

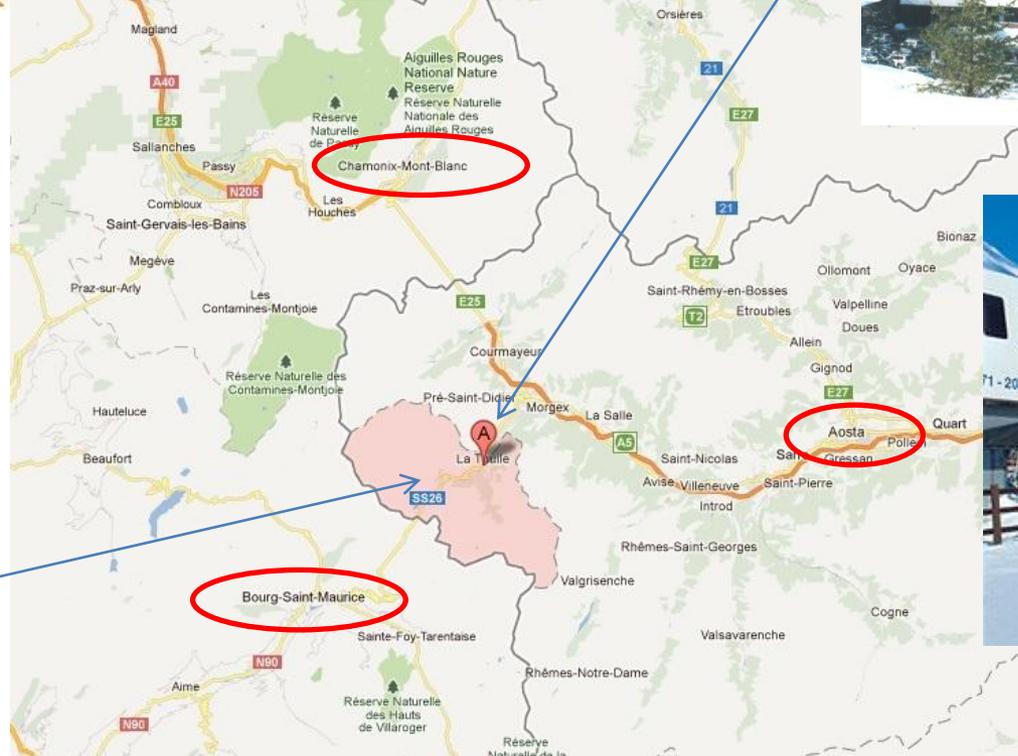
XXVI Rencontres de Physique de la Vallée d'Aoste  
Feb 26th-Mar 3rd 2012, La Thuile, Italy



La Thuile, Aosta Valley, Italy  
February 26 - March 3, 2012

Julie Kirk

<http://agenda.infn.it/conferenceOtherViews.py?view=standard&confId=4116>



Last at La Thuile in 1994 much is still the same:

About 120 participants  
2 sessions a day (8.30-11.30 and 4.30-7.30) covering:

- Cosmology and Astrophysics
- Neutrino Physics and Lepton Flavour Violation
- QCD Physics/Hadronic Interactions
- Flavour Physics/CP Violation and Rare Decays
- Electroweak and Top Physics
- Higgs Searches
- Searching for New Physics
- Perspectives

- Physics and Society
- Young Scientist Forum

Thorium reactors  
Physics in Latin America

New this year:  
Experiments asked to nominate good, young scientists to talk about their research.

What can I do from 11.30 'til 4.30??



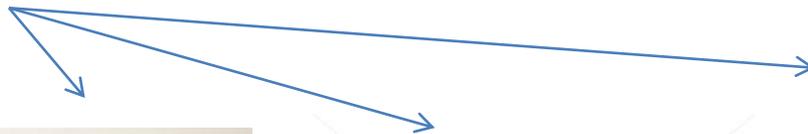
Each session about 7-8 20 minute talks with questions  
→ [selected highlights here](#)

# To Wi-Fi or NOT to Wi-Fi ??

No Wi-fi in conference room (7<sup>th</sup> floor). Good as people pay attention to the talks....



..BUT attendance not great at some sessions. Not everyone was out skiing

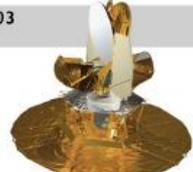




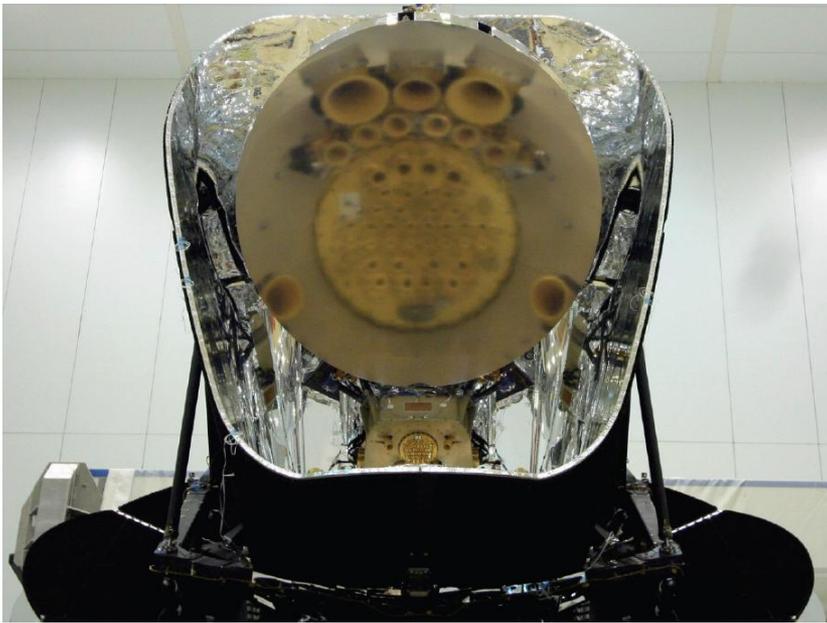
Summary (sadly)  
“Come back next week”  
i.e. for Moriond!

# Planck – Cosmic Microwave Background

## Cosmic Microwave Background Radiation Overview

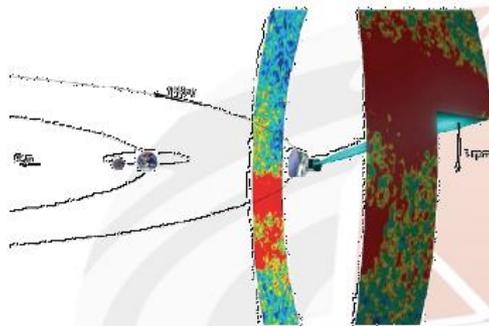
1965		Penzias and Wilson	<b>The oldest light or the first light of the Universe</b> Discovered the remnant afterglow from the <b>Big Bang</b> . → <b>2.7 K</b>
1992		COBE	<b>Blackbody radiation,</b> Discovered the patterns ( <b>anisotropy</b> ) in the afterglow. → <b>angular scale ~ 7°</b> at a level $\Delta T/T$ of $10^{-5}$
2003		WMAP (Wilkinson Microwave Anisotropy Probe):	→ <b>angular scale ~ 15'</b>
2009		Planck	→ <b>angular scale ~ 5',</b> $\Delta T/T \sim 2 \times 10^{-6}$ , 30~867 Hz

- 3<sup>rd</sup> generation of CMB experiment:
- Better angular resolution
  - Larger frequency range



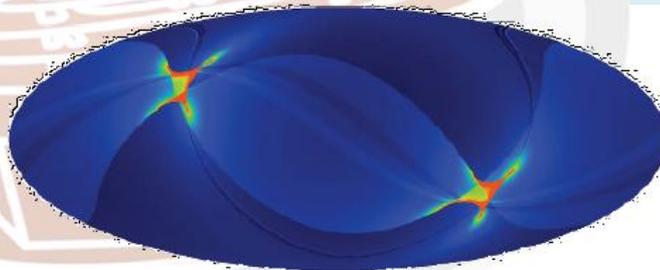
## Current Status

- 1019 days since launch.
- Satellite and instruments have been working nominally and continuously since start of sky surveys (mid August 2009)
  - HFI ran out of He on 14 Jan 2012 and stopped taking scientifically useful data
  - LFI still data gathering (several months )
- All the sky has been surveyed about five times with both instruments. LFI is already into its sixth sky survey.



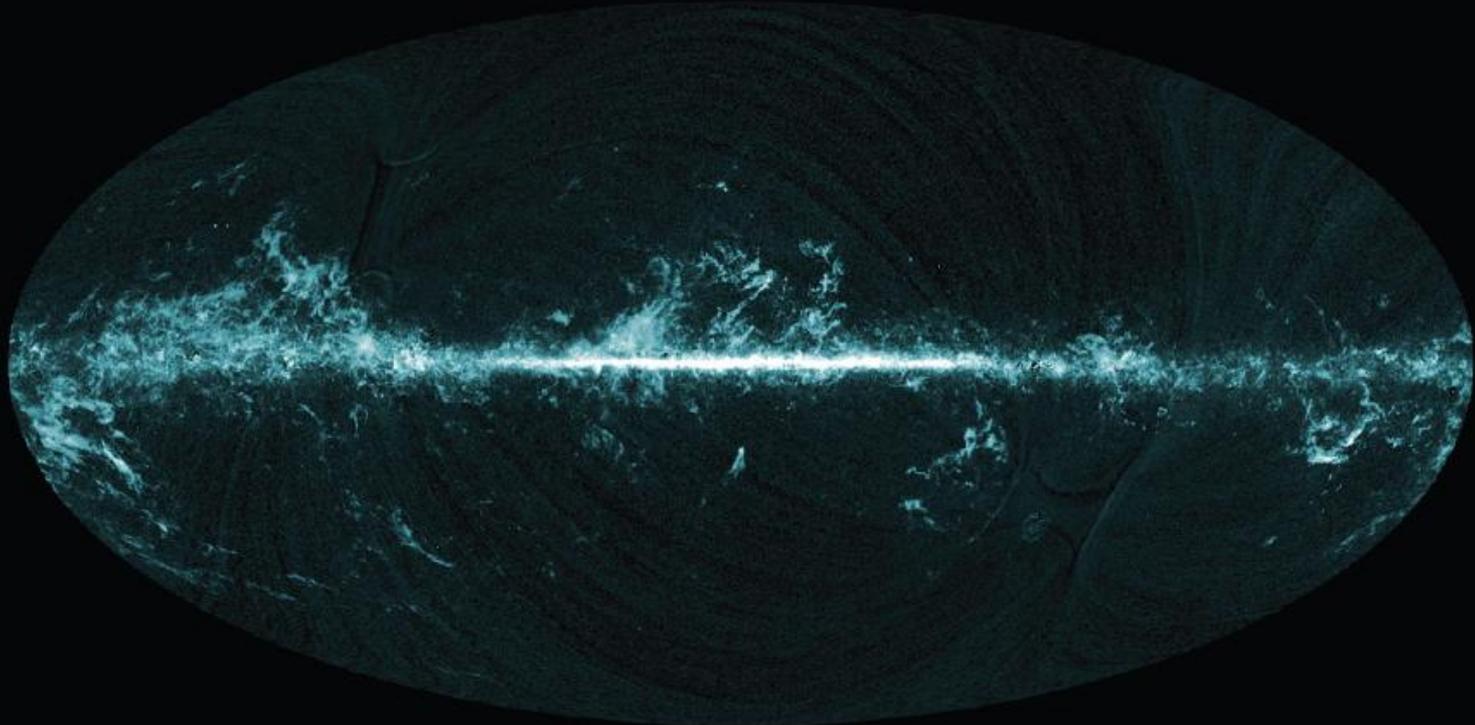
Planck is a  
survey mission

**About 6 months are  
needed to cover  
~95% of the sky.**

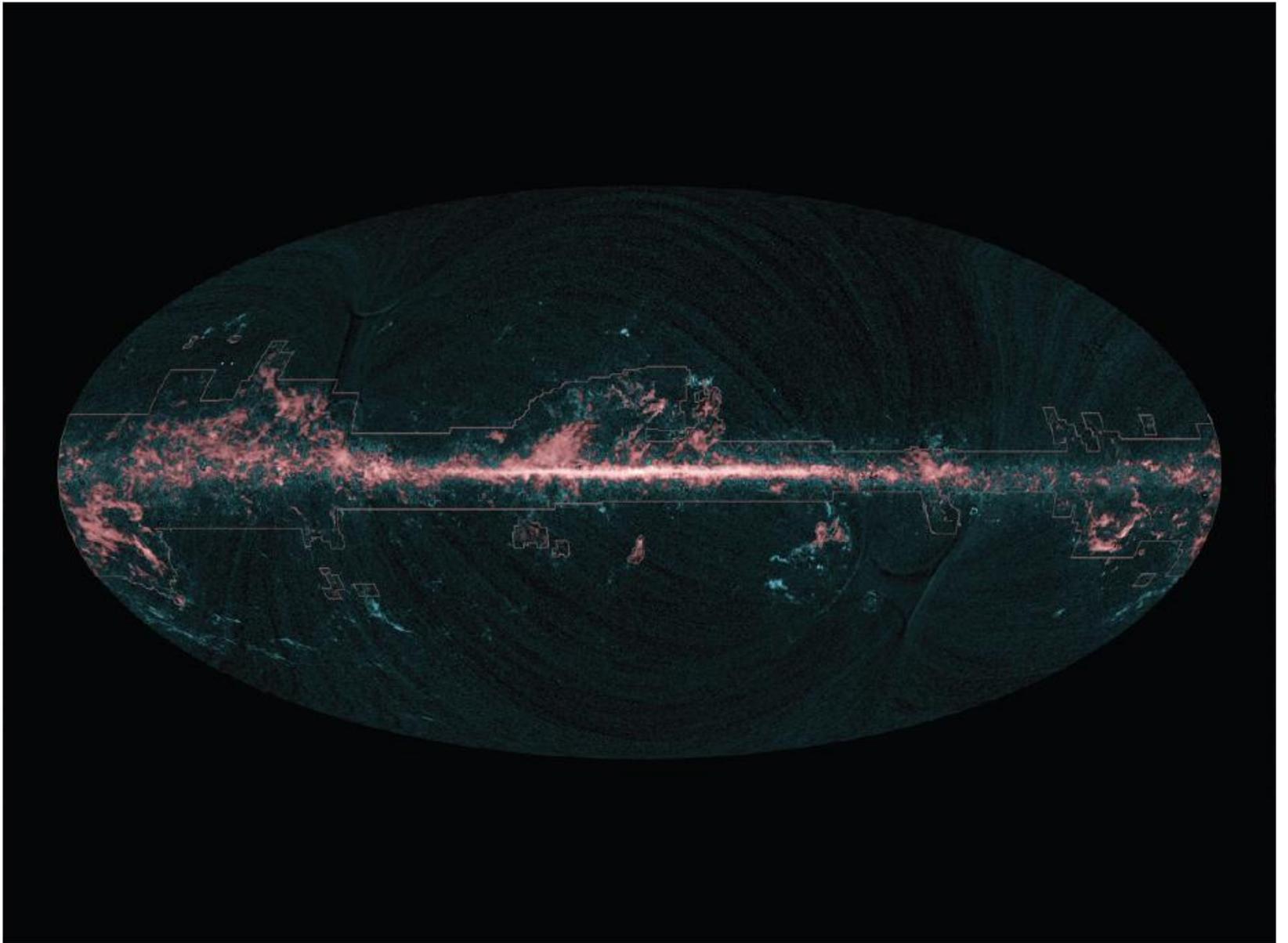


Current results of interest to astrophysics  
Expect results from cosmological  
programme in 2013

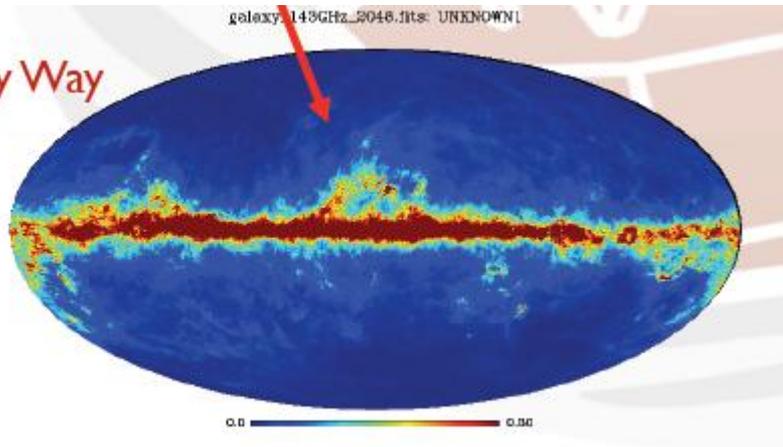
# First full sky CO Map



Shows star forming regions



## The Milky Way

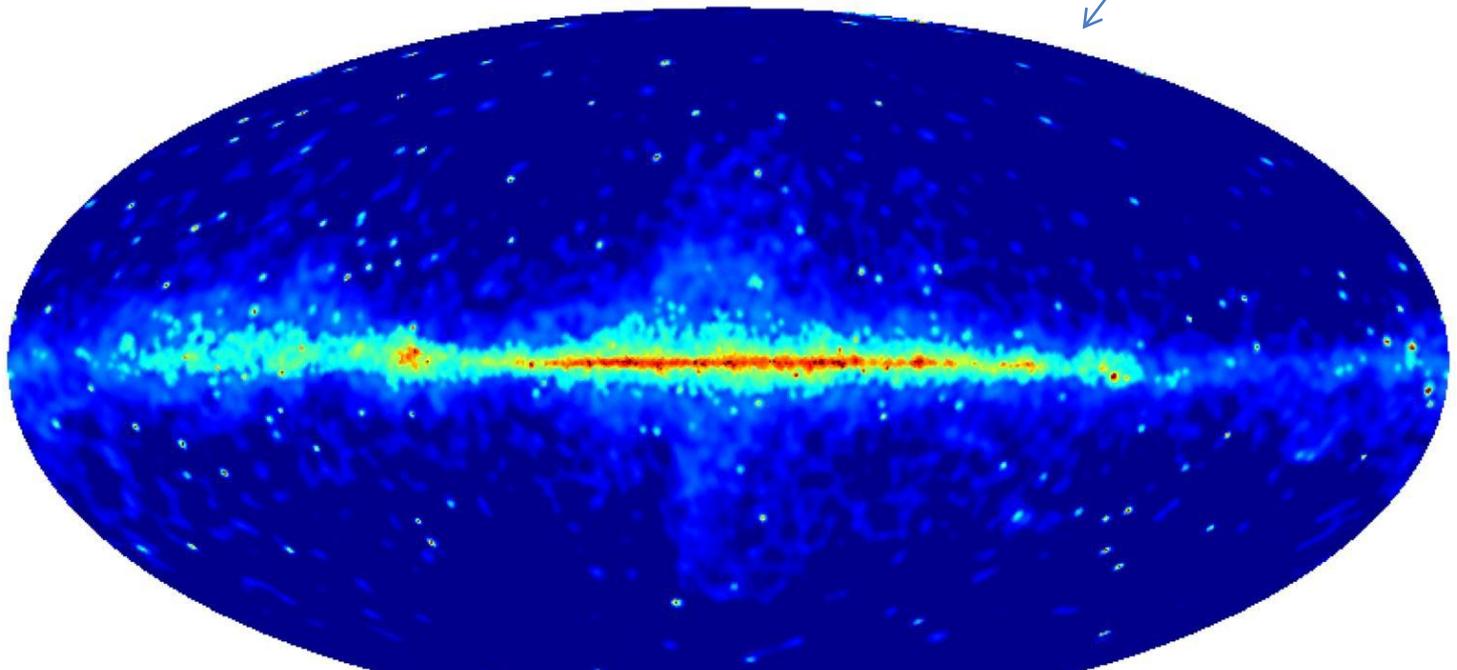


## Cosmic Haze:

Synchrotron emission from centre of galaxy – not consistent with expectations.

