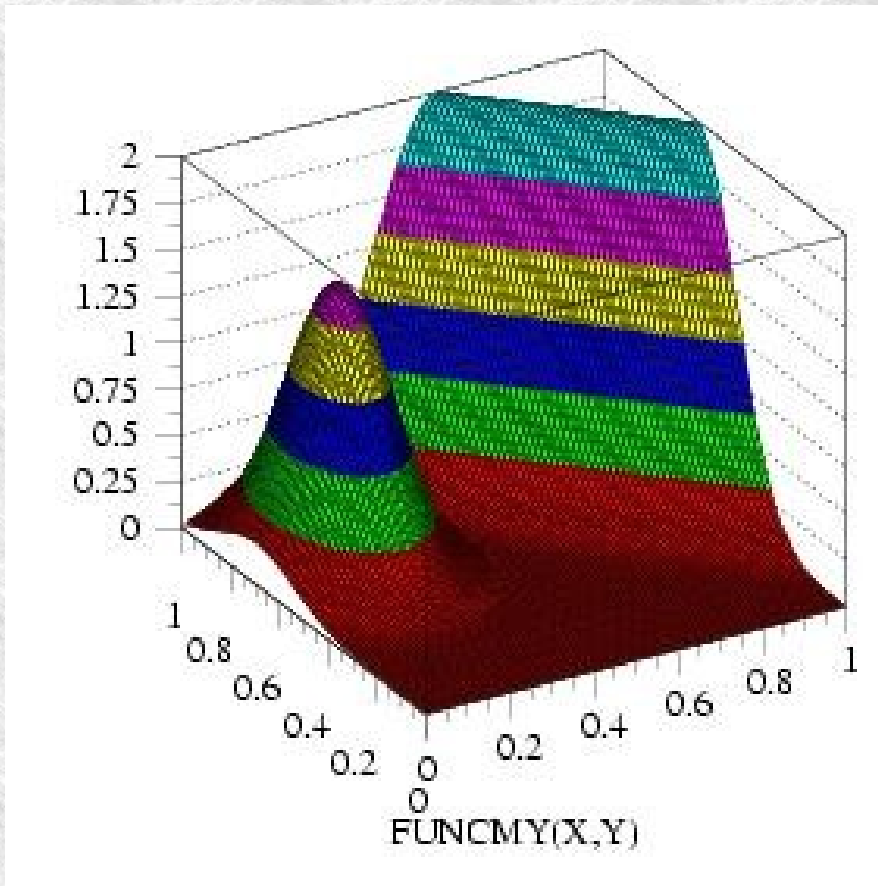


# Parameter optimisation

- What value of the free parameters best describes the data?
- **'MINUIT'** from root/CERNLIB a very common minimization package
- User must specify WHAT is minimized ( $\chi^2$ , likelihood, etc.)
- **MINUIT** will (usually) find the optimal parameter set
- Complex functions may have **secondary maxima / convergence failure**



# Example of a hard maximisation



- Subsidiary minimum may confuse algorithm
- Primary maximum has very high parameter correlation
- Likely convergence failure



# The right way to fit?

|                               | $\chi^2$    | <i>likelihood</i> |
|-------------------------------|-------------|-------------------|
| <b><i>Binned</i></b>          | Yes         | Yes               |
| <b><i>Unbinned</i></b>        | No          | Yes               |
| <b><i>Works low stats</i></b> | No          | Yes               |
| <b><i>Unbiased</i></b>        | Sometimes   | Yes               |
| <b><i>Goodness-of-fit</i></b> | $P(\chi^2)$ | Hard              |

The techniques converge for high statistics

Both can be handled by 'Minuit'

No *right way*



# Definition of $\chi^2$

→ Compatibility of results with expectation:

$$\chi^2 = \sum_i \frac{(x_i^{obs} - x_i^{pred})^2}{\sigma_x^2}$$

→ If counting events in bins then:

→ Beware: Is  $\sigma = \sqrt{N^{obs}}$  or  $\sqrt{N^{pred}}$ ? - Both are wrong!

→  $\sqrt{N^{obs}}$  Biased down if data down... (Mean 5 & 3 = 3.75)

→  $\sqrt{N^{pred}}$  Error depends upon theory - biased up (~4.12)



# Definition of likelihood

$$L_r = e^{-r} \times \prod_i R_i$$

$$LR = \frac{L_{s+b}}{L_b} = e^{-s} \times \prod_i \frac{S_i + B_i}{B_i}$$

→r: The total rate

→ $R_i$ : The density at point i

→s: signal

→b: background

→Likelihood ratio compares two hypotheses.