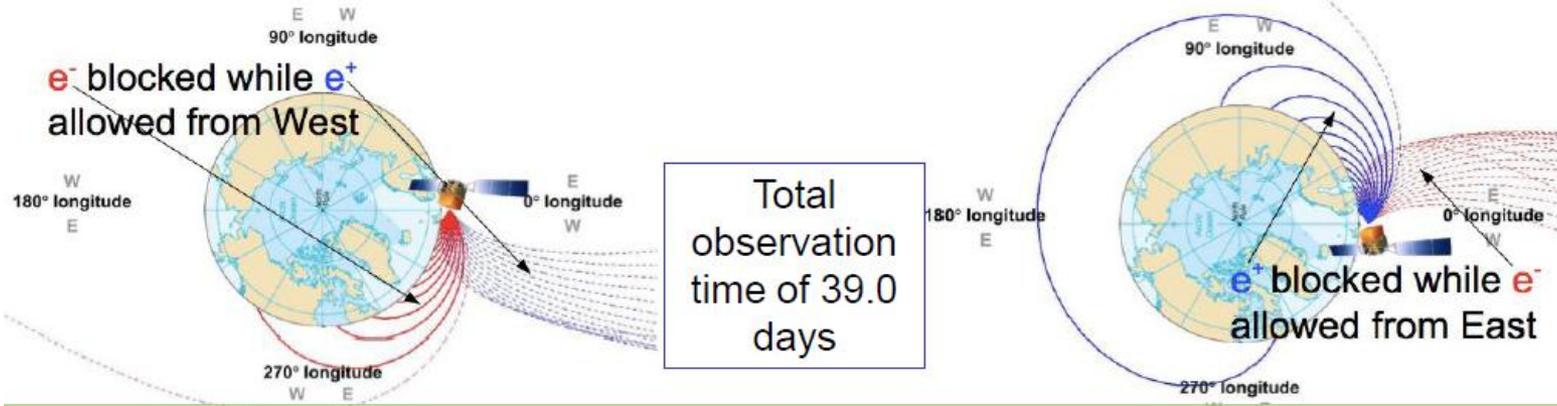
SN/ new CR component/ DM??

Also seen by Fermi

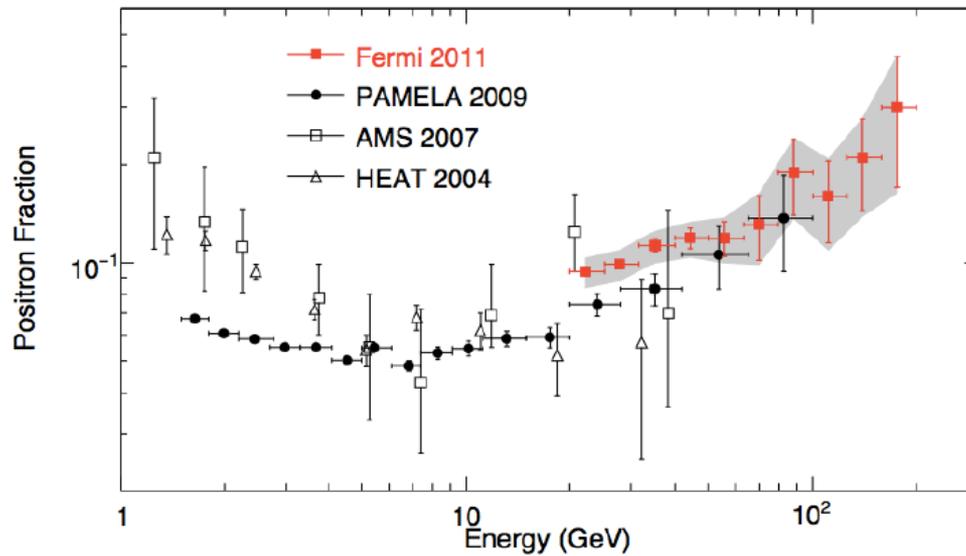


**3 years smoothed count map above 10 GeV**

# Electron-positron distinction with Fermi



Positron fraction rises with energy. Confirms PAMELA result



# Dark forces – Kloe.....

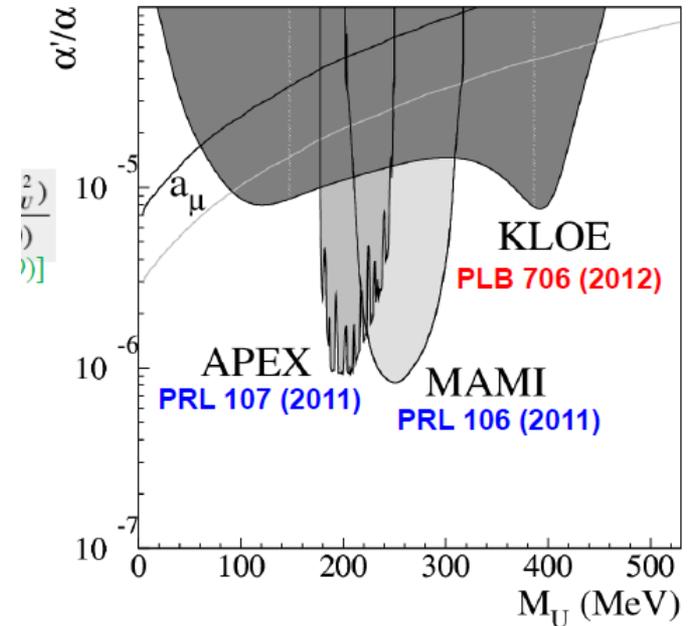
## Low energy dark forces

Several unexpected astrophysical observations (PAMELA, ATIC, INTEGRAL, DAMA/LIBRA, CoGent...) could be explained with the existence of a hidden gauge sector weakly coupled with SM through a mixing mechanism of a new gauge boson ( $U, A', V...$ ) with the photon:



[Arkani-Hamed et al. PRD79 015014 (2009)]  
[Essig et al., PRD80 015003 (2009)]

- ✓ U mass range: **1 MeV – few GeV**
- ✓ Coupling constant of electric charge to U:  $\epsilon \leq 10^{-3}$
- ✓ U production/decay through photon mixing



S. Giovannella – La Thuile 2012 – KLOE searches on Dark Forces

2

- ✗ The most advanced search is through the  $\phi \rightarrow \eta U, \eta \rightarrow \pi^+ \pi^- \pi^0$ , decay, which allow to set a limit of:

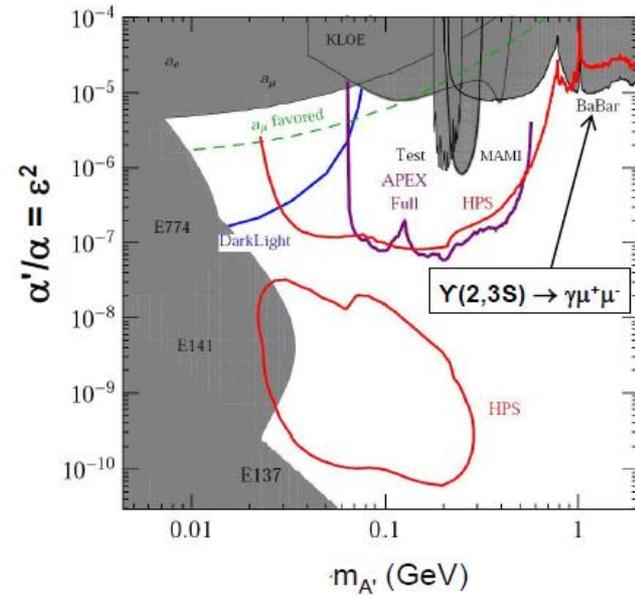
$$\alpha' / \alpha \leq 2 \times 10^{-5} \text{ @ 90\% C.L. for } 50 < M_U < 420 \text{ MeV}$$

⇒ A dark photon can be readily produced in

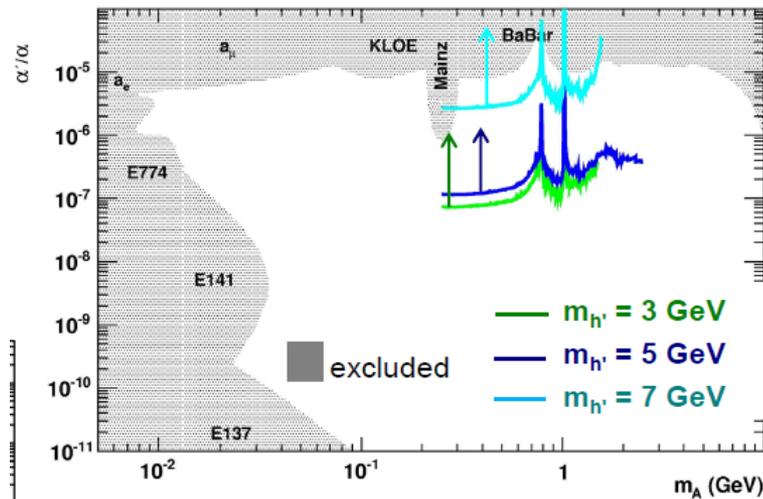
$$e^+e^- \rightarrow \gamma A', A' \rightarrow f \bar{f}$$

The limits on  $e^+e^- \rightarrow Y(2S,3S) \rightarrow \gamma \mu^+ \mu^-$  can be reinterpreted as limits on dark photon production.

Constraints on  $\alpha'/\alpha = \varepsilon^2$



Limit on  $\varepsilon^2 = \alpha'/\alpha$  for various Higgs mass (assuming  $\alpha_D = \alpha_{em}$ )



Substantial improvement over existing limits for  $m_{h'} < 5 - 7$  GeV if light dark Higgs boson exists

⇒ The Higgs'-strahlung process

$$e^+e^- \rightarrow A'^* \rightarrow h' A', h' \rightarrow A' A'$$

is very interesting, as it is **only suppressed by  $\varepsilon^2$**  and is expected to have a **very small background**.

$\Gamma_{h'}$  (GeV)

# Neutrino Physics

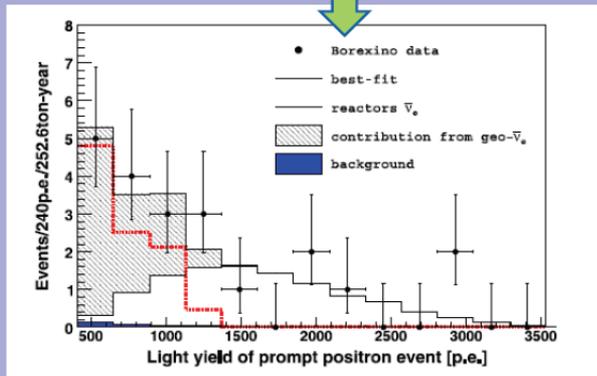
Search for delayed coincidences in the Borexino detector  
Main background sources:

- Li-9, He-8, untagged muons, accidentals.....

$$3.9^{+1.6}_{-1.3} \left( \begin{matrix} +5.8 \\ -3.2 \end{matrix} \right) \text{ ev}/(100 \text{ tons} \cdot \text{yr})$$

68.3 %CL

99.73 %CL



La Thuile - 27 Feb 2012

15

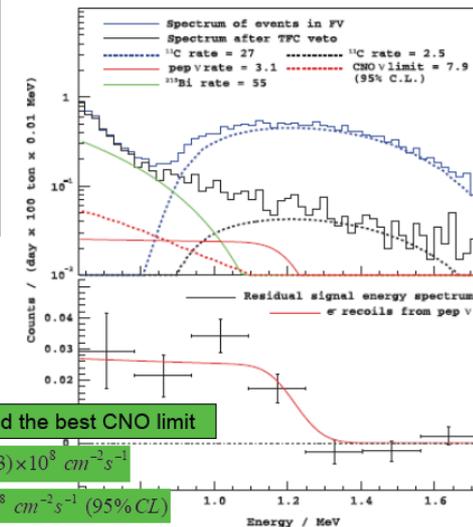
## BOREXINO

### Geo-neutrinos

AntiNeutrinos emitted in beta decays of naturally occurring radioactive isotopes in the Earth's crust and mantle

C-11 reduction strategy:

- Threefold coincidence (muon,neutron,C11)
- Pulse shape discrimination electron/gamma/positron (Ps formation)



First pep measurement and the best CNO limit

$$\Phi_{\text{pep}} (\text{MSW} - \text{LMA}) = (1.6 \pm 0.3) \times 10^8 \text{ cm}^{-2} \text{ s}^{-1}$$

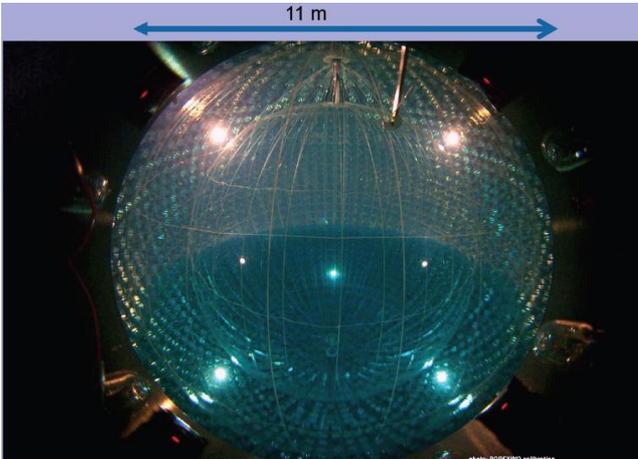
$$\Phi_{\text{CNO}} (\text{MSW} - \text{LMA}) < 7.7 \times 10^8 \text{ cm}^{-2} \text{ s}^{-1} \text{ (95\% CL)}$$

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20

P-e-p

11 m

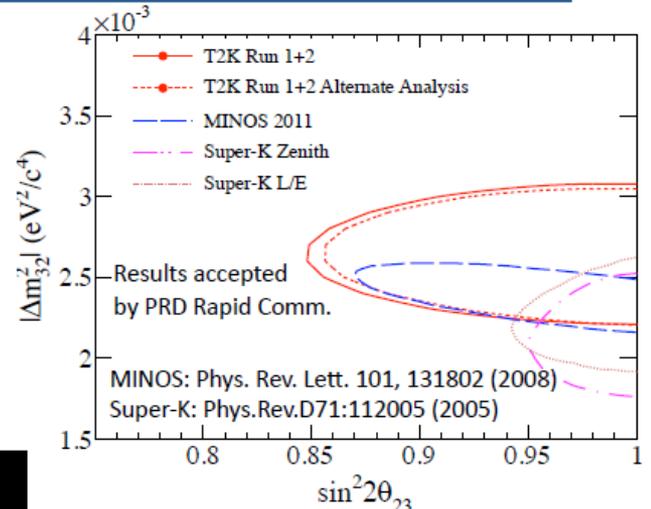
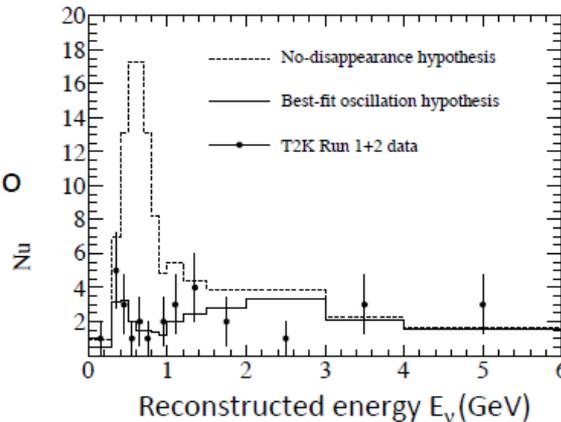


# T2K

Open questions:

- Is  $\theta_{23}$  exactly 45 degrees, or not?
- Is  $\theta_{13}$  zero, or just small?
- Is there CP violation in the neutrino sector?

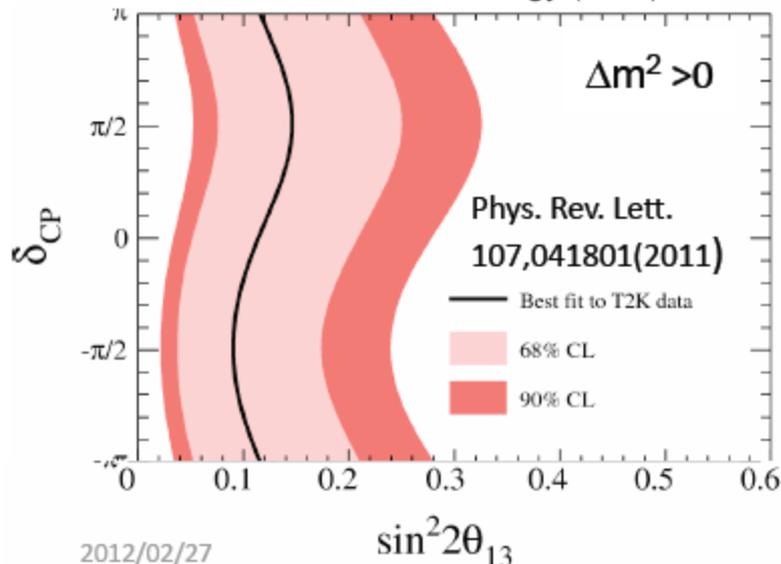
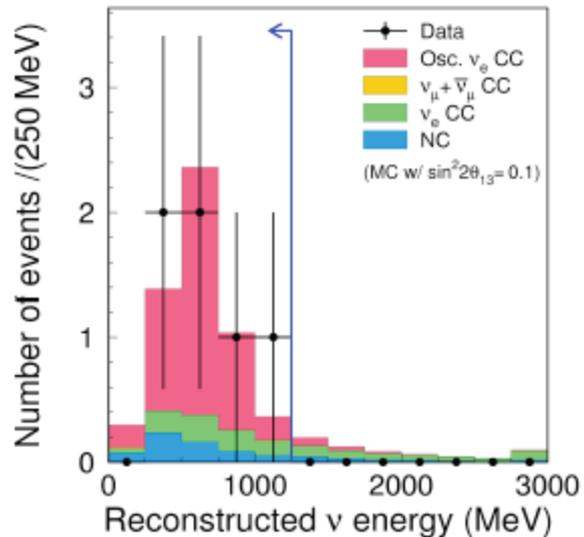
## $\nu_\mu$ disappearance results



Summary of uncertainties	$\nu_\mu$ signal $\Delta m^2_{23} = 2.4 \times 10^{-3} \text{ eV}^2$ $\sin^2 2\theta_{23} = 1.0$
$\nu$ flux	$\pm 4.8\%$
$\nu$ interactions	$+8.3 - 8.1\%$
Near detector	$+6.2 - 5.9\%$
Far detector	$\pm 10.3\%$
Total	$+15.4 - 15.1\%$

- 31 events pass  $\nu_\mu$  selection criterion
- $103.6^{+13.8}_{-13.4}$  expected for no osc, excluded at  $4.5\sigma$
- Fit  $E_\nu$  distribution for 2 flavor osc. parameters (binned  $\chi^2$  fit)
- **Best fit:  $|\Delta m^2_{32}| = 2.65 \times 10^{-3} \text{ eV}^2$   
 $\sin^2 2\theta_{23} = 0.98$**

# $\nu_e$ appearance results



- 6 candidate events observed for background of  $1.49 \pm 0.34$
- Probability to see 6 events or more for  $\sin^2 2\theta_{13}=0$  is 0.007 ( $2.5\sigma$  equivalent)
- For  $|\Delta m_{32}^2|=2.4 \times 10^{-3} \text{ eV}^2$ ;  $\sin^2 2\theta_{23} = 1$   
Best fit:  $\sin^2 2\theta_{13} = 0.11$   
 $0.03 < \sin^2 2\theta_{13} < 0.28$  at 90% C.L.
- Recent results are consistent with T2K:  
MINOS:  $\sin^2 2\theta_{13} < 0.12$  at 90% C.L.  
Double Chooz:  $0.017 < \sin^2 2\theta_{13} < 0.16$
- Combined fit to all three experiments:  
 $\sin^2 2\theta_{13} > 0$  at  $\sim 3\sigma$

References:

MINOS: Phys. Rev. Lett. 107, 181802 (2011)

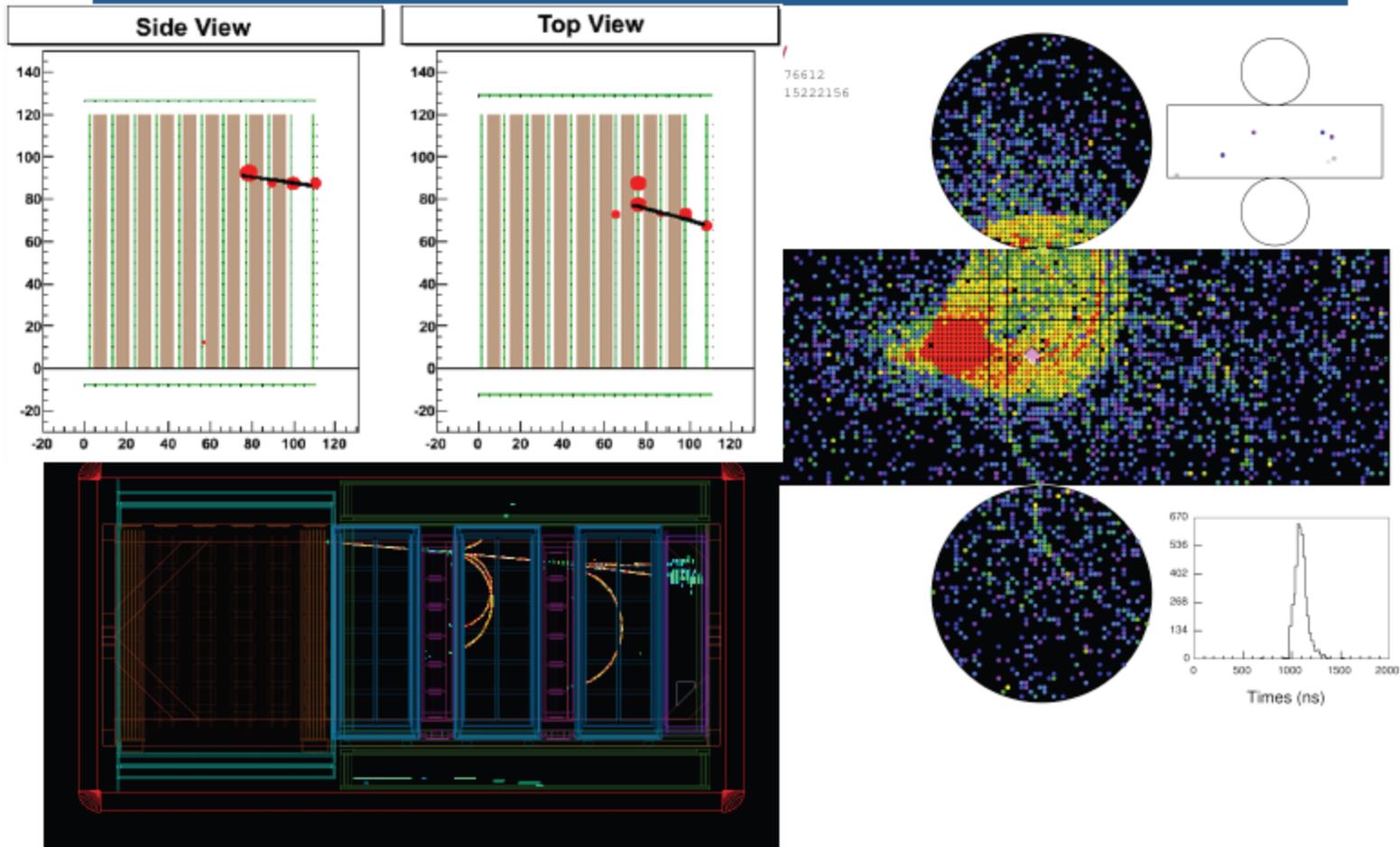
Double Chooz: hep-ex/1112.6353

Global fit: hep-ph/1111.3330

2012/02/27

21

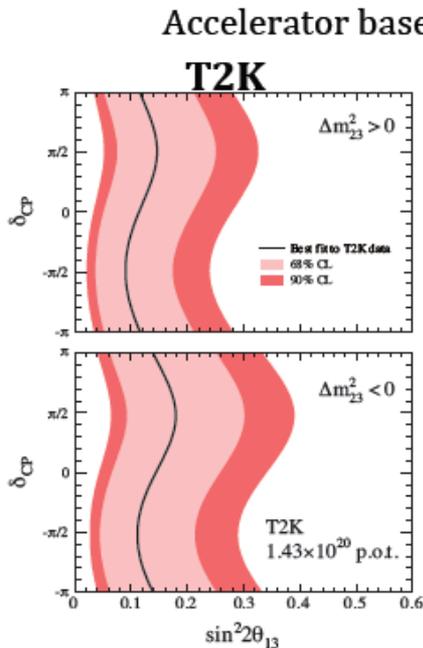
# First neutrinos post-earthquake!



Enormous amount of work by collaborators, labs (KEK, J-PARC) and funding agencies) to make this happen, to which we are very grateful

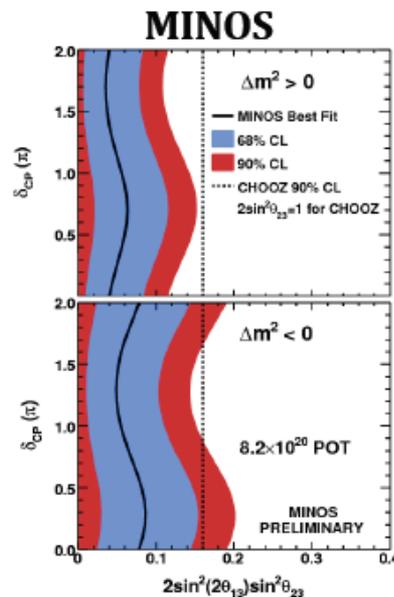
# DAYA BAY

## Recent Experimental Results (2011): $\theta_{13}$ may be large



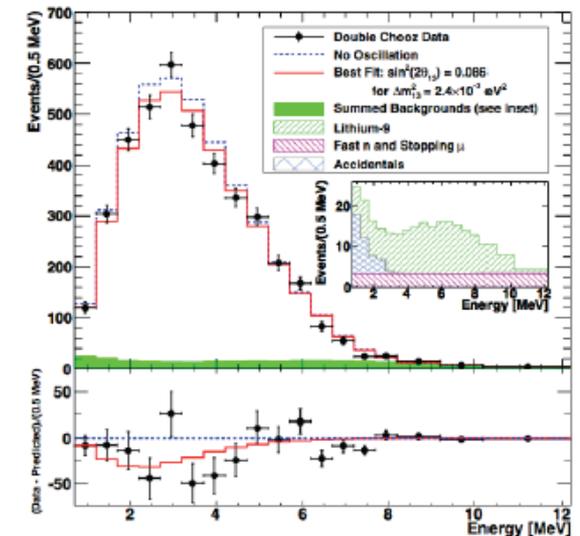
PRL 107, 041801 (2011)

$0.03(0.04) < \sin^2 2\theta_{13} < 0.28(0.34)$  at 90% CL



PRL 107, 181802 (2011)

## Reactor based disappearance expt Double Chooz



arXiv:1112.6353v2

$\sin^2 2\theta_{13} = 0.086 \pm 0.041(\text{stat}) \pm 0.030(\text{sys})$   
( $0.015 < \sin^2 2\theta_{13} < 0.16$  at 90% CL)

# Daya Bay Design Principle

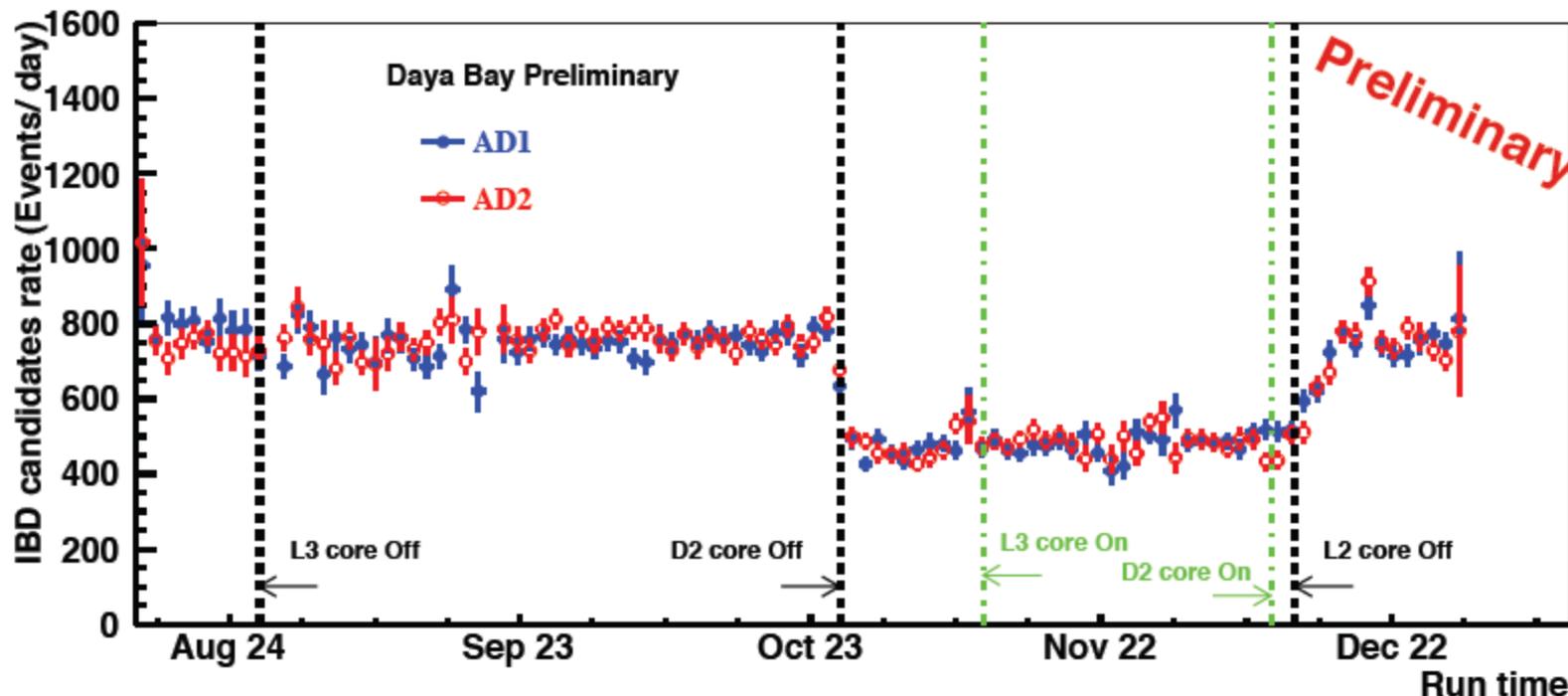
- Identical near and far detectors cancel many systematic errors
- Multiple modules boost statistics while reducing systematic errors with multiple independent measurements
- Three zone detector design eliminates the need for spatial cuts which can introduce systematic uncertainties
- Shielding from cosmic rays and natural radioactivity reduces background rates
- Movable detectors allow possible cross calibration between near and far detectors to further reduce systematic errors.

## Experimental Layout

	Overburden	D. B.	L. A.	L. A. II
EH1	280	360	860	1310
EH2	300	1350	480	530
EH3	880	1910	1540	1550
	mwe	m	m	m



# IBD Candidate Rate



Dashed lines indicate reactor shut down or turn on.

Expected events (/day/detector)

	DYB Site	LA Site	Far Site
IBD Evt	840	760	90
BKG Evt	<0.6%	<0.5%	<0.4%

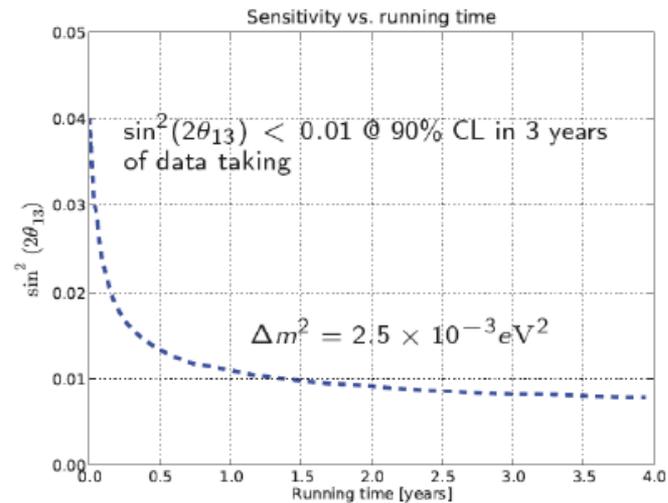
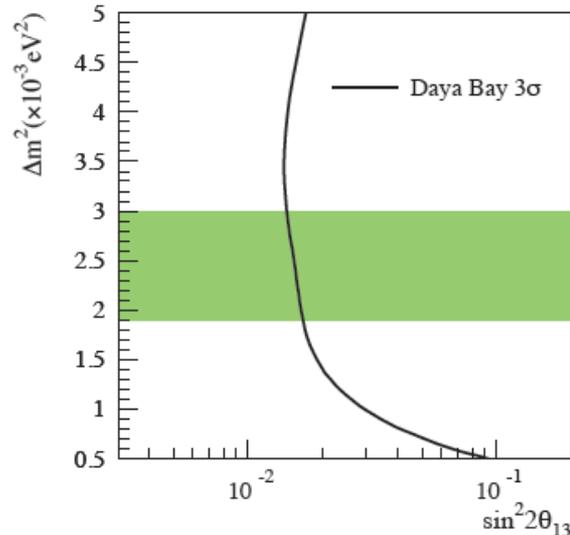
Expected events with all cores ON



“2 months of running to measure  $\theta_{13}$  as non-zero”



## Summary

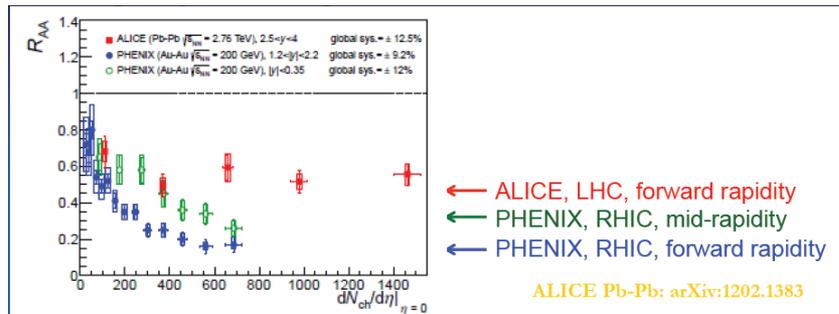
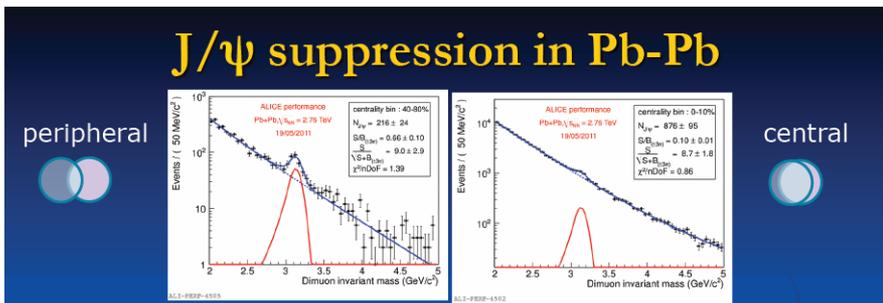


- We have demonstrated the identicalness of the first two detectors at Daya Bay near site
- With all 8 detectors running by summer 2012, Daya Bay is going to provide the most precise measurement of  $\theta_{13}$ .

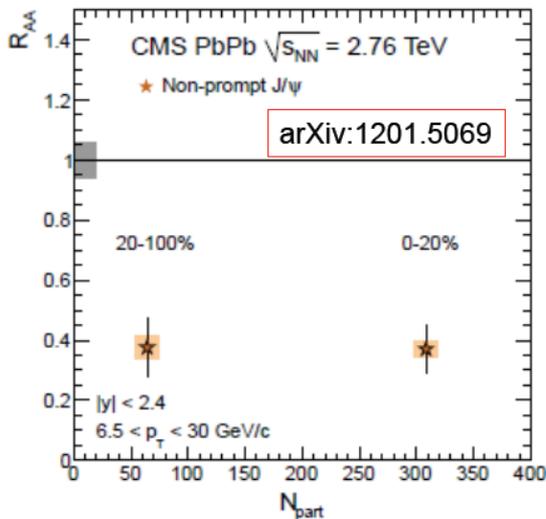
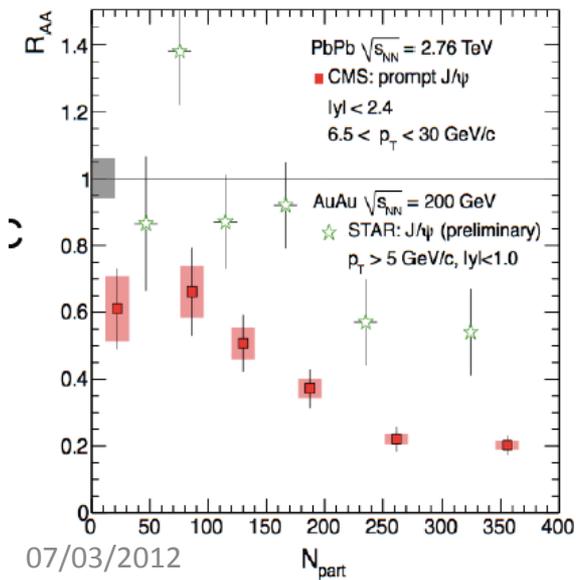


# Heavy Ions

(about 30 people in audience and no chairperson!)