→Or vary R to maximize L. Maximum likelihood powerful estimator.

$$\log(L_R) = -r + \sum_i \ln R_i \quad \leftarrow \text{weighted sum of events.}$$



# Comparison of definitions

$$\chi^2 = \sum_i \frac{(x_i^{obs} - x_i^{pred})^2}{\sigma_x^2}$$

$$\log(L_R) = -r + \sum_i \ln R_i$$

Take a Gaussian centred on 0, width  $\sigma$

$$\chi^2 = \sum_i \frac{x_i^2}{\sigma^2}$$

$$\log(L_R) = -r + \sum_i \ln \left( e^{-0.5 \frac{x_i^2}{\sigma^2}} \right)$$

$$\log(L_R) = -r + \sum_i -0.5 \frac{x_i^2}{\sigma^2}$$

So:  $\delta \chi^2 = -2 * \delta \log(L_R)$  **IF GAUSSIAN**



# Goodness of fit



# Averaging two numbers

- Suppose we have two measurements of  $x$ :  
 $10 \pm 1$  and  $11 \pm 1$
- We know the average:  
 $10.5 \pm 0.7$
- But what about  
 $10 \pm 1$  and  $20 \pm 1$
- Are we happy with:  
 $15.0 \pm 0.7$
- If the errors are Gaussian we should be happy
  - Combining two Gaussians gives a Gaussian
- Or we conclude we have a 'bad fit'





# Goodness of fit

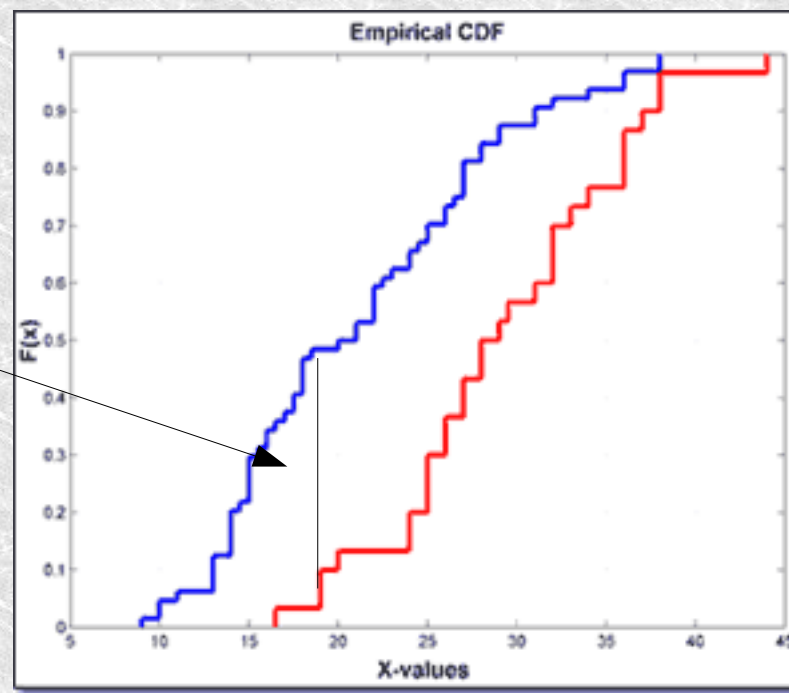
Do I believe my fitted results?

- For a  $\chi^2$  fit (much the most common) easy
  - Look up  $\text{Prob}(\chi^2, \text{NDF})$  in a table.
  - Good general test
    - often abused, e.g.  $\chi^2/\text{DoF}$
- For a likelihood fit: hard
  - Can sometimes use simulated trials to find Probability of getting observed result OR `larger' one