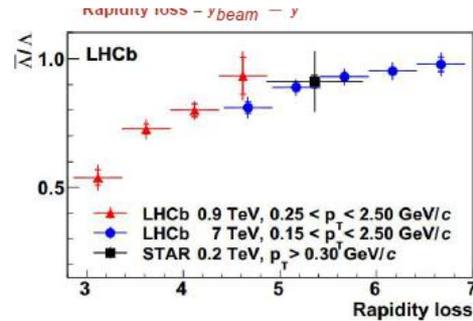
- ALICE: less suppression than RHIC at forward rapidity
- Similar suppression (not for central collisions) as RHIC at midrapidity



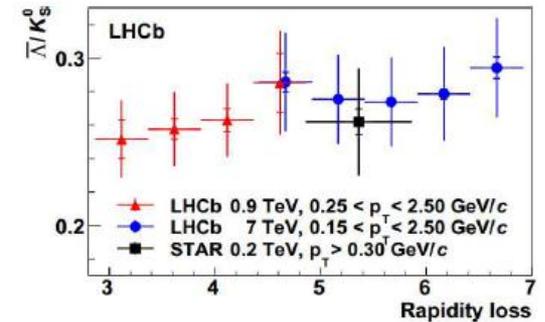
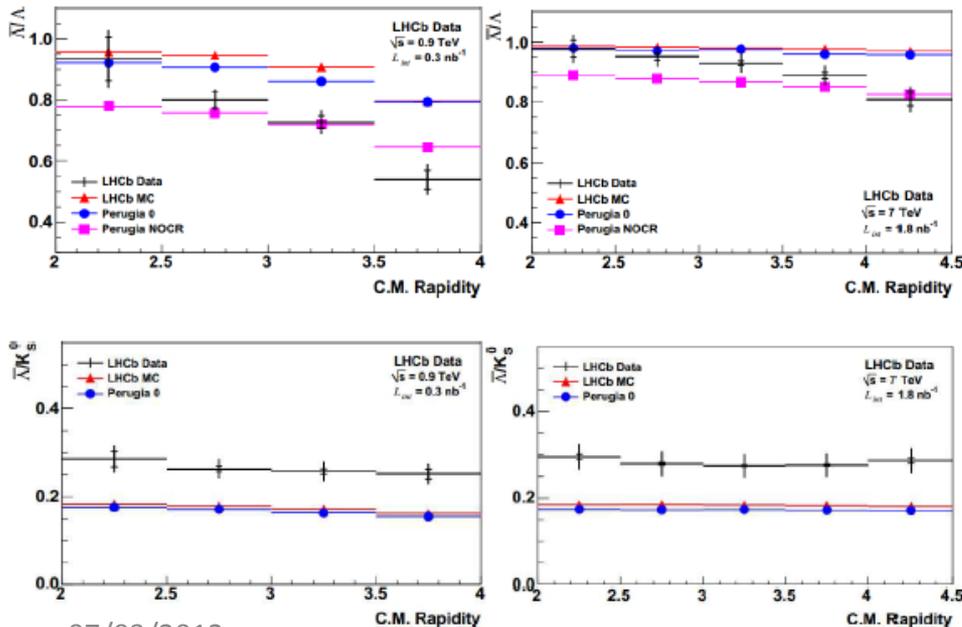
- First measurement in heavy-ion collisions
- J/ψ coming from B decay are strongly suppressed

# Production ratio: $\bar{\Lambda}/\Lambda$ , $\bar{\Lambda}/K_S^0$

- $\bar{\Lambda}/\Lambda = \frac{\sigma(pp \rightarrow \bar{\Lambda} X)}{\sigma(pp \rightarrow \Lambda X)}$  probes the baryon number transport.
- $\bar{\Lambda}/K_S^0 = \frac{\sigma(pp \rightarrow \bar{\Lambda} X)}{\sigma(pp \rightarrow K_S^0 X)}$  probes the strange baryon suppression.

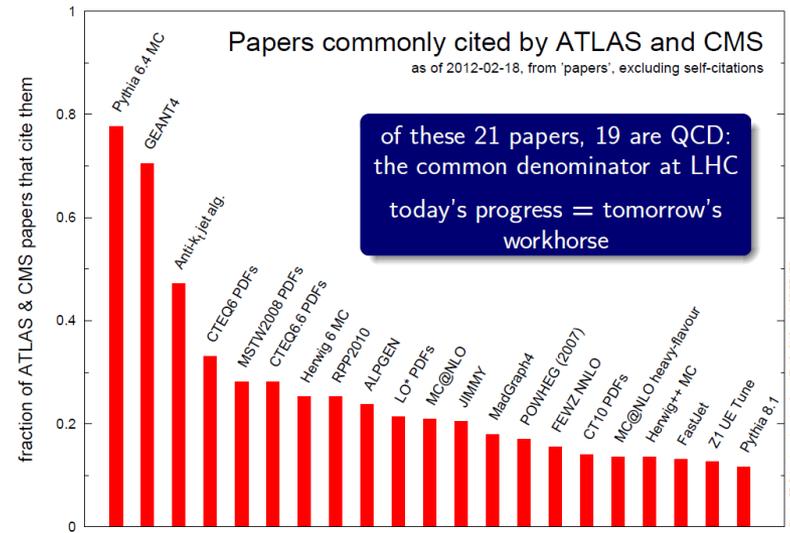
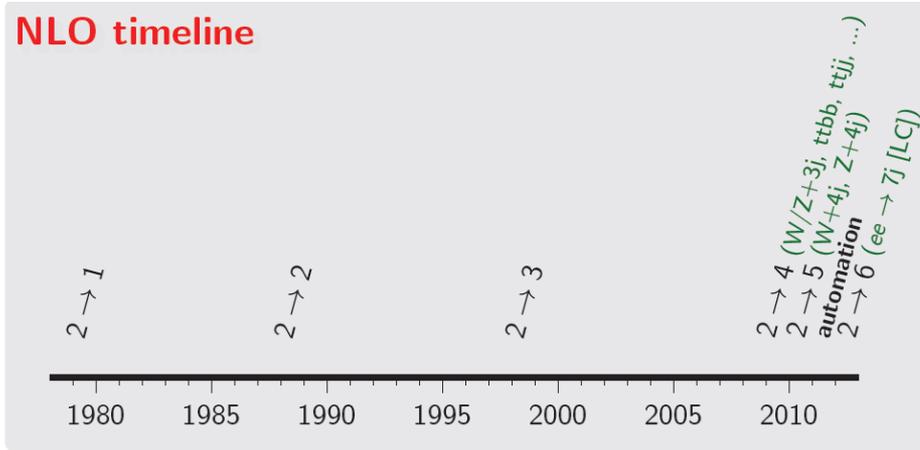


- Consistency between the two measurements ( $\sqrt{s} = 900$  GeV and 7 TeV) and the previous results.



- $\bar{\Lambda}/\Lambda$ : good agreement with Perugia0 at low rapidity while at high rapidity Perugia NOCR looks to be favoured.
- $\bar{\Lambda}/K_S^0$  measured ratio is significantly larger than predicted by the generators, i.e. more baryons are produced in strange hadronization than expected.

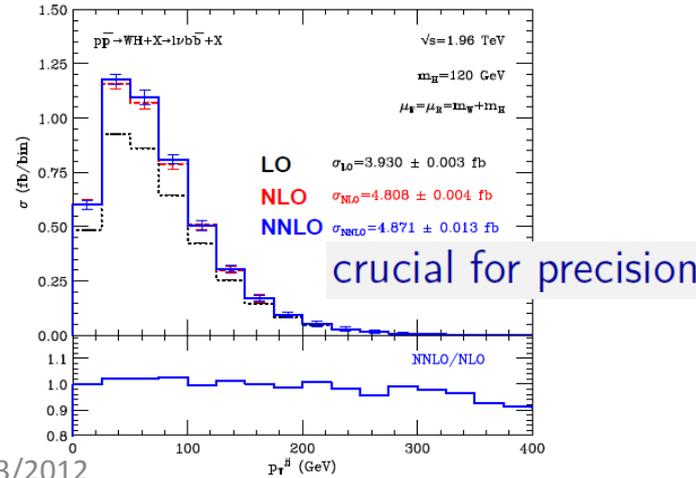
## NLO timeline



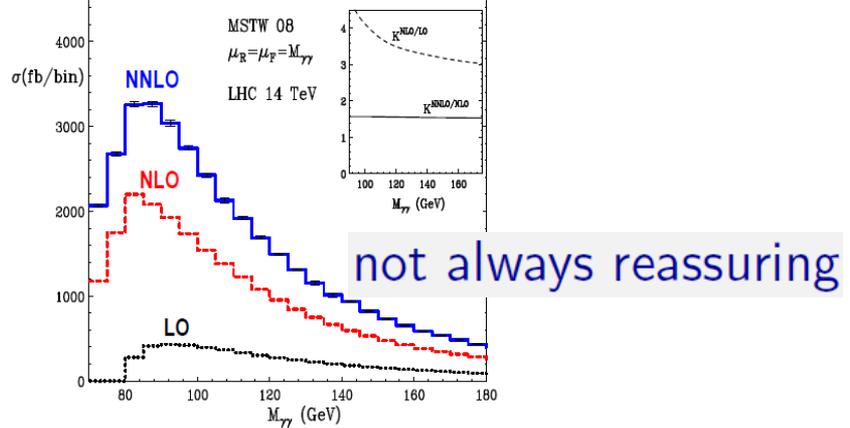
High precision — NNLO — is crucial for key processes, but not yet always available:

- ✓  $W, Z, \text{Higgs}, \gamma\gamma, \text{VBF}, \text{VH}$
- ✗  $VV, t\bar{t}, \text{inclusive jets, etc.}$

## New in 2011: NNLO WH (differential)



## New in 2011: NNLO $\gamma\gamma$



Important also to develop methods so that we're less sensitive to limits on our precision. Generally by finding ways to distinguish signals from the background more efficiently, i.e. increasing  $S/B$ .

# NEWS SCIENCE & ENVIRONMENT

29 February 2012 Last updated at 17:40 GMT

890 Share f t e

## New twist in antimatter mystery

By Paul Rincon  
Science editor, BBC News website

Physicists have taken a step forward in their efforts to understand why the Universe is dominated by matter, and not its shadowy opposite antimatter.

A US experiment has confirmed previous findings that hinted at phenomena outside our understanding of physics.

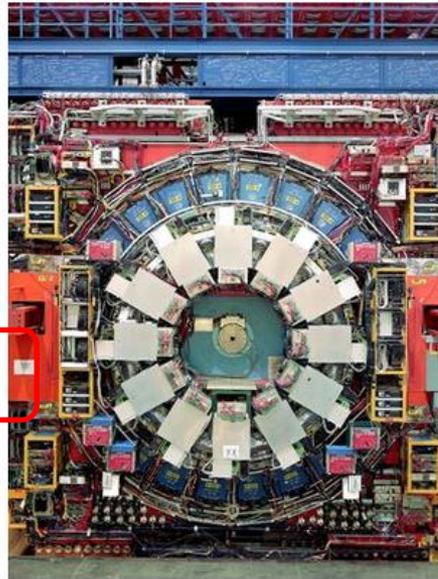
The results show that certain matter particles decay differently from their antimatter counterparts.

Such differences could potentially help explain why there is so much more matter in the cosmos than antimatter.

The findings from scientists working on the CDF experiment have been presented at a particle physics meeting in La Thuile, Italy.

CDF was one of two multi-purpose experiments at the now-defunct Tevatron particle smasher in Illinois.

Physicists think the intense heat of the Big Bang should have forged equal amounts of matter and its "mirror image" antimatter. Yet today we live in a Universe composed overwhelmingly of matter.



CDF was one of two multi-purpose experiments at the US Tevatron accelerator near Chicago

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[BP reaches \\$7.8bn oil spill deal](#)

[Capsized Costa ship hearing opens](#)

### Features & Analysis



#### Trees or SUVs

Five interpretations of Dr Seuss's The Lorax



#### Tough times

Power cuts and candle shortages blight daily life in Syria



#### 52 mouths to feed

How many orphans could you take into your home?



#### Year of change

John Simpson on new leadership among key powers

# CP violation

CDF updated  $\Delta A_{CP}$  using full runII dataset (LHCb result shown in PPD seminar last week)

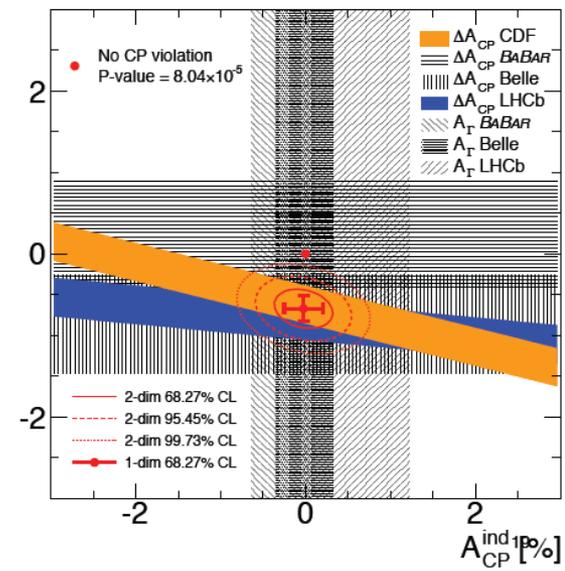
- New CDF result confirms LHCb result: same resolution,  $<1\sigma$  difference in central value  

$$\Delta A_{CP}(\text{LHCb}) = [-0.62 \pm 0.21 \pm 0.10]\%$$
- When combining à la HFAG with other available measurements, no CPV point is at  $\sim 3.8\sigma$  and  

$$\Delta A_{CP}^{\text{dir}} = (-0.67 \pm 0.16)\%$$

$$A_{CP}^{\text{ind}} = (-0.02 \pm 0.22)\%$$
- CDF will also produce separate measurements of  $A_{CP}(K^+K^-)$  and  $A_{CP}(\pi^+\pi^-)$

More information in [CDF Public Note 10784](#)



**Results**

$B_S^0 \rightarrow J/\psi\phi$

Most precise measurement<sub>INTS</sub>

$\Gamma_s = 0.656 \pm 0.009 \text{ (stat)} \pm 0.008 \text{ (syst)} \text{ ps}^{-1}$

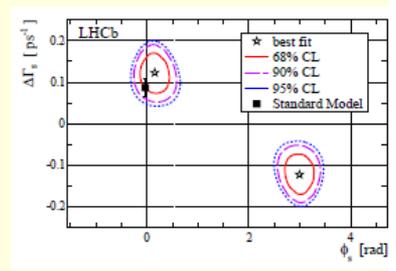
$\phi_s = 0.15 \pm 0.18 \text{ (stat)} \pm 0.07 \text{ (syst)} \text{ rad}$

$\Delta\Gamma_s = 0.123 \pm 0.029 \text{ (stat)} \pm 0.011 \text{ (syst)} \text{ ps}^{-1}$

First direct evidence of non-zero  $\Delta\Gamma_s$

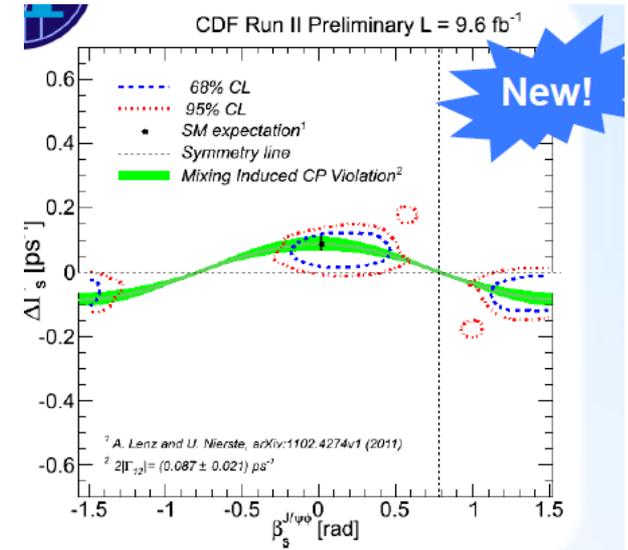
2 solutions, 1 in agreement with SM prediction

→ Current uncertainties still leave room for New Physics



COMBINATION OF  $B_S^0 \rightarrow J/\psi\phi$  AND  $B_S^0 \rightarrow J/\psi f_0$

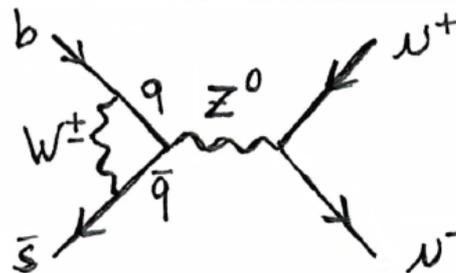
$$\phi_s = 0.03 \pm 0.16 \text{ (stat)} \pm 0.07 \text{ (syst)} \text{ rad}$$



# Rare Decays

sensitive to NP

▪ Z-penguin suppressed diagram  
doubly suppressed : FCNC & helicity



[The CKMFitter group - hep-ph 1106.4041]

▪ SM-expected Branching Ratio :

$$B_{SM}(B_s \rightarrow \mu^+ \mu^-) = (3.64^{+0.17}_{-0.31}) \times 10^{-9}$$

$$B_{SM}(B_d \rightarrow \mu^+ \mu^-) = (1.13^{+0.06}_{-0.11}) \times 10^{-10}$$

Previously..

NEW at La Thuile

## Experimental status

- 
**CDF**  $Br(B_s \rightarrow \mu\mu) = 1.8^{+1.1}_{-0.9} \times 10^{-8}$   
PRL107, 191801 (2011)
- 
**DØ**  $Br(B_s \rightarrow \mu\mu) < 5.1 \times 10^{-8}$  @95% C.L.  
PLB693, 539 (2010)
- 
**LHCb**  $Br(B_s \rightarrow \mu\mu) < 1.4 \times 10^{-8}$  @95% C.L.  
PLB 708, 55 (2012)
- 
**CMS**  $Br(B_s \rightarrow \mu\mu) < 1.9 \times 10^{-8}$  @95% C.L.  
PRL107, 191802 (2011)

7 → 10 fb<sup>-1</sup>

$$BR(B_s \rightarrow \mu^+ \mu^-) = (1.3^{+0.9}_{-0.7}) \times 10^{-8}$$

$$0.8 \times 10^{-9} < BR(B_s \rightarrow \mu^+ \mu^-) < 3.4 \times 10^{-8} \text{ 95\% C.L.}$$

MORIOND 13th March

0.37 → 1 fb<sup>-1</sup>

$$BR(B_s \rightarrow \mu\mu) < 4.5 \times 10^{-9} \text{ @ 95\% C.L.}$$

CERN seminar 6th March

1.14 → 4.9 fb<sup>-1</sup>

$$BR(B_s \rightarrow \mu\mu) < 7.7 \times 10^{-9} \text{ @ 95\% C.L.}$$

CERN seminar 13th March

ATLAS:

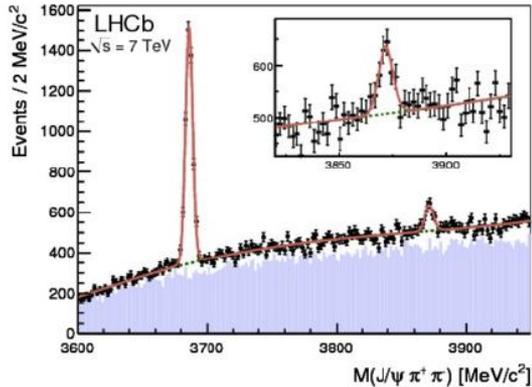
2.4 fb<sup>-1</sup>

$$BR(B_s \rightarrow \mu\mu) < 2.2 \times 10^{-8} \text{ @ 95\% C.L.}$$

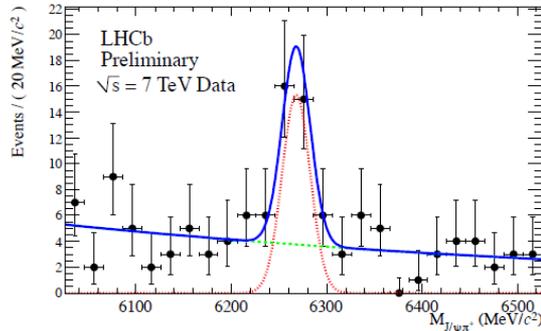
27

# Lots of other B physics

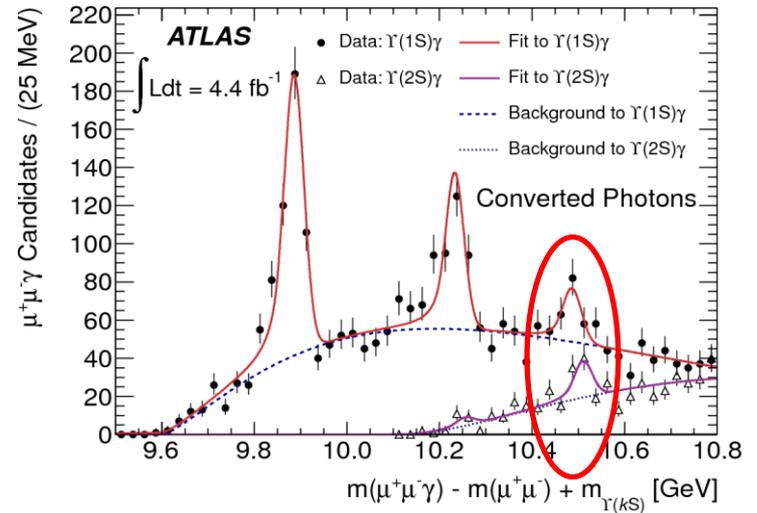
$M_{\chi(3872)} = 3871.95 \pm 0.48 \text{ (stat)} \pm 0.12 \text{ (syst)} \text{ MeV}$



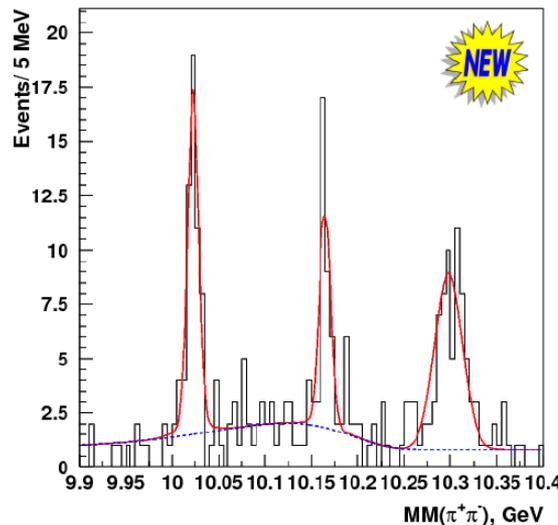
$M(B_c^+) = 6268.0 \pm 4.0 \text{ (stat)} \pm 0.6 \text{ (syst)} \text{ MeV}$



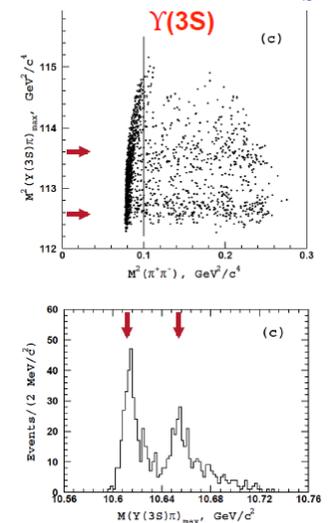
Mass ( $\chi_b(3P)$ ) =  $10.530 \pm 0.005 \text{ (stat.)} \pm 0.009 \text{ (syst)} \text{ GeV}$



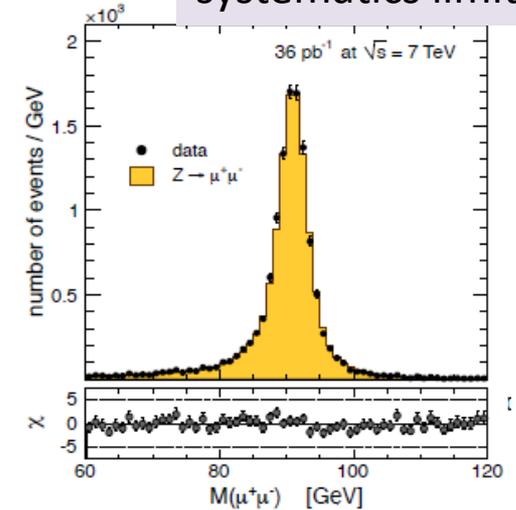
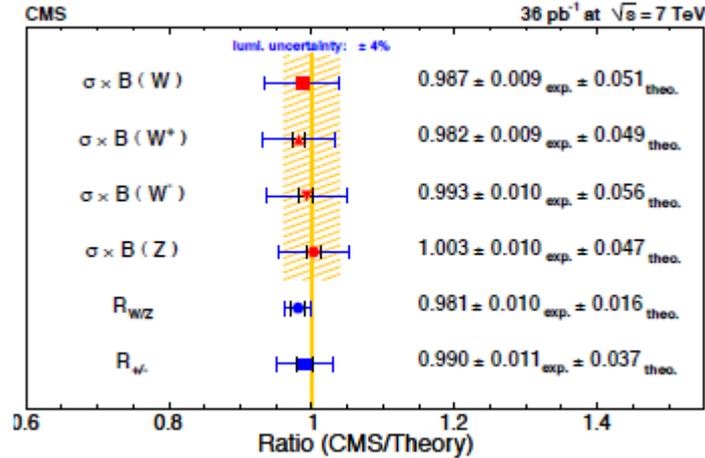
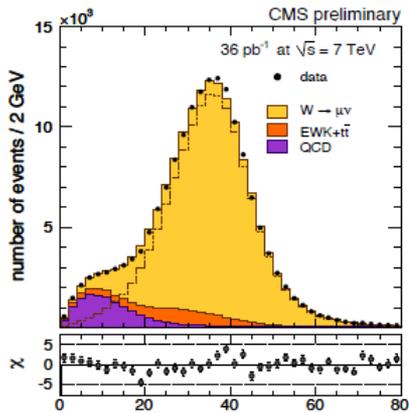
Observation of  $Y(5S) \rightarrow Y(1D)\pi^+\pi^-$



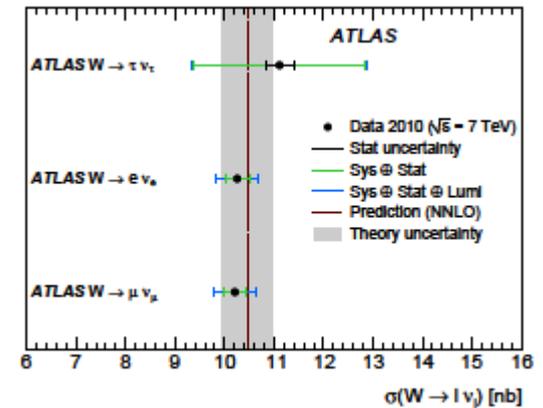
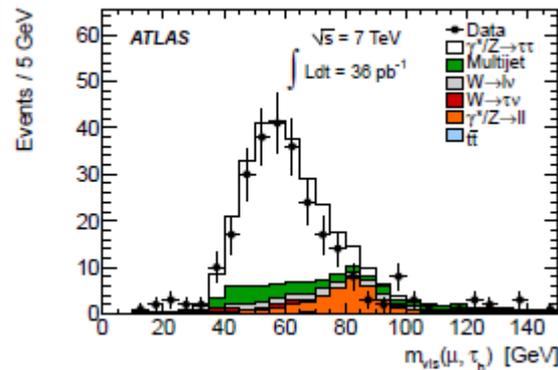
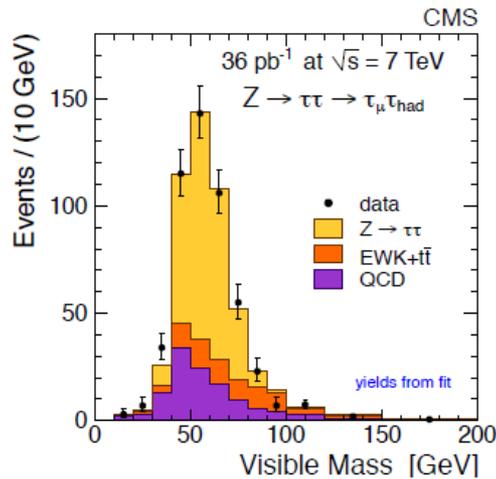
Discovery of  $Z_b$



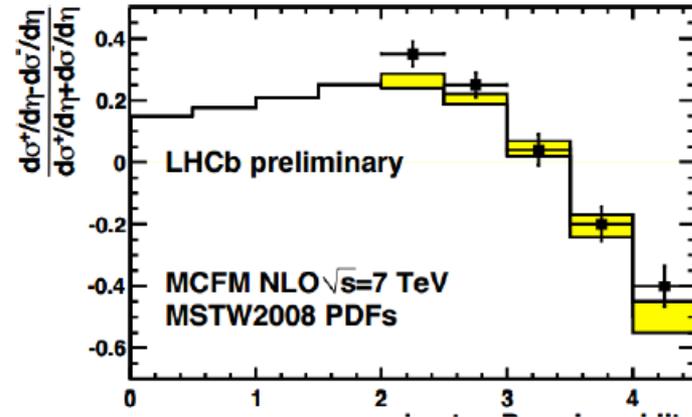
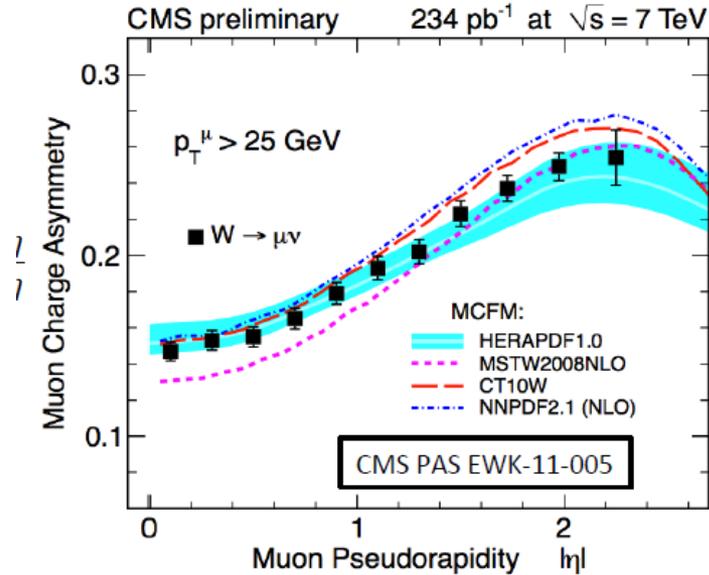
systematics limited



$\tau$  channels still statistics limited



# W charge asymmetry

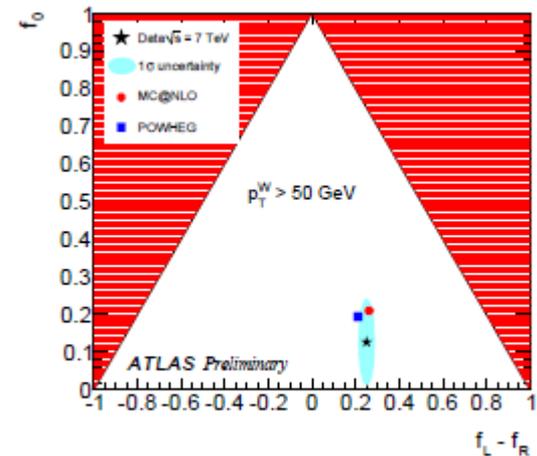


- Asymmetry changes sign in the forward region.
- In this region, PDF uncertainty on  $A_W$  is of the order of 2 - 8% (MSTW2008NLO, 68% C.L.)
- LHCb measurements already constrain PDFs with present analysis (analysis of the 2011 data ongoing).

Good agreement with HERAPDF  
 More flat than MSTW ( $\chi^2=5.3$ ), CT10 ( $\chi^2=2.1$ ), NNPDF ( $\chi^2=4.1$ )  
 Provides significant constraints to the PDF global fits

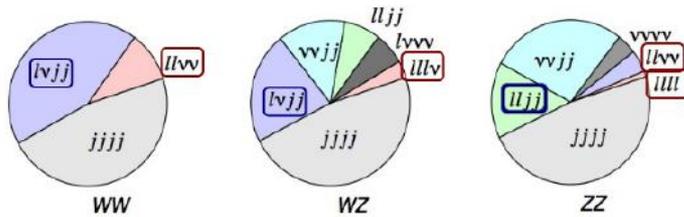
## W polarisation measurements

$$f_L + f_R + f_0 = 1.$$

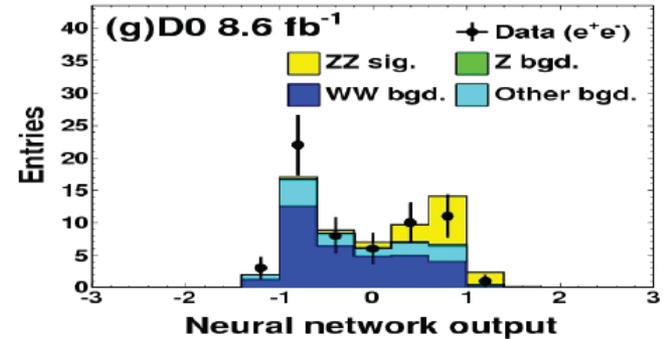


# Di-bosons @ Tevatron

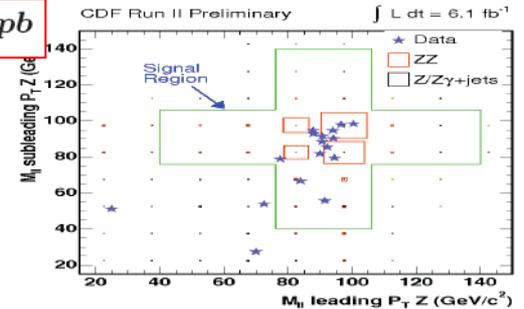
## Heavy Diboson Production



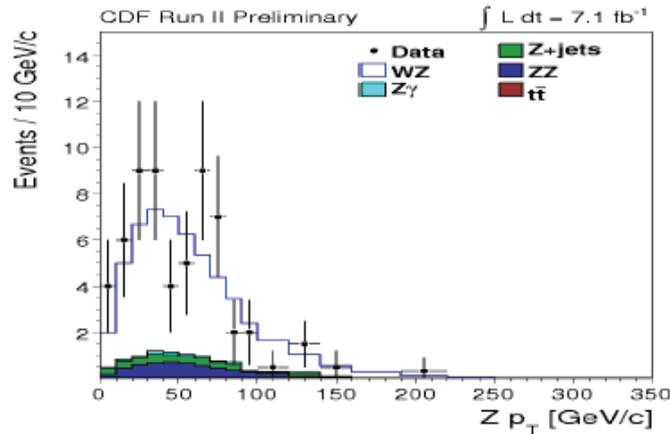
$$\sigma(ZZ) = 1.64 \pm 0.44(\text{stat.})_{-0.15}^{+0.13}(\text{syst.}) \text{ pb} \quad (SM = 1.30 \pm 0.10 \text{ pb})$$



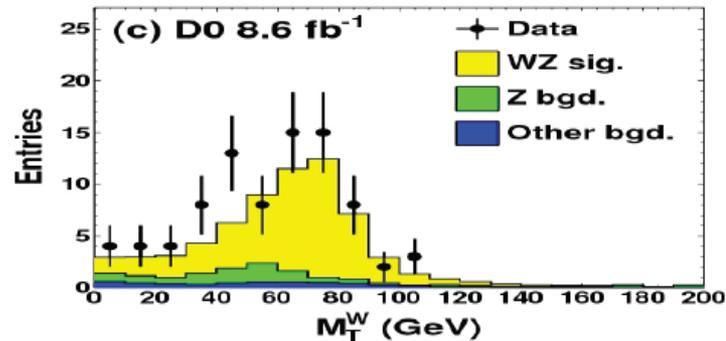
$$\sigma(ZZ) = 1.64_{-0.38}^{+0.44} \text{ pb}$$



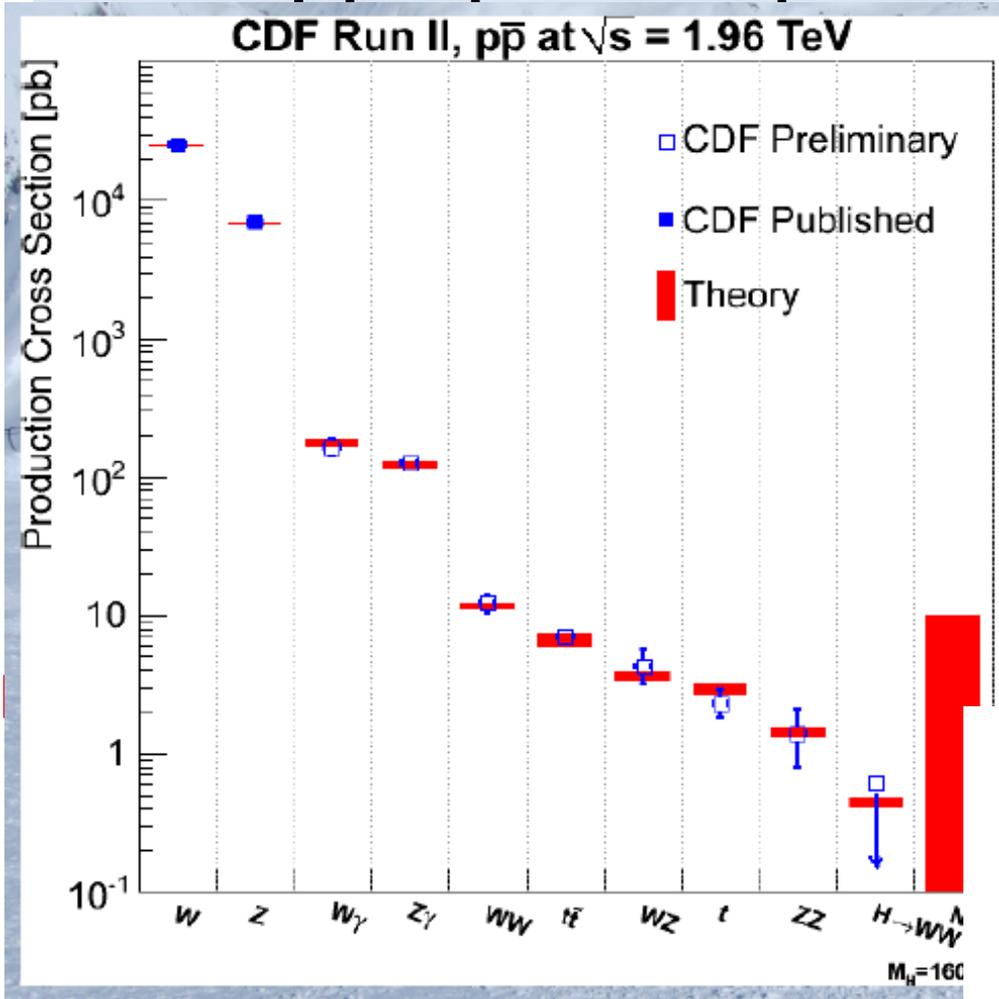
$$\sigma(WZ) = 3.9_{-0.5}^{+0.6}(\text{stat.})_{-0.4}^{+0.6}(\text{syst}) \text{ pb} \quad (SM = 3.46 \pm 0.21 \text{ pb})$$