# Kolmogorov-Smirnov Test

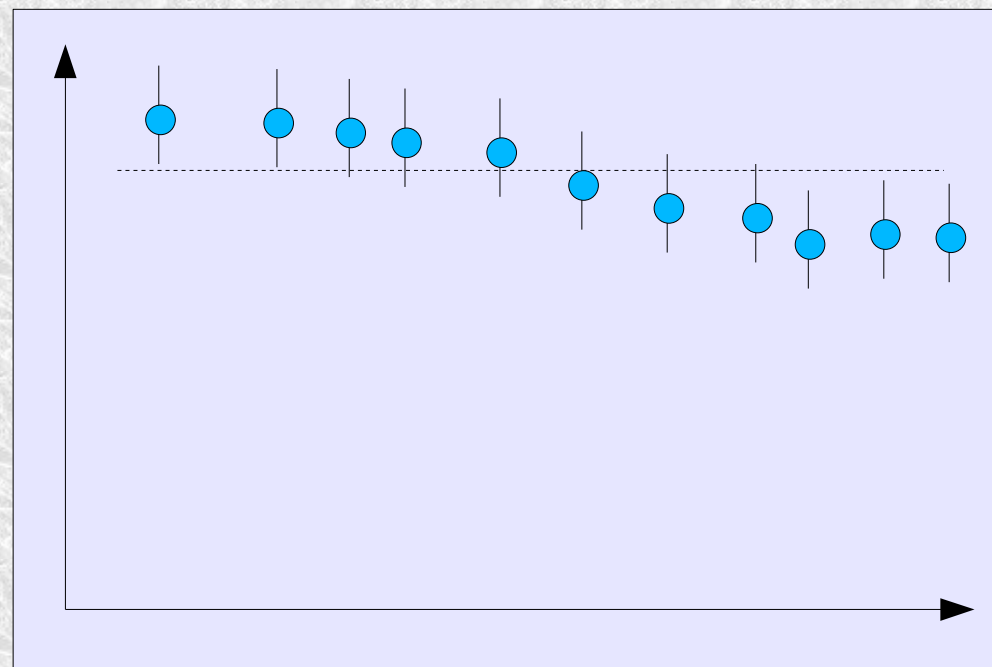
- Are two 1D distributions drawn from a common parent?
- Find integral distribution
- Observe maximum difference between two integrals
- Probability is calculable
- Very bad here!
- K-S very good
  - No allowance for fit
  - Or systematic error





# Run test

- Measures 'runs': How often one distribution is above (or below) the other
  - Find length of maximum run
  - Probability of this length is calculable
- This data has good  $\chi^2$
- Run test would be poor





# Is there a better test than $\chi^2$ ?

- Yes...and no
- For any given problem, a more sensitive test can be defined
- But you can only define the sensitivity if you know what you are choosing between
- There is no 'most powerful' goodness of fit test.



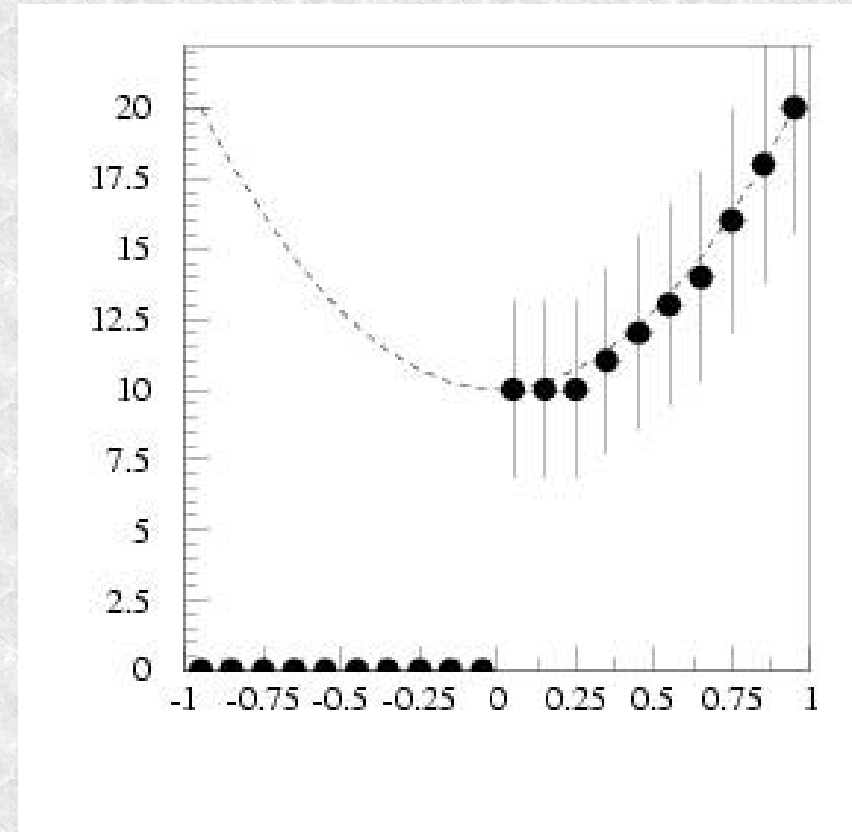
# How things can go wrong...

- Example: Fit function

$$y = a + b \cos \theta^2$$

- Likelihood fit is not sensitive to sign on  $\theta$
- Concludes that dist<sup>n</sup> shown is very good fit!

Nothing is perfect...





# Conclusions

- The likelihood ratio underpins everything
  - Use it
- Cost of computing becoming important
  - Optimal methods not *necessarily* optimal
  - But the data is very expensive too.
  - Systematic errors may dominate anyway
- Need to see statistics as a tool
- Please:
  - Ask me if you have statistical questions