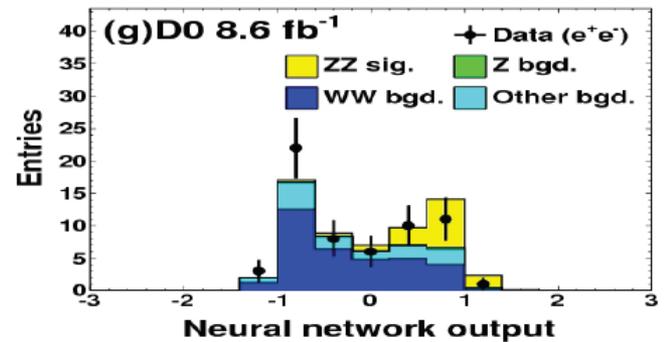
$$\sigma(WZ) = 4.50 \pm 0.61(\text{stat.})_{-0.25}^{+0.16}(\text{syst.}) \text{ pb}$$



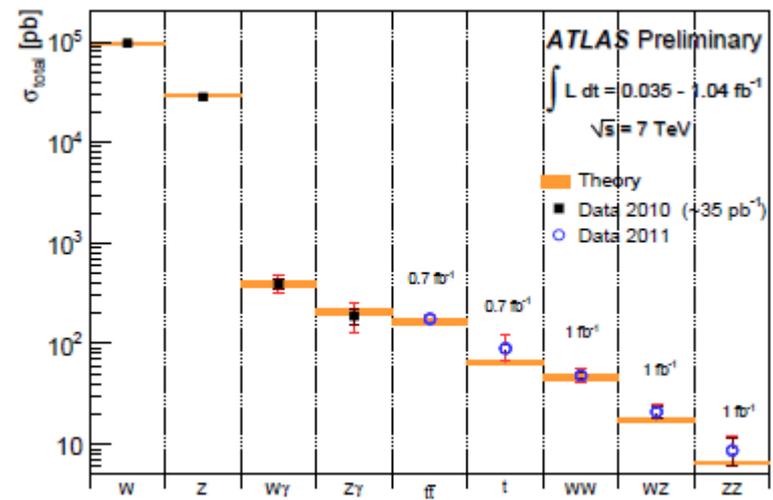
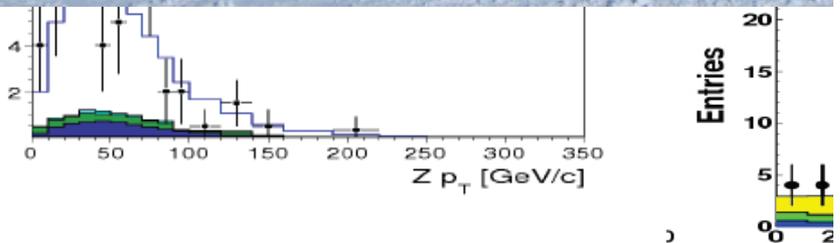
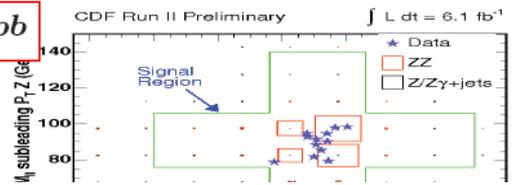
# Di-bosons @ Tevatron



$$\sigma(ZZ) = 1.64 \pm 0.44(\text{stat.})_{-0.15}^{+0.13}(\text{sys.}) \text{ pb} \quad (SM = 1.30 \pm 0.10 \text{ pb})$$

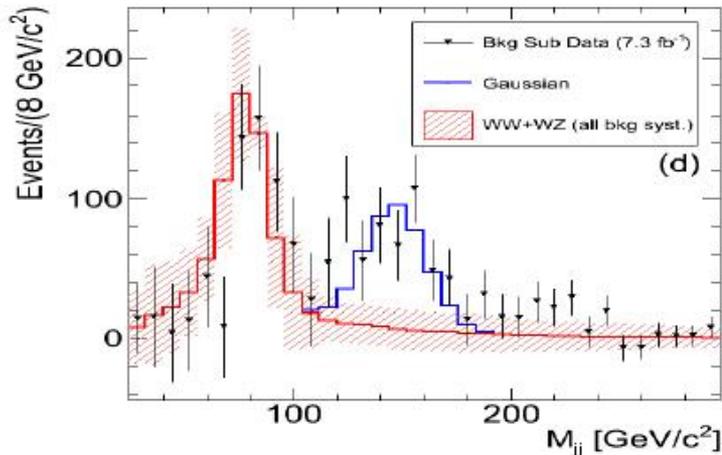
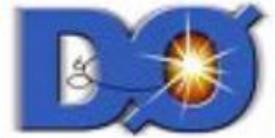


$$= 1.64_{-0.38}^{+0.44} \text{ pb}$$





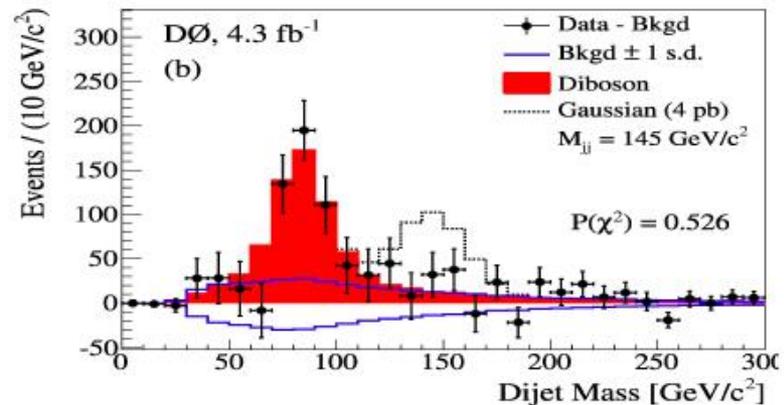
# Wjj anomalous production



- **4.1 $\sigma$  excess seen in dijet mass spectrum of W+2jet sample**

- Binned  $\chi^2$  fit to  $M_{jj}$  distribution consistent with  $\sigma = 3.0pb \pm 0.7$
- Many cross checks performed: various bkg control regions, W+jets modelling etc
- PHYS. REV. LETT. 106, 171801 (2011), and [Public Webpage](#)

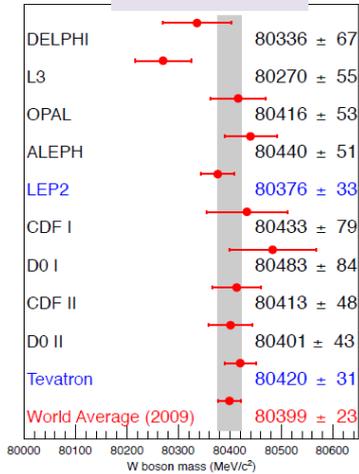
- DØ repeated CDF analysis → with some minor differences
- **No significant discrepancy w.r.t. background model**
- Results are  $2.5\sigma$  apart
- Phys.Rev.Lett. 107 (2011) 011804



- **S/B 10x worse at LHC**
- **Hard to understand W + jets at that level**

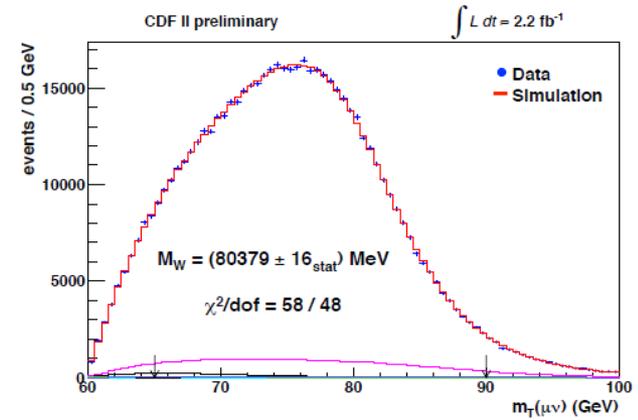
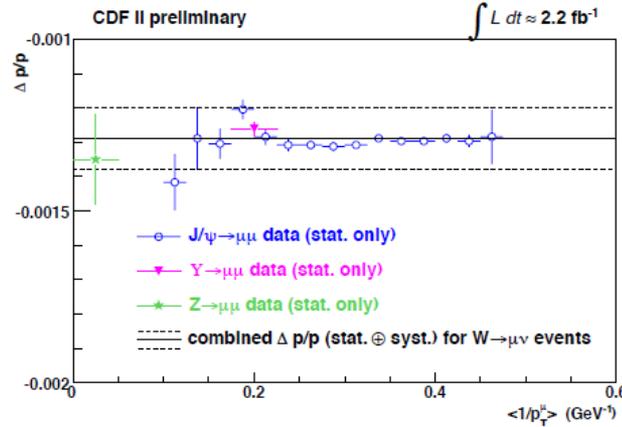
# New W mass from CDF

Previously

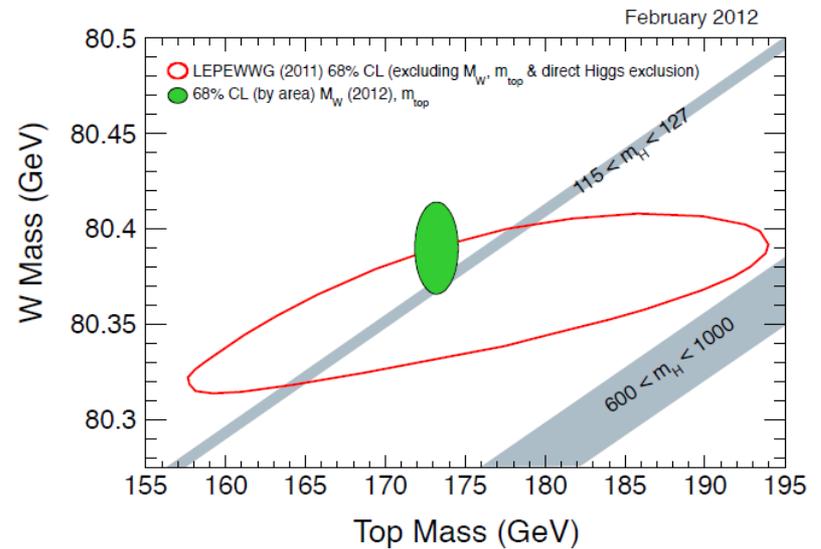
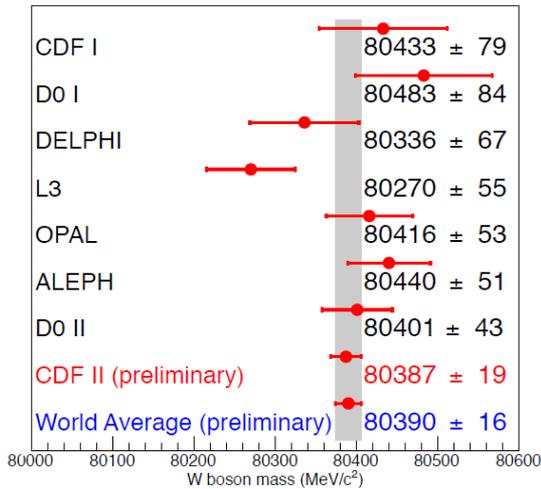


SM Fit  $M_H = 92^{+34}_{-26}$  GeV

$M_H < 161$  GeV @95% CL



$$M_W = 80387 \pm 12_{\text{stat}} \pm 15_{\text{syst}} \text{ MeV}/c^2$$



$M_H = 90^{+29}_{-23}$  GeV

$M_H < 145$  GeV @95% CL

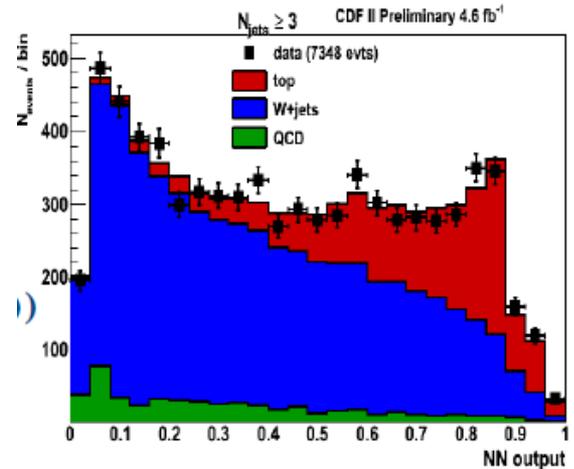
# Top physics

◆ **D0 l+jets 5.3fb<sup>-1</sup>:**  $\sigma_{t\bar{t}} = 7.78^{+0.77}_{-0.64}$  (stat+sys) pb

◆ **CDF l+jets 4.6 fb<sup>-1</sup>:**  $\sigma_{t\bar{t}} = 7.82 \pm 0.55$  (stat+sys) pb

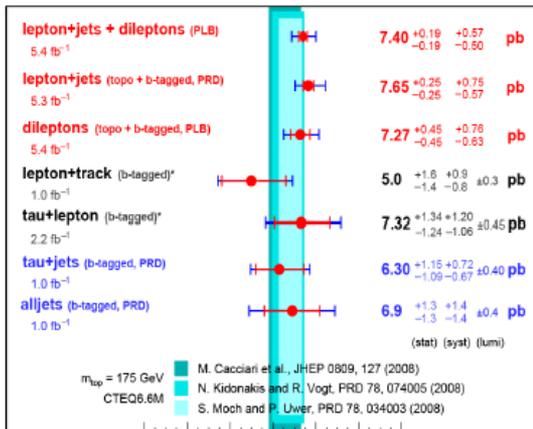
with a theoretical NNLO<sub>approx</sub> cross section of

• • • **7.46 pb @ m<sub>top</sub> = 172.5 GeV** (PRD 78, 034003 (2008))



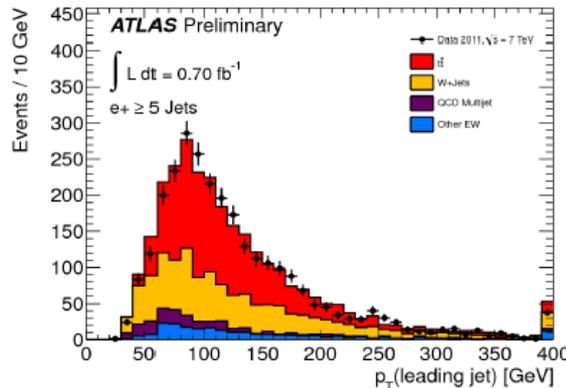
## DØ Run II

July 2011

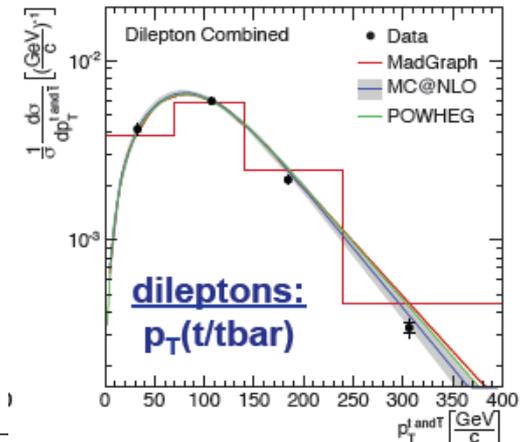


\* = preliminary  
 red = 2011 result  
 blue = 2010 results

$\sigma(p\bar{p} \rightarrow t\bar{t} + X)$  [pb]



$$\sigma_{t\bar{t}} = 179.0 \pm 9.8(\text{stat.} + \text{syst.}) \pm 6.6(\text{lumi.}) \text{ pb}$$



CMS combined 2011:

$$\sigma_{t\bar{t}} = 165.8 \pm 2.2(\text{stat.}) \pm 10.6(\text{syst.}) \pm 7.8(\text{lumi.}) \text{ pb}$$

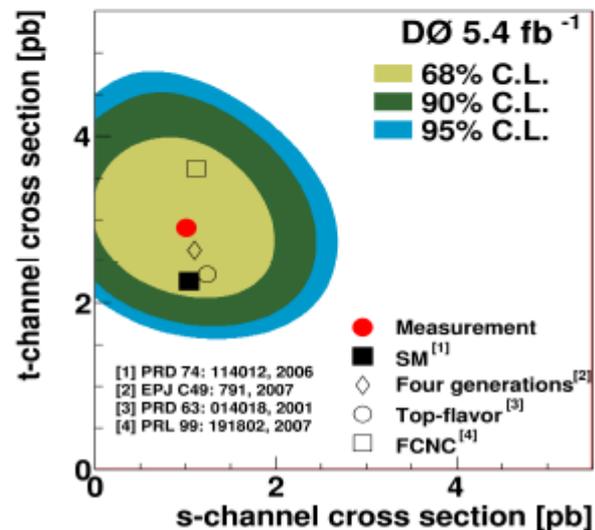
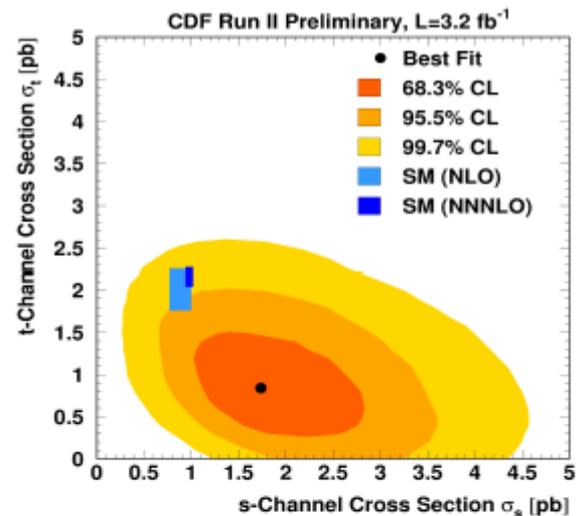
CMS Preliminary

# Single top @ tevatron

## Separate t- and s- Channel Production

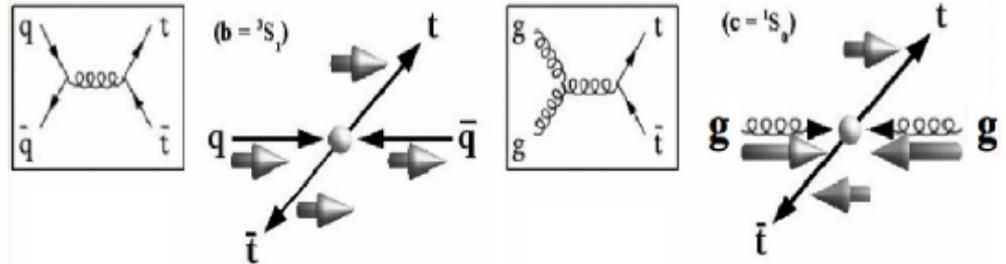
- ◆ **2-dimensional** measurement of s- and t-channel
  - t-channel sensitive to **anomalous couplings**
  - s-channel sensitive to **resonances**
- ◆ **strategy**: train separately for s- and t-channel
- ◆ **CDF 3.2 fb<sup>-1</sup>**:
  - $\sigma(t) = 0.8 \pm 0.4$  pb (PRD 82, 112005 (2009))
  - $\sigma(s) = 1.8^{+0.7}_{-0.5}$  pb
- ◆ **D0 5.4 fb<sup>-1</sup>**:
  - $\sigma(t) = 2.90 \pm 0.59$  pb (PLB 705, 313 (2011))
  - $\sigma(s) = 0.98 \pm 0.63$  pb
- ◆ **t-channel observation with 5.5  $\sigma$  at D0**
  - main systematics from background

LHC also starting to show results



- ◆ even **tops** are **not** produced in **polarized** state, the **spins** are **correlated**
- ◆ the correlation strength **A**

$$A = \frac{N_{\uparrow\uparrow} + N_{\downarrow\downarrow} - N_{\uparrow\downarrow} - N_{\downarrow\uparrow}}{N_{\uparrow\uparrow} + N_{\downarrow\downarrow} + N_{\uparrow\downarrow} + N_{\downarrow\uparrow}}$$



- depends on the **production mode** → different for Tevatron and LHC
  - choice of **spin basis** (here beam basis)
- ◆ due to the short top lifetime the **spin** does **not flip** and is **reflected** in the **angular** distributions of the **decay products**

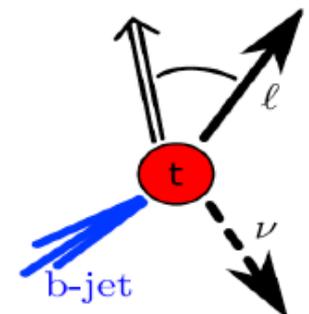
$$\frac{1}{\sigma} \frac{d\sigma}{d\cos\theta_i} = \frac{1}{2} (1 + \alpha_i \cos\theta_i)$$

with  $\alpha = 1$  for charged leptons and down-type quarks

- ◆ thus spin correlation can be measured by studying e.g.

$$\frac{1}{\sigma} \frac{d^2\sigma}{d\cos\theta_1 d\cos\theta_2} = \frac{1}{4} (1 - C \cos\theta_1 \cos\theta_2)$$

where  $C = A\alpha_1\alpha_2$



- ◆ even **tops** are **not** produced in **polarized** state, the **spins** are **correlated**
- ◆ the correlation strength  $A$



- ◆ D0 5.4 fb<sup>-1</sup> dilepton :

$$C_{\text{beam}} = 0.57 \pm 0.31 \text{ (stat+syst)}$$

- ◆ D0 5.3 fb<sup>-1</sup> l+jets :

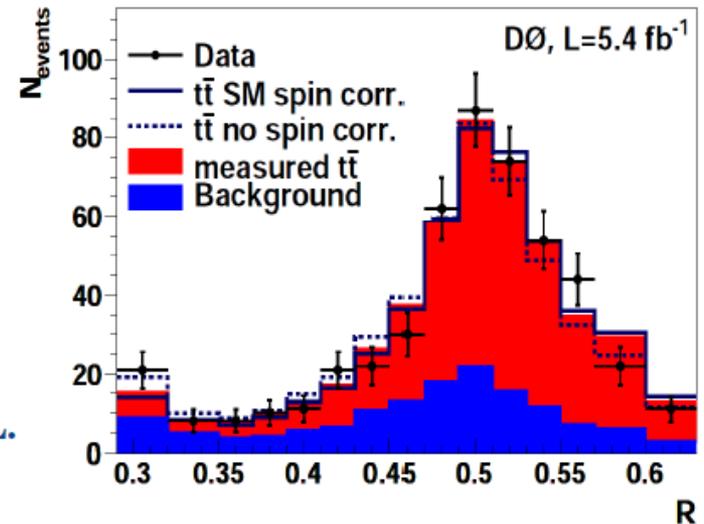
$$C_{\text{beam}} = 0.89 \pm 0.33 \text{ (stat+syst)}$$

- ◆ 30% increased sensitivity
- ◆ excellent agreement with SM
- ◆ combining statistically independent results:

$$C_{\text{beam}} = 0.66 \pm 0.23 \text{ (stat+syst)}$$

- $C < 0.26$  @ 95% C.L. and  $C < 0.04$  @ 99.7% C.L.
- $C = 0$  @ 3.1  $\sigma$  SD (PRL 108, 032004 (2012))

=> first **evidence** for **non-vanishing** spin correlation !

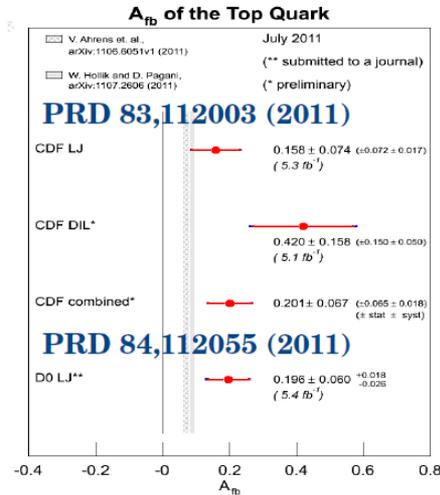


- ◆ thus spin correlation can be measured by studying e.g.  $b$ -jet

$$\frac{1}{\sigma} \frac{d^2 \sigma}{d \cos \theta_1 d \cos \theta_2} = \frac{1}{4} (1 - C \cos \theta_1 \cos \theta_2)$$

where  $C = A \alpha_1 \alpha_2$

# Top Asymmetry $A_{fb}^t$



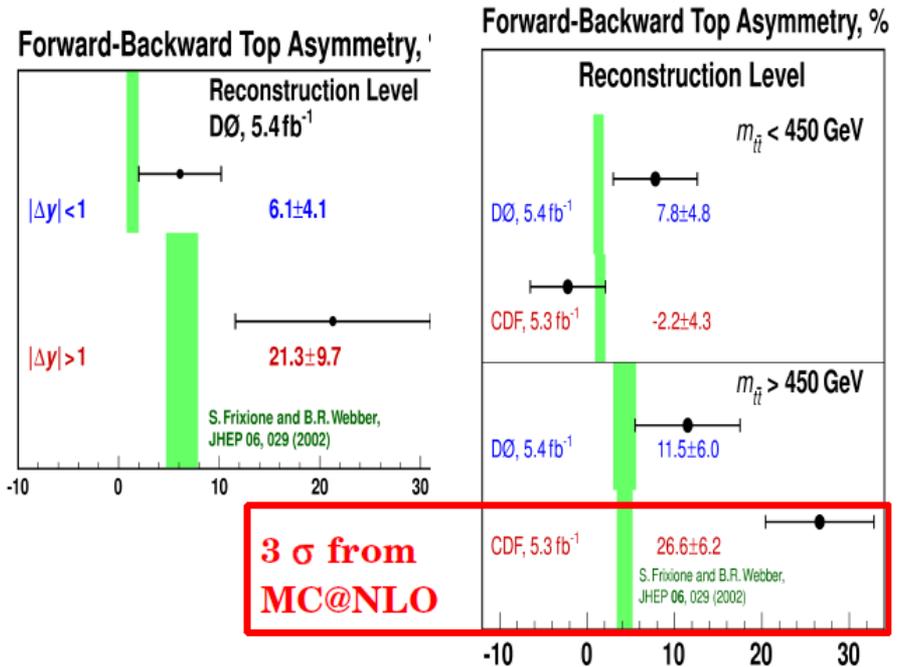
◆ measurements higher than prediction

◆ even larger difference for  $A_{fb}^t$ :

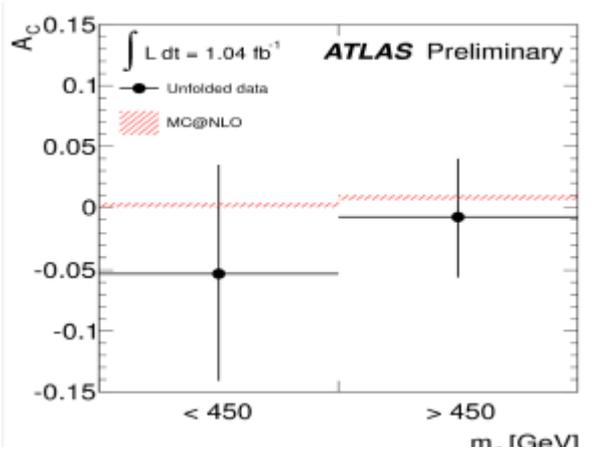
- D0 l+jets:  $A_{fb}^t = 14.2 \pm 3.8 \%$  where MC@NLO :  $A_{fb}^t = 0.8 \pm 0.6 \%$
- CDF dilepton:  $A_{fb}^t = 21 \pm 7 \%$

asymmetry depends on several variables like  $m_{\bar{t}t}$ ,  $|\Delta y|$

- e.g. new physics could lead to a different mass dependency



# ... And at LHC

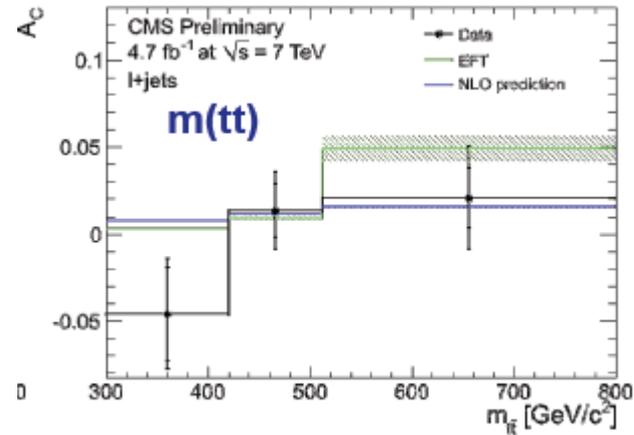


Result:

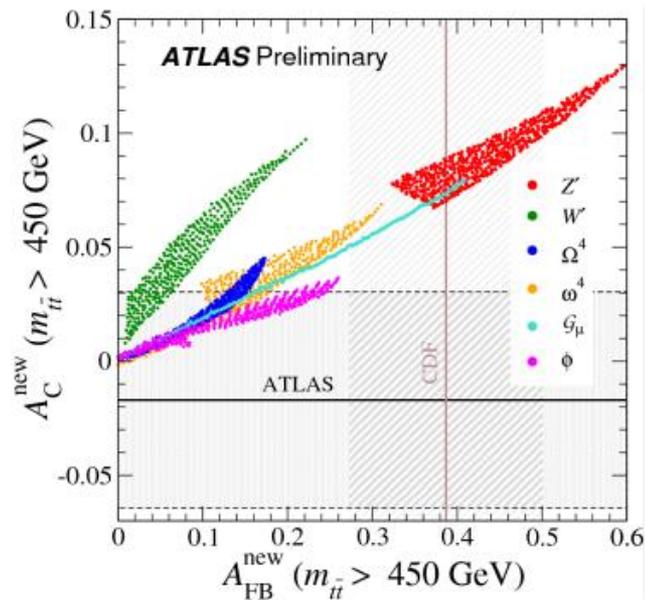
$$A_C = -0.018 \pm 0.028 \text{ (stat.)} \pm 0.023 \text{ (syst.)}^a$$

- consistent with the SM value of  $A_C^{pp} = 0.006 \pm 0.002$

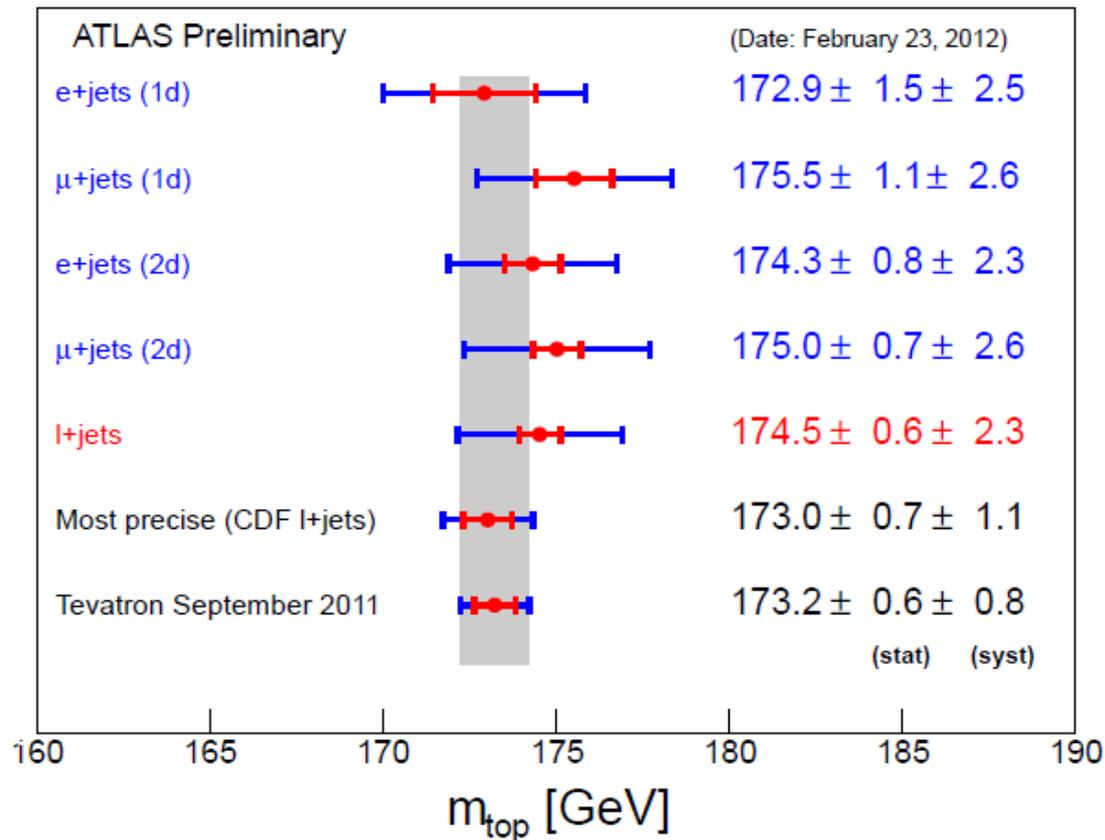
- limits on allowed  $A_C(\text{LHC})$  and  $A_{FB}(\text{Tevatron})$  regions can constrain some models proposed to explain the Tevatron  $A_{FB}$



Final corrected	$0.004 \pm 0.010 \text{ (stat.)} \pm 0.012 \text{ (syst.)}$
Theory prediction (SM)	$0.0115 \pm 0.0006$



# Top Mass



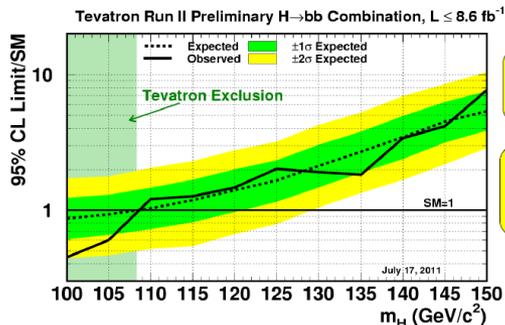
- reduction of the systematics uncertainty needed to reach Tevatron measurement uncertainty.
- Challenging, but there is a lot of space for improvement wrt. currently published measurement, incl. reducing JES, MC modelling uncertainties, reducing MC stats. effects by generating huge samples ...

# HIGGS - Tevatron



## Limits for $H \rightarrow b\bar{b}$

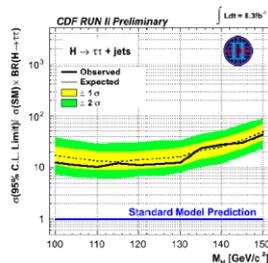
Summer 2011



New results coming very soon



## Searches with Taus



Limits at  $M_H = 115$  GeV:  
Exp:  $13 \times \sigma_{SM}$   
Obs:  $12 \times \sigma_{SM}$

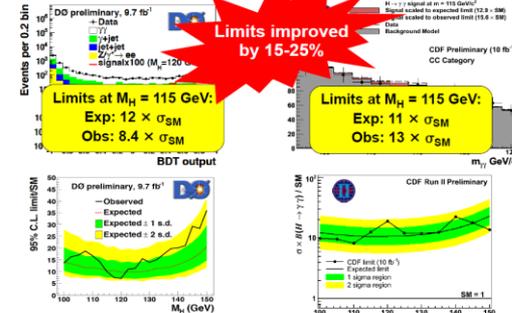
15% improvement  
in limits

Limits at  $M_H = 115$  GeV:  
Exp:  $14 \times \sigma_{SM}$   
Obs:  $11 \times \sigma_{SM}$

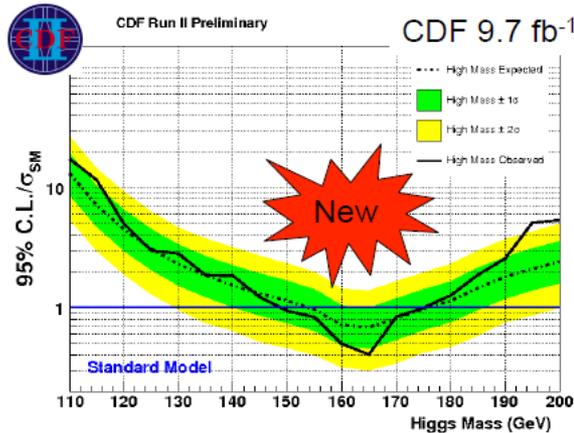


## Diphoton final states

Simple event selection: two photons



Figures Searches at the Tevatron



High mass search updated: 12%  
better at 165 GeV

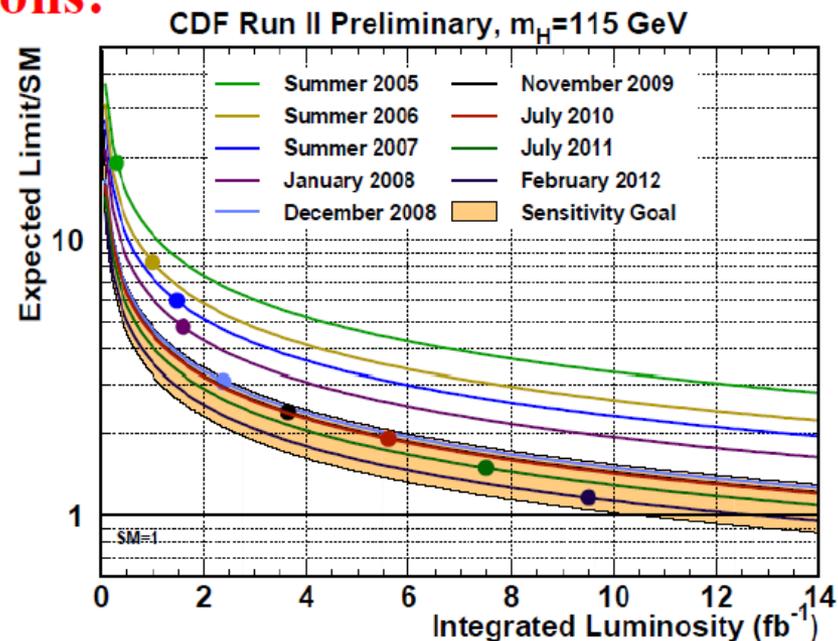
CDF Sensitivity close to TeV  
sensitivity in July 2011.

Analyses using  
dataset  
Stay Tuned!!

CDF gave us a taste of what is to come??

# Conclusions

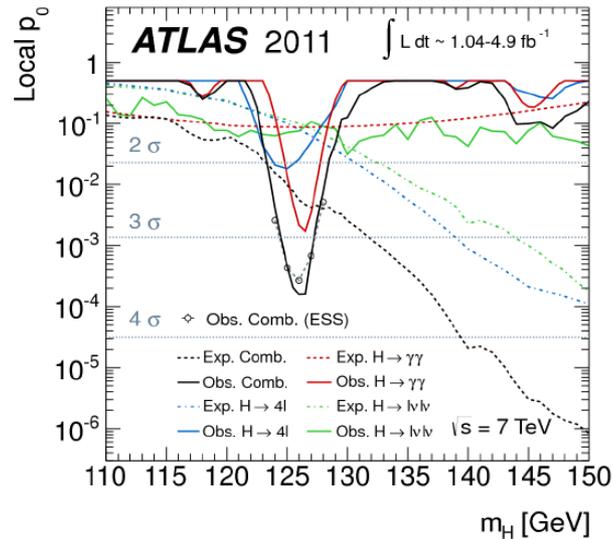
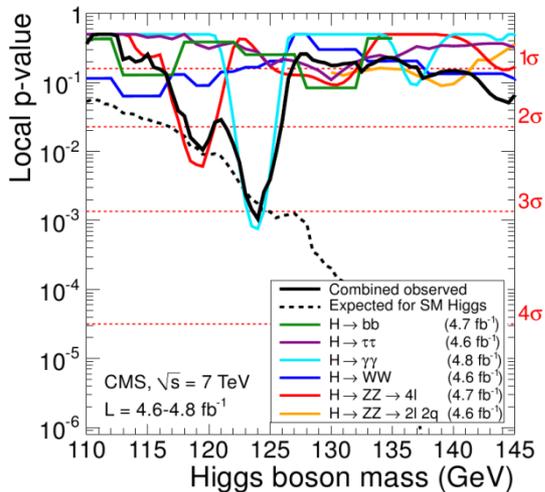
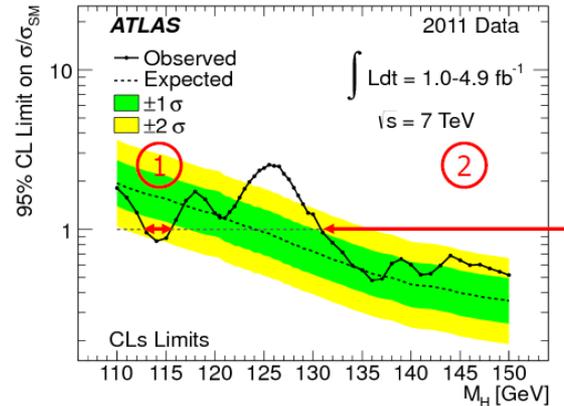
- The CDF Collaboration has produced Higgs searches with expected sensitivities **a factor of 2 better than 2007 beyond luminosity additions!**
- CDF is sensitive to  $< \sqrt{2} \times \text{SM}$  Everywhere
  - **2XCDF** would have **>25% chance of 3-sigma!**
- Tevatron Leads in  $H \rightarrow b\bar{b}$
- The Tevatron full dataset combined Higgs search will be exciting!



- Release observed limits next Weds at Moriond!
- **Tevatron Combination** talk at Moriond EWK by Wade Fisher!

# HIGGS - LHC

- **Higgs search at LHC/ATLAS is very exciting now!**
  - ATLAS still have several channels (WW/ $\tau\tau$ /bb etc) to be updated.
  - > Wait for new results until Morinod conf.
- In the low mass region, we have a small un-excluded mass window of **115.5-131GeV**.
- Small excess around 126GeV is real? -> **No conclusion at the moment**
  - > Still  $\sim 2$ -3 sigma level with the look elsewhere effect.
  - > We need more data in this year.

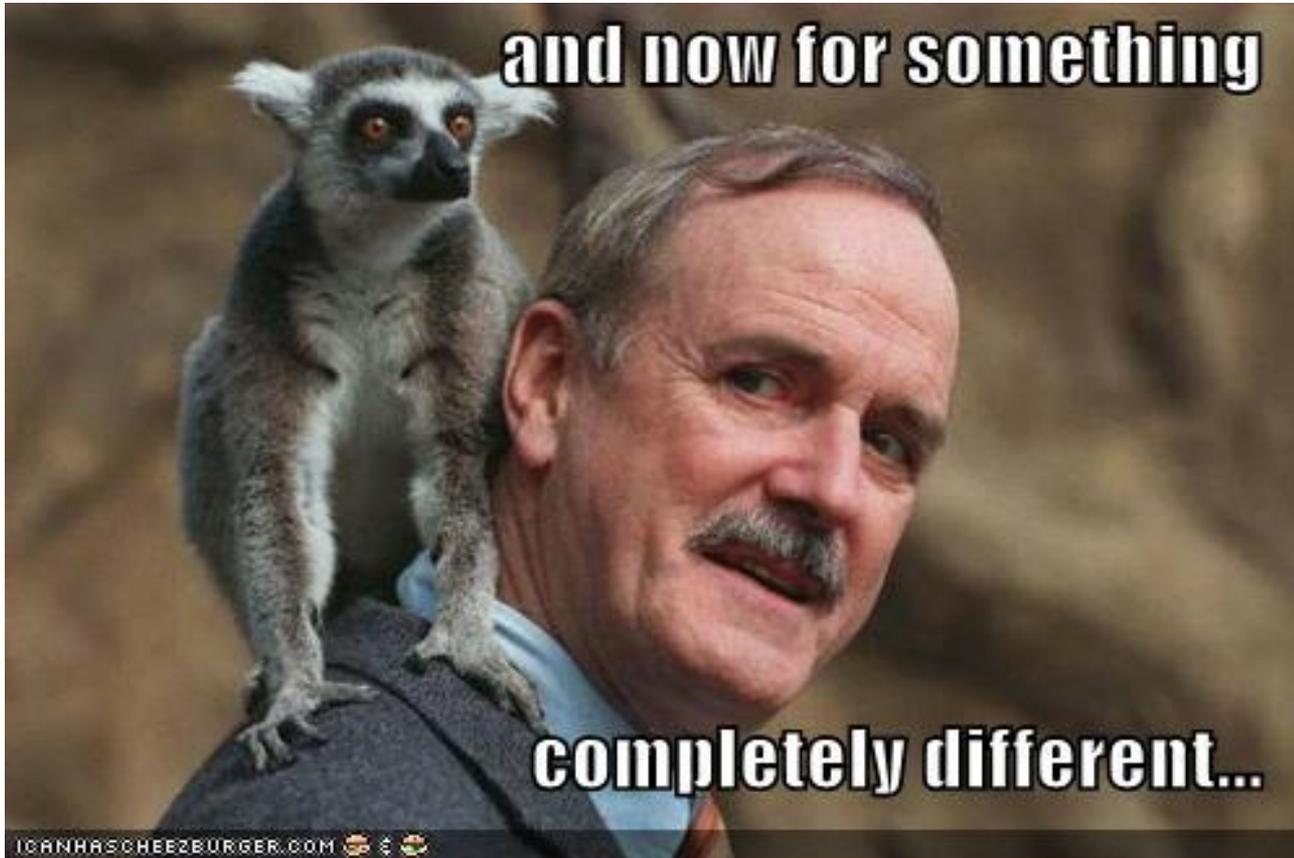


"More data are required to ascertain the origin of the observed excess."

# Beyond Standard Model

Lots of searches  
No signals



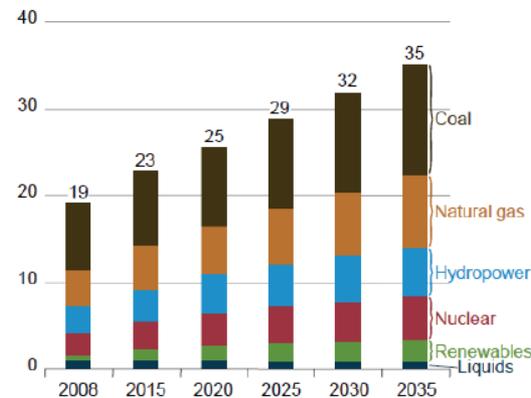


# Thorium Energy and Accelerator-Driven Reactors

## The Challenge

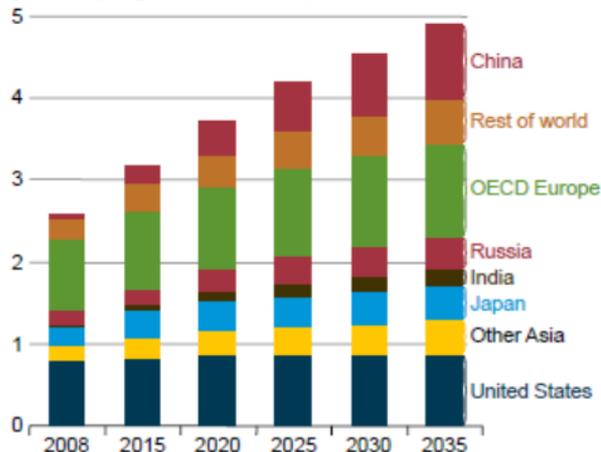


Figure 17. World net electricity generation by fuel type, 2008-2035 (trillion kilowatt-hours)



- World net electricity consumption increases by 84%
- World coal consumption increases by 50%
- World Nuclear power generation increases by ~90%, but remains flat at ~14% of total electricity generation

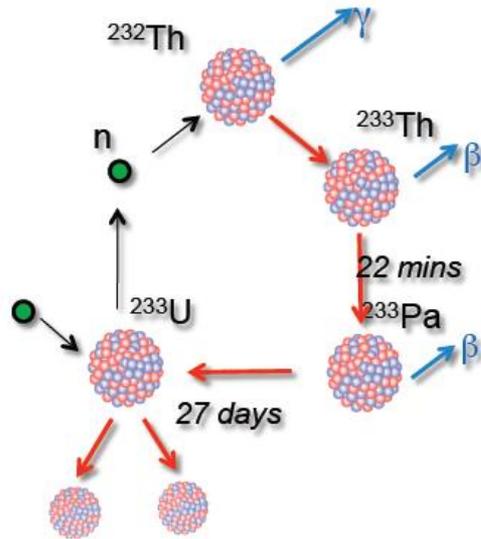
Figure 76. World net electricity generation from nuclear power by region, 2008-2035 (trillion kilowatt-hours)



## Problems for nuclear power:

- Public acceptance
- Real accidents
- Radioactive waste
- Proliferation concerns
- Not competitive in terms of cost

# Thorium/Uranium Fuel Cycle



- U-235 is the only naturally occurring fissile material
- U-233 is also fissile but does not exist in nature
- U-233 can be produced from Th-232 via neutron capture
- Reactors that generate fissile U-233 from Th-232 are Breeder Reactors

“Previous work on Thorium reactors dropped because they didn’t produce enough Plutonium!”

~4 times more abundant than U-238,  
400 times more abundant than U-235  
As abundant as lead

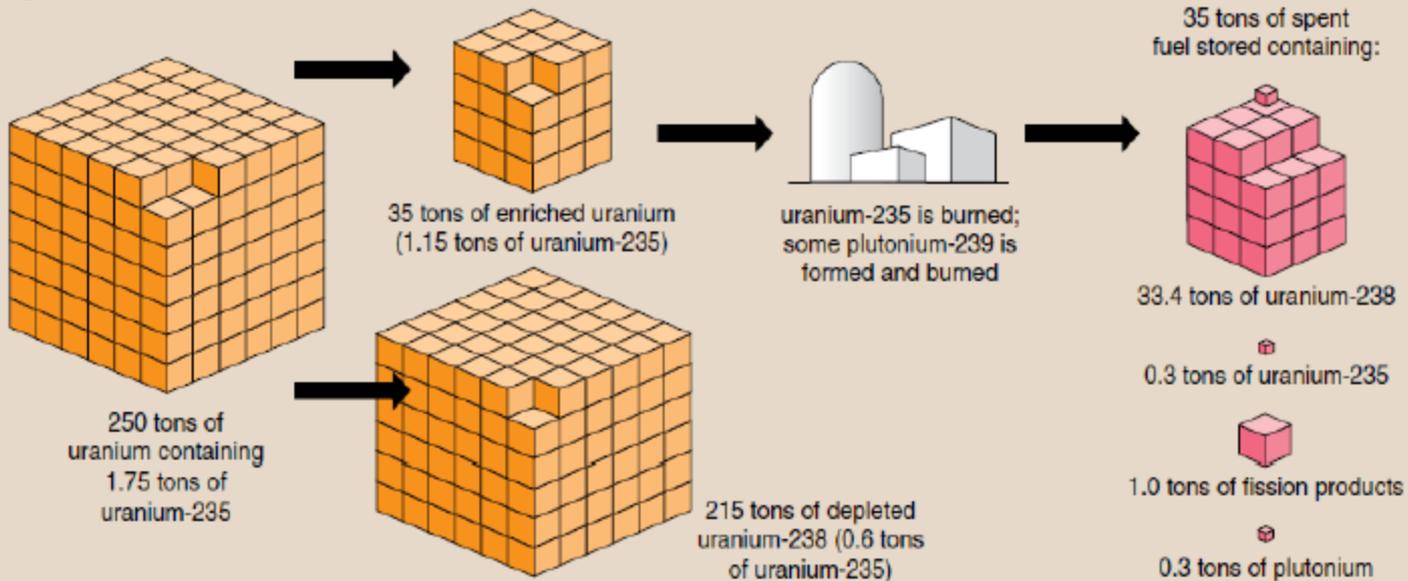
Thorium currently costs only US\$30/kg, while the price of Uranium has risen above \$100/kg, not including costs for enrichment and fuel fabrication.

89 Ac Actinium (227)	90 Th Thorium (232, 230)	91 Pa Protactinium (231, 230, 229)	92 U Uranium (238, 235, 234, 233, 232, 231)	93 Np Neptunium (237)	94 Pu Plutonium (244)	95 Am Americium (243)	96 Cm Curium (247)	97 Bk Berkelium (247)	98 Cf Californium (251)	99 Es Einsteinium (252)	100 Fm Fermium (257)	101 Md Mendelevium (268)	102 No Nobelium (259)	103 Lr Lawrencium (262)
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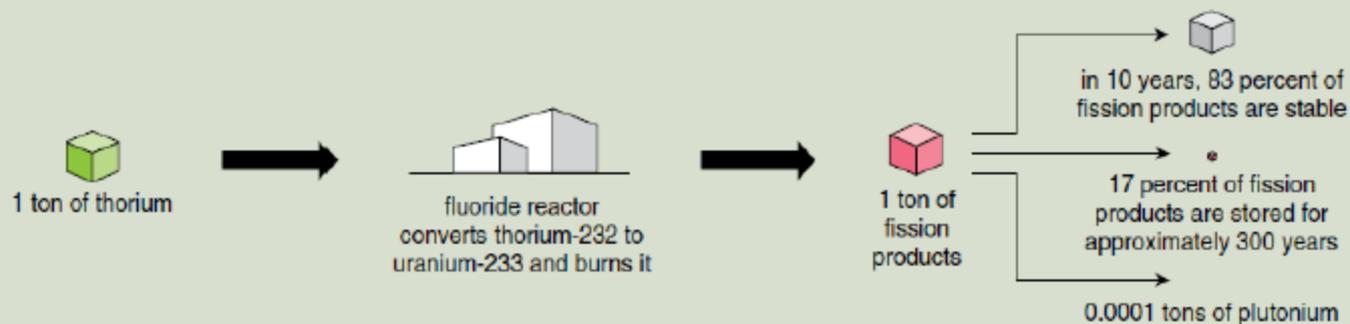
Minor actinides -> long lived radioactive waste

# Advantages of Thorium: Raw Material and Resource Utilization for 1GW-year

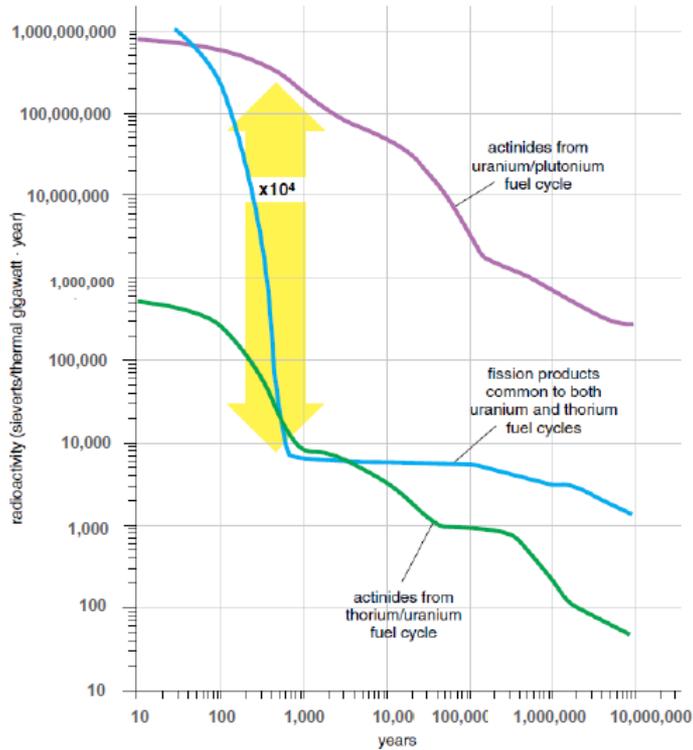
## light water reactor



## liquid fluoride thorium reactor



# Waste Storage

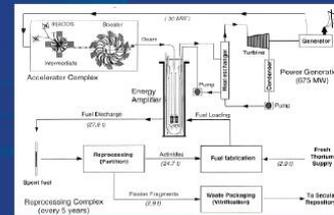


## Deploying Thorium Energy: Three Approaches



Thorium fuel in solid form in conventional reactors

Thorium as fuel in molten-salt reactors



Accelerator-driven Subcritical Reactors using Thorium fuel

C. Rubbia, Energy Amplifier

- **Europe:** Belgium has committed to build MYRRHA, the first ADS demonstration reactor
- **India:** National nuclear power strategy is based on Thorium; ADS is needed to breed U-233
- **China:** Announced development program to build 1 GW ADS by 2032



**EPLANET** European Particle physics Latin American NETWORK (EPLANET)

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*Credits:  
Black hole simulation realized  
by ATLAS-CERN Team*

Visitor No. \_\_\_\_\_

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- 4 years, budget 3.245 ME, 1803 months;
- started 1feb 2011(about 2 years gap from HELEN).

A wide-angle photograph of a snowy mountain landscape. In the foreground, a red and black pole stands in the snow. A ski lift line with several chairs is visible in the middle ground, winding up a slope. The background shows more snow-covered mountain peaks under a clear sky.

Thankyou  
for  
listening!

<http://agenda.infn.it/conferenceOtherViews.py?view=standard&confId=4116>